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# 1 Shock effects in feldspars: an overview

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## 18 **ABSTRACT**

19 **Feldspars are the dominant mineral in the crust of most terrestrial planetary bodies,**  
20 **including Earth, Earth's moon, and Mars, as well as in asteroids, and thus in meteorites.**

21 **These bodies have experienced large numbers of hypervelocity impact events and so it is**  
22 **important to have a robust understanding of the effect that shock waves exert on feldspars.**  
23 **However, due to their optical complexity and susceptibility to weathering, feldspars are**  
24 **under-utilized as shock barometers and indicators of hypervelocity impact. Here, we**  
25 **provide an overview of the work done on shocked feldspars so far, in an effort to better**  
26 **frame the current strengths and weaknesses of different techniques, and to highlight some**  
27 **gaps in the literature.**

## 28 **1. INTRODUCTION**

29 Hypervelocity impacts result in distinctive microstructural deformation in rocks and  
30 minerals during passage of a shock wave (e.g., French and Short, 1968; Roddy et al., 1978;  
31 Stöffler and Grieve, 2007; French and Koeberl, 2010). This shock metamorphism records the  
32 pressure, temperature, and strain rate conditions experienced by geo-materials during crater  
33 formation. Understanding the shock metamorphic overprint of planetary materials is crucial to  
34 the study of meteorites, terrestrial impact structures, and samples returned to Earth by space  
35 missions. Clear criteria for characterizing and quantifying the degree of shock metamorphism are  
36 required together with an understanding of how shock metamorphism affects different  
37 lithologies.

38 Shock metamorphic effects vary between rock types due differences in the mineralogy  
39 and microstructural characteristics of the target (e.g., French and Koeberl, 2010; Ferrière and  
40 Osinski, 2013 and references therein). This observation has been used to propose multiple shock  
41 classification schemes for different rock types including meteorites (e.g., Fritz et al., 2017),  
42 quartzofeldspathic rocks, sandstones, basaltic-gabbroic rocks, particulate rock material,  
43 chondritic and dunitic rocks (Stöffler et al., 2018). Shock effects are mineralogically selective,

44 and heterogeneous in spatial distribution, such that neighbouring minerals, even if the same  
45 composition, may record different effects. This is due to differences in the shock impedance, a  
46 physical property of matter that describes how efficiently a shock wave propagates through a  
47 material. The mineral grains and pore spaces composing a polymineralic rock have different  
48 shock impedances, hence the shock wave propagates with variable efficiency, and this leads to a  
49 heterogeneous distribution of the maximum shock pressure and resultant shock metamorphic  
50 effects within a sample (e.g., Ogilvie et al., 2011). In addition, the degree and type of  
51 deformation is affected by the mineral's crystal structure and composition (Stöffler, 1972), pre-  
52 impact temperature (e.g., Langenhorst et al., 1992; Huffman et al., 1993; Langenhorst and  
53 Deutsch, 1994; Fritz et al., 2011) and pre-impact stress conditions (Daniel et al., 1997; Sims et  
54 al., 2019). Shock deformation of most minerals is evidenced by a progressive degradation of the  
55 crystal structure until melting occurs.

56         Feldspar group minerals (alkali feldspar and plagioclase) are both chemically and  
57 optically complex. Various publications address shock metamorphism for plagioclase, a major  
58 component in planetary materials, including asteroids, Mars, Earth, and Earth's moon. However,  
59 the K-rich alkali feldspars that are rare in meteorites but abundant in terrestrial rocks, have not  
60 been studied in as much detail. Feldspar group minerals suffer from alteration in a hydrous  
61 environment rendering them challenging to study using solely optical microscopy. Additionally,  
62 pre-existing microstructures, such as twin planes, cleavage planes, and exsolution lamellae,  
63 might be masking or possibly even precluding the formation of shock-generated planar features.  
64 The aforementioned challenges to studying feldspar microstructures coupled with inconsistencies  
65 in the features associated with different shock levels in different rock types (e.g., Stöffler et al.,  
66 2018; Fritz et al., 2017) make the utility of feldspars as a shock thermo-barometer variable.

67           Despite their abundance in rocky planetary materials, feldspars have been under-utilized  
68 for shock barometry compared to quartz and olivine. However, feldspar is becoming increasingly  
69 popular for shock barometry (e.g., Jaret et al., 2014; Kayama et al., 2018; Fritz et al., 2019a). In  
70 this paper, we discuss all impact cratering related effects on the feldspar group, but the reader  
71 should note that only definitive planar deformation features (PDFs), and diaplectic glass are  
72 considered *diagnostic shock effects* on their own (French and Koeberl, 2010). Certain high-  
73 pressure polymorphs can also be diagnostic, but only given a proper geologic context. Diagnostic  
74 shock effects are those that only form via passage of a shock wave with a strain rate and a  
75 pressure-temperature-time (P-T-t) path that cannot be produced by endogenic processes.  
76 Additional microstructures can form during impact cratering events but can also form by other  
77 geological processes. For example, fractures, kink bands, mosaicism, and darkening can inform  
78 the degree to which feldspars have been shocked, but these features can also be produced  
79 endogenically. Even tectonically-induced amorphization is possible, although only under  
80 extreme and rare conditions such as those reported for silica by Janssen et al. (2010).

