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# Numerical simulation of homojunction *p-i-n* In0.4Ga0.6N solar cell with different absorber layer configurations

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https://doi.org/10.1016/1 \*Corresponding author: <u>shashiong@usm.ng</u>(6)<sup>11</sup> The effects of thickness, carrie were evaluated. Numerical studies were performed to optimize the homojunction *p-i-n* In<sub>0.4</sub>Ga<sub>0.6</sub>N solar cell using SCAPS-1D. With the optimized thicknesses and carrier densities, the *p-i-n*  $In_{0.4}Ga_{0.6}N$  solar cell shows a maximum conversion efficiency of 18.74%. By changing the fixed indium composition of the absorber layer into step- and linear-graded configurations, the conversion efficiency, open-circuit voltage, and short-circuit current density of the solar cell were further improved. The results showed that the solar cells with step- and linear-graded absorbers attained maximum conversion efficiency of 19.77% and 19.84%, respectively. These results implied that the absorber with the graded compositional design helps to reduce the barrier height of electron where transport for band-to-band and interband absorption? This eventually reduces the codigy loss caused by phonon exprision and heat thermalization. 10101,1001,1001,1001,1001,1001,1000,1

# Highlights

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Optimization of intermediate indium composition *p-i-n* homojunction solar cell by using • SCAPS-1D

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- Implementation of novel step-graded and linear-graded absorber layer to improve the performance of the solar cell
- Achieved quantum efficiency (QE) close to 100% by using novel absorber configurations.

Achieved quantum efficiency (QE) close to 100% by using novel absorber configurations.
 Reduction of photon energy losses through thermolization and phonon emission with graded absorber system configurations. Awarned the potential of intermediate indium composition InGaN-based solar cells in the Awarned the potential of intermediate indium composition InGaN-based solar cells in the normalization of the potential of intermediate indium composition InGaN-based solar cells in the normalization of the potential of intermediate indium composition InGaN-based solar cells in the normalization of the potential of intermediate indium composition in GaN-based solar cells in the normalization of the potential of intermediate indium composition in GaN-based solar cells in the normalization of the potential of intermediate indium composition in GaN-based solar cells in the normalization of the potential of intermediate individual of the potential of intermediate individual of the potential of intermediate indicates in the normalization of the potential of intermediates indicates in the normalization of the potential of intermediates indicates in the normalization of the potential of intermediates indicates in the normalization of the potential of intermediates indicates in the normalization of the potential of intermediates indicates in the normalization of the potential of intermediates indicates in the normalization of the potential of t



# 1. Introduction

Various photovoltaic (PV) technologies have been developed to mitigate the reliability of fossil fuels in producing electricity. Such PV technologies are crystalline silicon, cadmium telluride (CdTe), copper indium gallium selenide (GICS), perovskite, organic, multijunction III-V, and dye-sensitized solar cell (DSSC) [1]. However, most of these technologies are susceptible to ionized electrons produced by radiation which will act as stable impurity complexes to alter the optical, electrical and structural properties of the solar array [2-4]. Therefore oraterial such as indium gallium nitride (InGaN) has gained attention due to its extraordinally properties. InGaN is a promising material for optoelectoric devices such as solar cells plaser diodes, lightemitting diodes, and high electron poblicity transistors. This is mainty attributed to its tunable direct bandgap energy characteristics from 0.70 eV (infrared) to 3.42 eV (ultraviolet) by ntips://doi.or varying the indium (In) mposition. Thus, InGaN is excellent candidate for solar cell application since it an cover almost the entire pectrum from solar irradiation [5]. The incorporation of indium nitride (InN) improved the radiation resistance of InGaN when subjected to proton irradiation, which dominates the radiation environment in space. Eventually, InGaN has a longer carrier lifetime when compared to gallium arsenide (GaAs) and indium gallium arsenide (InGaAs), which makes it a promising solar cell material for space exploration [6].