## 81 **2. Feldspar Group Minerals**

82           Feldspars are the most abundant mineral group in crustal rocks on Earth (Deer et al.,  
83 2001), as well as in the crustal rocks of terrestrial planetary bodies (Papike et al., 1991). The  
84 feldspar group is composed of two solid solutions: alkali feldspar ( $\text{NaAlSi}_3\text{O}_8$ - $\text{KAlSi}_3\text{O}_8$ ) with a  
85 density of 2.62 - 2.55  $\text{g/cm}^3$  and plagioclase feldspar ( $\text{NaAlSi}_3\text{O}_8$ - $\text{CaAl}_2\text{Si}_2\text{O}_8$ ) with a density of  
86 2.62 - 2.76  $\text{g/cm}^3$ . Intergrowths of alkali and plagioclase feldspars are common, such as perthite  
87 (albite lamellae in K-feldspar host) and antiperthite (K-feldspar lamellae in albite). The crystal  
88 system is triclinic or monoclinic, depending on composition, thermal history, and crystallization  
89 temperature. Planar microstructures that are common in feldspar are good cleavage ( $\{001\}$  and

90 {010}) and the formation of twins (Seifert, 1964; Zoltai and Stout, 1984; Xu et al., 2016 and  
91 references therein). Common twins in alkali-feldspar are Carlsbad, with the z-axis as twin axis,  
92 and Baveno, with (021) as twin axis, and a combination of albite and pericline twinning,  
93 resembling a "tartan" pattern (Deer et al., 2001).

94 On planetary bodies with active hydrospheres, plagioclase routinely alters to sericite,  
95 zeolites, or clay group minerals. These products can locally be replaced by secondary calcite, due  
96 to hydrothermal alteration (e.g. Leichmann et al., 2003; Pittarello et al., 2013). A typical product  
97 of hydrothermal alteration of plagioclase is "saussurite", an old term used to indicate a mixture  
98 of zoisite, epidote, sericite, and other components.

### 99 **3. Shock effects in Feldspars**

100 Shocked feldspar from impact structures on Earth, meteorites, and shock recovery  
101 experiments displays a range of shock-induced microstructures including undulatory extinction,  
102 mosaicism, planar microstructures, diaplectic glass, and melted (i.e., flow-textured and/or  
103 vesiculated) feldspar glass (Table 1, e.g., Stöffler, 1967; Stöffler et al., 2018). These effects  
104 result from progressive breakdown of the crystal lattice, and are linked with a specific range of  
105 shock pressures, temperatures, and anorthite content, as determined by high-pressure  
106 experiments (Fig. 1).

107 The state of matter under shock compression can be described with Hugoniot curves  
108 (e.g., Melosh, 1989; Langenhorst, 2002; Langenhorst and Deutsch, 2012; Fritz et al., 2017),  
109 which are depicted, for a specific mineral, in a density (or specific volume) vs. pressure plot.  
110 Figure 2 shows experimental data on anorthosite (rock composed mainly of Ca-rich plagioclase)  
111 derived from particle and shock wave velocity measurements at given shock pressures, with the

112 specific volume calculated via the Rankine-Hugoniot relations. The Hugoniot is not a  
113 representation of a gradual increase in pressure and density, but describes the possible states  
114 achieved by a material that discontinuously jumps across the shock front from the uncompressed  
115 to the compressed state. The Hugoniot Elastic Limit (HEL) of a material indicates the transition  
116 from a purely elastic state to an elastic-plastic state. For feldspars, the HEL has been measured at  
117 3.5-4.5 GPa. Pressures in excess of this range cause permanent changes in feldspar crystal  
118 structures (Ahrens et al., 1973; Grady and Murri, 1976; Huffman and Reimold, 1996). In  
119 addition, the Hugoniot of feldspar is subdivided into Regions I to III, which mark pressures  
120 above which the compressibility of the material during shock changes substantially. These  
121 regions correspond to specific shock metamorphic effects – the physical manifestation of the  
122 pressure-volume work during shock. Hence, shock effects from fracturing, kinking, and  
123 mosaicism, to formation of planar microstructures, diaplectic glass, and melting can be linked to  
124 a range of shock pressures. In order to appropriately extrapolate a shock pressure estimate to the  
125 whole rock, one must consider numerous grains throughout the sample, ideally in a variety of  
126 minerals. In the following sections, we provide an overview of the effects produced in feldspars  
127 during hypervelocity impact, both diagnostic (unique to impact) and non-diagnostic (formed both  
128 during impact and endogenically).