Typically, for the InGaN-based single-junction solar cell to achieve a conversion efficiency of more than 20%, InGaN layers with an indium composition of more than 40% and several hundred nanovaters are required [7–10]. However, the growth of thick and high quality In-rich InGaN (i.e., with In composition of more than 40%) is extremely challenging due to several constraints such as high equilibrium nitrogen (N) vapor pressure, low dissociation temperature of InN (around 650 °C), difficulty to obtain lattice-patched substrate, and the other modynamic stability [11], desides that, thick In-rich InGeN thin films often encountered phase separation and compositional fluctuation due to the miscibility gap between InN and GaN species [12]. The presence of compositional fluctuation and other defects within the thin film is detrimental to the open-circuit voltage ( $V_{oc}$ ) and short circuit current density ( $J_{sc}$ ) of the In-rich InGaN-based solar cell [13]. Hence, InGaN layers with intermediate In composition is favorable for the solar cell application.

Islam et al. (2013) have successfully grown thick InGaN thin film with intermediate In composition (~40%) asing metal-organic vapor phase epitaxy. X-ray diffraction profile  $(2\theta - \omega)$ showed no phose separation, indicating good compositional homogeneity of the InGaN thin study also showed that the the rmal expansion coefficient minimatch between the film. Their (2010) optimized the grawth conditions of InGaN by value. and trimethylindium/ (trimethylindium + triethylgallium) ratio. Although the studies showed the maximum In incorporation up to 100%, phase separations were observed for InGaN thin films with In composition above 60% due to thermodynamic instability. They managed to grow single-phase InGaN samples with In composition between 10% to 45% [15]. Sang et al. (2021) reported the growth of In<sub>0.4</sub>Ga<sub>0.6</sub>N with a thickness of more than 300 nm by using high pressure MOCVD with the growing pressure up to 2.5 atm, and the growth temperature varied from 680-830 °C. It was found that high pressure growing conditions could improve the crystalline quality and better surface morphology of the InGaN thin film. Phase separation in the In-rich InGaN thin was suppressed under high-pressure groups g conditions [16]. Hence, the prowth of intermediate In composition InGaN (~40%) can be assumed to be consistent based on the current advorces. Apart from that, the growth of In-rich p-type InGaN tion film is another challedge for developing In-rich InGen solar cells. Several studies found that the magnesium https://doi.org/Mg) dopant flow rate, growth temperature, and annealing temperatures play important roles in forming *p*-InGaN [11]. Heavy Mg doping concentration deteriorates the crystalline quality of the *p*-InGaN by phase separation [17]; while *n*-type electrical characteristic increases due to electron accumulation on the surface of the thin film in *p*-type InGaN with increasing In composition [18]. Therefore, the p-InGaN is limited to having In composition of less than 40% and a maximum hole concentration of around  $10^{19}$  cm<sup>-3</sup> [11,19–21]. Hence, the growth of InGaN solar cells with intermediate In composition (~40%) is feasible with the current growth

advances.

From LoFature reviews, several growth hallenges of In-rich InGaN to P pinpointed. Those childenges limit the growth stability of the In-rich InGaN with Information of more than 90% and hole concentration of more than 10<sup>19</sup> cm<sup>-3</sup>. Next, the cMN/InGaN heterostructure hole on only introduces more disocation into the thin films but also induces piezoelectric polarization, which will create drift currents opposite in direction to reduce the carrier collection efficiency [22]. In this work, we studied a three-layered *p-i-n* homojunction solar cell with intermediate indium composition of 40%, which used the current growth limitations to evaluate the performance of the InGaN solar cell. The homojunction *p-i-n* was proposed to minimize the effect of piezoelectric polarization and the lattice mismatch issue in the solar cell structure. The carrier density investigation was carried out to pinpoint the best carrier densities required to achieve maximum conversion efficiency. Several studies showed that the defect and dislocation densities are the main reason limiting the performance of InGaN *p-i-n* solar cells [23,24]. Therefore, the effect of dislocation densities was taken into consideration.

In this paper, the optimization of standard APn In<sub>0.4</sub>Ga<sub>0.6</sub>N has been done to varying the thickness and carrier density. To further paper we have the solar cell performance, novel stepand linear graded absorber designs were proposed. The simulation data for standard, step- and linear-graded In<sub>0.4</sub>Ga<sub>0.6</sub>N were compared to identify the best graded absorber configuration.

#### 2. Theoretical modeling and simulation

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In this paper, three different solar cell structures were investigated, and their respective configurations are shown in Figure 1. The first one is named standard *p-i-n* layer structure, utilizing a fixed absorber layer of In<sub>0.4</sub>Ga<sub>0.6</sub>N material. The second and third are linear- and ulations were performe University of Gobt; Belgium. step-graded absorber structures using graded In<sub>x</sub>Ga<sub>1-x</sub>N absorber layers. The configured absorber layers were proposed using the optimized structure of the standard *p-i-n*. All the simulations were performed using the SCAPS-1D splar cell simulator developed by the University of Gopt; Belgium.  $10^{10}$   $M^{1.5}$   $10^{10}$   $10^{10}$   $M^{1.5}$   $10^{10}$   $p-In_0 Ga_0 N$ p-In<sub>0.3</sub>Ga<sub>0.7</sub>N *i*-In<sub>0.4</sub>Ga<sub>0.6</sub> p-In<sub>0.4</sub>Ga<sub>0.6</sub>N n-In<sub>0.4</sub>Ga<sub>0.6</sub>N n-In0.4Ga0.6N i-In<sub>0.4</sub>Ga<sub>0.6</sub>N (b) (a) n-In<sub>0.4</sub>Ga<sub>0.6</sub>N (c)

Figure 1. Schematic diagrams of InGaN solar cells with different configurations: (a) standard,(b) linear-graded absorber, and (c) step-graded absorber *p-i-n* solar cell structures.



where x is the In composition of the  $In_xGa_{1-x}N$ ;  $E_g^{InN}$  is the bandgap energy of indium nitride (InN) and is equal to 0.70 eV;  $E_g^{GaN}$  is the bandgap energy of gallium nitride (GaN) and is equal to 3.42 eV; b is the bowing factor equal to 1.43 eV. The electron and hole mobilities can be calculated as a function of carrier density with the Caughey-Thomas approximation [10], as shown in Eq. (2)

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Table 1. Parameters used to calculate the mobilities of the InGaN

	$\frac{\mu_{min,e}}{(cm^2/Vs)}$	$\frac{\mu_{max,e}}{(cm^2/Vs)}$	γe	$\mu_{min,h}$ (cm <sup>2</sup> /Vs)	$\frac{\mu_{max,h}}{(cm^2/Vs)}$	$\gamma_{ m h}$	Ref.
InN	429	3970	1	20	220	2	[25,26]
GaN	55	1000	1	3	170	2	[27]

The dielectric permittivity ( $\varepsilon$ ) and the electron affinity ( $\chi$ ) of In<sub>x</sub>Ga<sub>1-x</sub>N can be expressed using



The effective density state of conduction  $(N_c)$  and the valence band  $(N_v)$  of the In<sub>x</sub>Ga<sub>1-x</sub>N can be expressed using Eqs. (5) and (6), respectively;

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$$N_c = (0.9x + 2.3(1 - x)) \times 10^{18}$$
<sup>(5)</sup>

$$N_{\nu} = (5.3x + 1.8(1 - x)) \times 10^{19} \tag{6}$$



 $\alpha_o$  is the reference absorption coefficient. The absorption coefficient of InGaN is assumed to be the same as GaN, which is approximately  $2 \times 10^5$  cm<sup>-1</sup> [29].