### 129 **3.1. Fracturing**

130 Fracturing is pervasive in rocks affected by impact (Fig. 3). The density of impact-  
131 induced fractures in quartz and feldspars has been found to increase with increasing pressure up  
132 to ~20 GPa (Lambert 1979). However fracture density decreases as pressure increases above  
133 ~20 GPa. Lambert (1979) concluded that the correlation between fracture density and pressure

134 was too weak to be used for quantitative pressure calibration. Fractures are rare or absent in  
135 partly or fully isotropic feldspars (see Section 3.7).

### 136 **3.2. Undulatory extinction**

137 Undulatory (or undulose) extinction is the optical effect of a wave of extinction sweeping  
138 through a single crystal as the microscope stage is rotated with the sample between crossed  
139 polarizers; i.e., the entire crystal does not go in and out of extinction simultaneously (Fig. 4).  
140 This optical effect is the result of bending of the crystal structure without fracturing. Undulatory  
141 extinction is a common effect in minerals that have experienced non-uniform strain, and is not  
142 diagnostic of shock metamorphism, however it can be informative. For example, the amount of  
143 stress that a crystal has experienced can be qualitatively estimated based on how undulatory the  
144 extinction is; i.e., approximately how much variation in extinction angle is displayed within a  
145 single crystal? A higher amount of variation records higher strain. There is potential to quantify  
146 this effect using a Universal stage (U-stage) or electron backscatter diffraction (EBSD)  
147 measurements of the misorientation of the optic axes, however, we are not aware of any  
148 systematic studies attempting to quantify undulatory extinction in feldspars. Such efforts would  
149 be unlikely to develop a robust correlation given the ubiquitous nature of undulatory extinction  
150 as a result of tectonic deformation ( $P \ll 1$  GPa).

### 151 **3.3. Optical Mosaicism**

152 Optical mosaicism is an extinction pattern observed in crystals composed of numerous  
153 subdomains, with differently oriented optic axes (e.g., Dachille et al., 1968; Stöffler, 1972;  
154 Stöffler and Langenhorst, 1994; French and Koeberl, 2010). This extinction pattern is different  
155 from undulatory extinction in which extinction passes smoothly through the crystal during

156 rotation of the microscope stage. In mosaicism, subdomains are only a few micrometers in size,  
157 hence several rotated subdomains overlap in a typical  $\sim 30 \mu\text{m}$  thick thin section. Mosaicism is  
158 best observed in transmitted light between crossed polarizers, where it appears as a patchwork of  
159 extinction domains; with EBSD, as a patchwork of crystallographic orientations; or with single-  
160 crystal or in situ X-ray diffraction (XRD), producing arcs of diffraction spots instead of single  
161 spots. A similar pattern of extinction is also caused by endogenic processes (e.g., Spry, 1969;  
162 Vernon, 1975), so optical mosaicism cannot be used as a unique diagnostic indicator of shock  
163 metamorphism.

#### 164 **3.4. Planar microstructures**

165 Shock related planar microstructures in feldspars include planar fractures (PFs), planar  
166 deformation features (PDFs), deformation bands, and kink bands (e.g., Stöffler, 1967, 1972;  
167 French and Short, 1968; French, 1998). PFs and PDFs in feldspar are not as well characterized as  
168 those in quartz, so we start with a brief discussion of these features in quartz to frame the  
169 discussion. In quartz, PFs are open cracks, typically  $>3 \mu\text{m}$  wide, and spaced  $\sim 15\text{-}20 \mu\text{m}$  apart;  
170 PDFs are lamellae composed of amorphous (or recrystallized)  $\text{SiO}_2$ ,  $< 2 \mu\text{m}$  wide, and spaced  
171  $\sim 2\text{-}10 \mu\text{m}$  apart (e.g., Engelhardt and Bertsch, 1969; Stöffler and Langenhorst, 1994;  
172 Langenhorst, 2002). Both PFs and PDFs are planar and form parallel to rational crystallographic  
173 orientations (French and Short, 1968; French et al., 2004).