This study performed the simulations at 300 K with an incident solar radiation of 1000 W/m<sup>2</sup> (AM 1.5G) as the environmental boundary conditions. Table 2 shows the initial boundary conditions and the parameters used to evaluate the performance of the standard *p-i*n InGaN solar cell. The optimizations were first performed on the standard *p-i-n* structure by varying the thickness of p-, i-, and n-layers consecutively to obtain the optimized thickness of each layer. Next, the surfier density of each layer was waried from  $10^{13}$  cm<sup>-3</sup> to  $10^{21}$  cm<sup>-3</sup> to study the effect of carrier density on the solar cell's overall efficiency. Subsequently, the solar cell performance of standard *p-i-n* was compared with the configured absorber structures (e.g., linear staded *p-i-n* and step graded *p* (*n*) to investigate the roles of thinkel absorber structure of the performance of the InGen solar cell. Several experimental results indicated that the https:///di.org/ defect density of the InGaN thin film is within  $10^{10} - 10^{12}$  cm<sup>-3</sup> [30,31], to consider the possible occurrence of the defect during the fabrication processes, the defect density,  $N_t$  has been increased by two orders of  $10^{14}$  cm<sup>-3</sup> for this study.

Table 2. Initial boundary conditions and the parameters used for evaluating the *p-i-n* InGaN-<br/>based solar cell.

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	Structural parameters	Values
-	<i>p</i> -layer initial thickness	25 nm
	<i>i</i> -layer initial thickness?	100 nm
	<i>n</i> -layer initial thickness	100 nm
	Acceptor doping density in <i>p</i> -layer, $N_A$	$1 \times 10^{18}  \mathrm{cm}^{-3}$
	Carrier density of intrinsic layer, $N_D = N_A$	$1 \times 10^{16} \text{ cm}^{-3}$
	Donor define density in <i>n</i> -layer, $N_D$	$1 \times 10^{18}$ cm
	Pre-existing defect density, $N_t$	$1 \times 10^{14}$ cm <sup>-3</sup>
	Rachative recombination coefficient, B	$9.6 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1} [32]$
	Auger electron and hole capture coefficient, $C_p$ and $C_n$	$4.7 \times 10^{-31} \text{ cm}^{6}/\text{s}$ [33]
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ULL .	3. Results and discussion	

# 3.1 Optimization of Standard p-In0.4Ga0.6N/ i-In0.4Ga0.6N/ n-In0.4Ga0.6N

The optimization of the InGaN solar cell was started by varying the absorber's thickness from 25 nm to 500 nm; the extracted simulation results are shown in Figures 2(a) and 2(b). Under initial boundary conditions, the fill factors (*FF*), conversion efficiency ( $\eta$ ), short circuit current density ( $J_{sc}$ ), and open-circuit voltage ( $V_{oc}$ ) are 90.42%, 15.33%, 11.29 mA/cm<sup>2</sup>, and 1.50 V, respectively. The  $\eta$  of the solar cell increases up to 16.68% at an absorber thickness of 150 nm. Beyond 150 nm, the  $\eta$  was reduced to 13.50% at the thickness of 500 nm. The  $V_{oc}$  and  $J_{sc}$  also showed a similar trendbas the  $\eta$ . In contrast, the *FF* increased with the absorber thickness. As the absorber thickness gradually increased to 150 nm, more electron-hole pairs of e generated, leading to enigher radiative recombinative rate and forward bias diffusive current, *J*. As the absorber layer thickness exceeded the critical thickness (~150 nm) the  $\eta$  was reduced due to the reduction of the electric field [34] to separate electron-hole pairs. The reduction of radiative reduction of the electric field [34] to separate electron-hole pairs.

recombination reduces the short-circuit current density,  $J_{sc.}$  The relations of  $V_{oc}$  with  $J_{sc}$  and  $J_{o}$ of the solar cell are described in Eq. (8).

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$$V_{oc} = \frac{kT}{q} ln \left[ \frac{J_{sc}}{J_o} + 1 \right],\tag{8}$$



Figure 2. Effects of absorber's (*p*-layer) thickness on (a) fill factors and conversion efficiency, and (b) open circuit voltage and short circuit current density of *p*-*i*-*n* InGaN-based solar cell.

Figures 3(a) and 3(b) show the effects of *i*-layer thickness on *FF*,  $J_{sc}$ ,  $V_{oc}$ , and  $\eta$  of *p*-*in* InGaN-based solar cells. The thickness of the intrinsic (*i*-) layer varied from 25 nm to 1000 nm. As the thickness of the *i*-layer increases to 675 nm, the maximum  $\eta$  of 18.44% was achieved. The *FF*,  $V_{oc}$ , and  $J_{sc}$  of the optimized *i*-layer thickness are 90.91%, 1.50 V, and 13.55 mA/cm<sup>2</sup>, respectively. There is no significant change to the  $\eta$  after 675 nm. In comparison, *FF* shows a slight fluctuation of 0.14% (i.e., from 90.91% at 675 nm to 90.78% at 1000 nm).  $J_{sc}$ and  $V_{oc}$  demonstrated the opposite trend. The  $V_{oc}$  reduced while the  $J_{sc}$  indicased with the intrinse layer thickness, as shown including 3(b). As the thickness of the *i*-layer increases, note photons will be absorbed (Sigenerate electron-hole pairs the to widening the depletion diffusion length. Hence, the carriers' lifetime and the arith current density will increase [35]. The opposite trends of  $J_{sc}$  and  $V_{oc}$  of the *p*-*i*-*n* solar cell can be explained by the  $J_o$  increment across the *i*-layer thickness by the Sah-Noyce-Shockley approximation [36], as expressed in Eq. (9).

$$J_o = \frac{q n_i W}{\sqrt{\tau_e \tau_h}},\tag{9}$$

where q is the elementary charge,  $n_i$  is the intrinsic carrier concentration, W is the width of the depletion zone,  $\tau_e$  and  $\tau_h$  are respectively, the electron and the hole lifetimes. A wider *i*-layer thickness contributed to **a** wider depletion zone, even**ba**lly increasing the saturation **b** rrent density,  $J_o$ . Hence, the  $V_{oc}$  is predicted to be lowered as the thickness of the *i*-layer increases, as shown in Figure 3(b).



Figure 3. Effects of i-layer thickness on (a) fill factors and conversion efficiency, and (b) open circuit voltage and short circuit current density of p-i-n InGaN-based solar cell.

Figures 4(a) and (b) show the effects of the donor (n) layer's thickness on FF,  $J_{sc}$ ,  $V_{oc}$ , and  $\eta$ . As the thickness of the *n*-layer increases from  $Q^{0}$  nm to 350 nm, the  $\eta$  of the Qlar cell increases from  $Q^{0}40\%$  to 18.48%. Beyond 350 pm, there are no changes to the pland FF. The  $J_{sc}$  and  $V_{Q}$  demonstrated the same trend. These values increased and remained constant after a certain *n*-layer thickness. It can be concluded that the *n*-layer is less susceptible to thickness the plane of the plane of the values of the plane of the the *FF* converge at 90.87% after exceeding 300 nm of *n*-layer thickness. It can be explained by the correlation of maximum output power ( $P_{max}$ ),  $V_{oc}$ , and  $J_{sc}$ , as expressed in Eq. (10).

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$$Fill factor (FF) = \frac{P_{max}}{V_{oc} \times J_{sc}},$$
(10)



As the thickness of the *n*-layer increased, the  $V_{oc}$  and  $J_{sc}$  also increased and remained constant across subsequent thicknesses. Hence, the *FF* of the solar cell also became constant across the thicker *n*-layer. In summary, the optimized thickness for the *p*-In<sub>0.4</sub>Ga<sub>0.6</sub>N/ *i*-In<sub>0.4</sub>Ga<sub>0.6</sub>N/*n*-In<sub>0.4</sub>Ga<sub>0.6</sub>N solar cell is 150 nm/ 675 nm/ 350 nm. The *FF*,  $\eta$ ,  $J_{sc}$ , and  $V_{oc}$  for this optimized structure are 90.87%, 18.48%, 13.58 mA/cm<sup>2</sup>, and 1.50 V, respectively.

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**3.2 Effect of carrier densities to standard p-In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*n*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/*i*-**In**<sub>0.4</sub>**Ga**<sub>0.6</sub>**N**/

Figures 5(a) and (b) show the effects of absorber (hole) concentration on *FF*,  $\eta$ ,  $V_{oc}$ , and  $J_{sc}$ . The  $\eta$  increased from 16.00% to 18.74% as the hole concentration ( $N_A$ ) increased from  $1 \times 10^{13}$  cm<sup>-3</sup> to  $1 \times 10^{17}$  cm<sup>-3</sup>. Subsequently, the  $\eta$  reduces to 6.44% at a hole concentration of  $1 \times 10^{21}$  cm<sup>-3</sup>. The increase of hole density ( $\Delta n$ ) improves the carrier lifetime, as expressed in Eq. (11).

$$L = \sqrt{D \frac{\Delta n}{R}},$$
(11)  
where *L* is the diffusion length;  $\frac{\Delta n}{R}$  is the minor of carrier life, *R* is the recombination rate, and  
*D* is the diffusivity. The increment of difference of the solar carrier lifetime increased the diffusion length, *L*. It  
convoluted higher carrier collection rate at the *p*- and *n*- terminal. The  $\eta$  of the solar cell  
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reduced significantly when the hole concentration was more than  $1 \times 10^{18}$  cm<sup>-3</sup>. This anomaly was due to increasing surface recombination velocity in the less passivated absorber layer at a higher doping concentration, subsequently reducing the radiative recombination [37]. The reduction of  $J_{sc}$  was due to the reduction of radiative recombination, while the improvement of  $V_{oc}$  across the  $N_A$  can be explained by the separation of the quasi-Fermi level by Zhao *et al.* [38] in Eq. (12).

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Figure 5. Effects of absorber's carrier density (N<sub>A</sub>) on (a) fill factors and conversion efficiency, (b) open circuit voltage and short circuit current density of *p-i-n* InGaN-based solar cell.

The optimized  $\mathcal{Y}_{4}$  of  $1 \times 10^{17}$  cm<sup>-3</sup> for the absorber was used to optimize the G and *n*layers. As the optier concentration in the *i*-layer increased, *FF* and  $\eta$  remained constant at 90.57% and 8.74%, respectively, until the carrier concentration reached 1×0<sup>16</sup> cm<sup>-3</sup>, as shown in Fighte 6(a). As the carrier concentration of the *i*-layer exceeded 1×0<sup>60</sup> cm<sup>-3</sup>, *FF* and  $\eta$  started by the reduce. The behavior of  $V_{op}$  and  $J_{sc}$  in Figure 6(b) was similar to the absorber layer due to the transmitted of the transmitted to the transmitted transmitted to the transmitted tot the transmitted t the excess carrier concentration. The reduction of the  $\eta$  from 18.74% to 18.54% at higher carrier concentration was due to the changes in quasi-Fermi level and the band energy between *p*-In<sub>0.4</sub>Ga<sub>0.6</sub>N/*i*-In<sub>0.4</sub>Ga<sub>0.6</sub>N and *i*-In<sub>0.4</sub>Ga<sub>0.6</sub>N/*n*-In<sub>0.4</sub>Ga<sub>0.6</sub>N interfaces.

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valence band  $(E_{\nu})$  for *i*-layer with carrier concentration of  $1 \times 10^{19}$  cm<sup>-3</sup> is much greater than  $1 \times 10^{16}$  cm<sup>-3</sup>. The valence band offset (VBO) with the carrier concentration of  $1 \times 10^{16}$  cm<sup>-3</sup> is 2.06 eV, which is much lower than 4.12 eV from the carrier concentration of  $1 \times 10^{19}$  cm<sup>-3</sup>. Hence, the holes require more energy for recombination. Therefore, the  $\eta$  of the standard *p-i-n* solar cell reduces with *i*-layer carrier concentration.

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The effects of electron concentration  $(N_D)$  on the electrical properties of the solar cell are shown in Figures (a) and 8(b). As the  $N_D$  increased, the  $\eta$  increases from 15.78% (918.74%. Beyond the cather concentration of  $1 \times 10^{17}$  crow, the  $\eta$  shows a slight reduction from 18.74% to 18.67%. The behavior of the J<sub>sc</sub> and Vie is similar to the effect of depine concentration on *p*- and *i*-layers. Therefore, it can be concluded that the carrier density has almost no significant of effect on the FF and  $\eta$  of the solar cell after the  $N_D$  exceeds  $1 \times 10^{17}$  cm<sup>-3</sup>. In brief, the optimized results are summarized in Table 3.









Figure 8. Effects of *n*-layer carrier density on (a) fill factors and conversion efficiency, (b) open circuit voltage and short circuit current density of *p*-*i*-*n* InGaN-based solar cell.

# 3.2 Effects of absorber design on the performance of solar cells.

This section proposed solar cells with graded bandgap absorbers to enhance the photon harness effectiveness at the absorber layer. The graded absorber layers were proposed to avoid abrupt change ho the energy bandgap in the solar cell structure as well will be reduce the lattice presented between GaN and OnGaN. Therefore, the strain induced piezoelectric and the strain induced p

	spontaneous polarization can be reduced to improve the carrier collection [22]. The optimize parameters for the standard <i>p-i-n</i> structure shown in Table 3 were used as the reference for absorber layer. For the linear-graded absorber, the absorber layer was changed to the $In_x$					
	<sub>x</sub> N layer with a linear increment of In composition from 0% (GaN) on top to 40% (In <sub>0.4</sub> Ga <sub>0.6</sub> N at the bottom of the absorber layer. The wider bandgap was placed on top to capture photor with higher bandgap energy into the intrinsic ( <i>i</i> ) layer, reducing the thermalization of "he carriers" at higher bandgap energy, and improving the <b>Section</b> -hole generation and co <b>Section</b>					
	due to transferring kinetic energy in the electron via band-to-band impact ionization [39].					
	Table 3. Optimized results for the standard $p$ - <i>i</i> - <i>r</i>	<i>i</i> InGaN soils cell.				
	Structural parameters	Values				
	Player initial thickness	NO nm				
న	<i>i</i> -layer initial thickness	<b>9</b> 75 nm				
, doi.	<i>n</i> -layer initial thickness so.	350 nm				
sille	Acceptor doping density in <i>p</i> -layer, $N_A$	$1 \times 10^{17} \text{ cm}^{-3}$				
nttp	Carrier density of mitrinsic layer, $N_i$	$1 \times 10^{16} \text{ cm}^{-3}$				
·	Donor doping density in <i>n</i> -layer, $N_D$	$1 \times 10^{18} \text{ cm}^{-3}$				
	Open circuit voltage, Voc	1.45 V				
	Short-circuit current density, $J_{sc}$	14.22 mA/cm <sup>2</sup>				
	Conversion efficiency, $\eta$	18.74%				
	Fill factor, FF	90.56%				

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Moreover, the linear-graded absorber was proposed in such a way as to enhance strain relaxation so that the effect of piezoelectric polarization was insignificant to the performance of the solar cell [11]. Song *et al.* (2003) investigated the In grading rate. They found that a higher grading rate will introduce more dislocations into the thin film. Eventually, the probability of dislocation entanglement will increase to impede the lattice relaxation process [40]. Therefore, the step-graded absorber *no-n* was introduced to prevent the possibility of a defect from the different In grading rates. The step-graded absorber *poin* has one layer of GaN defect from the different In grading rates. The step-graded absorber *poin* has one layer of GaN the step-graded absorber of 30 nm. The GaN layer on top of the stepthesiludit.

graded acts as the wide bandgap energy collector as well as the GaN cap to protect InGaN from In desorption during the annealing process. The growth of thick In-rich InGaN thin film is always associated with large 3D-island InN on the growing surface due to Stranski-Krastanov (SK) growth mode [41]. Therefore, a dual temperature GaN cap growth method helps dissolve In-rich 3D-island, spreading InN across the GaN surface more uniformly across the growing surface [42] and preventing In desorbing from the InGaN layer [43]. The step-graded absorber with 10% In increment of each layer helps reduce the logice mismatch between the subsequent layers. It is relatively easier to grow by the MOCVD technique. The band diagrams of different absorber configurations are shown in Figure 9. Linear-graded absorber p-i  $\sigma$  shows a smooth reduction of bandgap energy from 3.42  $\Theta$  (GaN) to 1.99 eV (In<sub>0.4</sub>Ga<sub>0.6</sub> $\Theta$ ). In contrast, the stepproduct of the standard p-i-n structure can only reach a maximum p-i-n structure in the standard p-ilinear- and step-graded absorber *p-i-n* structures improve the  $V_{oc}$ ,  $J_{sc}$ , and  $\eta$  compared to standard *p-i-n* structures, as shown in Table 4. The performance of linear-graded absorber *p-in* was also found to be superior to step-graded absorber *p-i-n*. This is mainly attributed to the lesser abrupt bandgap energy changes during the band-to-band and inter-band absorption and lesser thermalization of energy across the absorber layer through the emission of phonon due to barrier height differences. Throughout this study, the best performance was achieved by linear-graded absorber *p-i-n* structure with FF,  $\eta$ ,  $J_{sc}$ , and  $V_{oc}$  of 90.33%, 19.84%, 14.48

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Table 4. Performance of solar cells with different absorber configurations.

	Solar cell configuration	Standard <i>p-i-n</i>	Linear-graded absorber <i>p-i-n</i>	Step-graded absorber <i>p-i-n</i>
	Open-circuit voltage, Voc (V)	1.45	1.52	1.52
	Short-circuit current density, $J_{sc}$ (mA/cm <sup>2</sup> )	14.22	14.48	14.43
	Conversion efficiency, $\eta$ (%)	18.74	19.84	19.77
	Fill Factor, FF (%)	90.56	<del>م</del> م <sup>90.33</sup>	90.34
https://doi.or	0110.1016/1.ije0.2022.1700 https://doi.org/10.	01611.ije0.2022.770	5:11001.010110.1016/1 <sup>11</sup>	160.2022.100

## 4.0 Conclusion

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In this study, optimization of intermediate In composition homojunction  $In_{0.4}Ga_{0.6}N$  solar cell with *p-i-n* structure was performed by varying thicknesses, carrier concentration, and absorber configurations. The results revealed that the absorber thickness and the hole concentration have profound effects on the solar cell's performance. The graded absorber layer was also proposed to improve solar cell performance. The results showed that the linear- and step-graded absorbers have better conversion efficiency and exhibit QE close to 100% compared to the standard structure. The results also revealed that the performance of the linear-ambied absorber was better conversion efficiency and exhibit QE close to 100% compared to the standard structure. The results also revealed that the performance of the linear-ambied absorber was better can the step-graded absorber due to the lesser abrupt bandgap energy changes and lesser thermalization of energy across the absorber layer. In summary these results implied the protential of intermediate In composition  $In_{0.4}Ga_{0.6}N$ . The implementation of graded In composition absorber configurations opens a new patholor the growth and fabrication of InGaN-based solar cells such as single junction and and more solar cells.



#### Authorship contribution statement

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A.K. Tan: Writing – original draft, conceptualization, and formal analysis. H.U. Manzoor: Conceptualization and software. S.S. Ng: Project administrator, review and editing, supervision, funding acquisition. N.A. Hamzah: Review and editing. M.A. Ahmad: Review and editing. Z. Hassan: Review and editing.

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