174 Early reports of what are now referred to as PDFs in feldspar were described as various  
175 types of lamellae in plagioclase by both Stöffler (1967) and Dworak (1969). Stöffler (1967)  
176 described  $1\text{-}8 \mu\text{m}$  thick isotropic (or low birefringence) lamellae spaced a few micrometers apart,  
177 similar to planar features described in quartz. Dworak (1969) described several types of lamellae,  
178 equivalent to deformation bands, PDFs, and PFs. The lamellae described as being most similar to

179 quartz PDFs are 2-5  $\mu\text{m}$  thick isotropic (or low birefringence) lamellae spaced about 10  $\mu\text{m}$   
180 apart, sometimes decorated with small inclusions. Dworak (1969) described feldspar PFs as up to  
181 10  $\mu\text{m}$  wide open fractures spaced a few micrometers apart and forming preferentially at grain  
182 margins.

183 In feldspar, PDFs occurring along with more widely spaced features (i.e., deformation  
184 bands or twins) form a ladder texture (Fig. 5) (e.g., Stöffler, 1967, 1972; Engelhardt and Stöffler,  
185 1968; French and Short, 1968). Pittarello et al. (2013) reported microstructures in plagioclase  
186 that resemble feather features (FFs) previously described for quartz (Poelchau and Kenkmann,  
187 2001), however no systematic study has yet been conducted on apparent FFs in feldspars. Sets of  
188 thin alternating isotropic lamellae ( $<10 \mu\text{m}$  wide) oriented parallel to the most common  
189 plagioclase twins have been reported as shock-induced microtwins (e.g., Stöffler, 1966; Dworak,  
190 1969; Stöffler, 1972), though it has yet to be confirmed whether the formation of the microtwins  
191 is due to shock metamorphism, nor whether the currently isotropic lamellae originated as twin  
192 lamellae.

193 A great many planar microstructures exist in feldspars to begin with, so using an optical  
194 microscope alone to fully characterize putative PDFs is challenging, and it is better combined  
195 with high-resolution techniques such as electron microscopy. Pickersgill et al. (2017)  
196 investigated planar microstructures in Or-rich alkali feldspar from granitoid rocks from the  
197 Chicxulub impact structure using EBSD combined with transmission electron microscopy  
198 (TEM). EBSD highlighted semi-planar microstructures visible in transmitted light parallel to the  
199 trace of  $\{001\}$ ,  $\{010\}$ , and  $\{100\}$ . Follow up TEM work revealed that the “planes” are in fact  
200 semi-coherent subgrain boundaries that are neither straight nor parallel (spacing varies from 0.2-  
201 0.6  $\mu\text{m}$  between adjacent boundaries), and have no indication of amorphous material, making

202 these subplanar microstructures distinct from the strict definition of PDFs as defined in quartz.  
203 The microstructures parallel to  $\{110\}$  form lamellar subgrains that, despite resembling strain-  
204 induced twins, are too closely spaced and narrow to be twins. Likewise, they are not exsolution  
205 lamellae as they show no chemical variation. Pittarello et al. (2020a) conducted EBSD on  
206 shocked plagioclase grains in a meta-granite from the Manicouagan impact structure, Canada.  
207 Results suggest that amorphization in plagioclase begins either within pre-shock twins or along  
208 shock-induced lamellae that are similar to both PDFs and micro-twins, likely representing  
209 structural failure of specific crystallographic planes during shock. Attempted indexing of these  
210 features using EBSD and U-stage resulted in orientations of  $\{001\}$ ,  $\{010\}$ , and  $\{1\bar{2}0\}$  or  $\{1\bar{3}0\}$   
211 (ambiguous).

212 PDFs in feldspars and quartz are reported to start to form at about the same pressure ( $>8$   
213 GPa, Stöffler, 1971; Ostertag, 1983), and are no longer present in samples subjected to pressures  
214 high enough to cause amorphization of the mineral (24-35 GPa). PDFs in feldspars have been  
215 reported for various impact structures including Ries (Stöffler, 1966; Engelhardt and Stöffler,  
216 1968); Sudbury (French, 1968); Manicouagan (Dworak, 1969; Dressler, 1990; White, 1993);  
217 Gardnos (French et al., 1997); Tenoumer (French et al., 1970; Jaret et al., 2014); Dhala (Pati et  
218 al., 2010); and El'gygytgyn (Pittarello et al., 2013). Overwhelmingly, these are noted in passing  
219 as an accessory to PDFs in quartz, with rare attention or detail given to the PDFs in feldspar. As  
220 a result, the compositional range of feldspars bearing PDFs are rarely noted, but from the limited  
221 examples, PDFs appear to be confined to alkali feldspar or low-Ca ( $<An_{50}$ ) plagioclase  
222 (Langenhorst et al., 1995; Trepmann et al., 2003; Gibson and Reimold, 2005; Nagy et al., 2008;  
223 Pittarello et al., 2013; Jaret et al., 2018), and are generally absent in higher Ca-plagioclase even  
224 for samples in shock ranges expected to form PDFs. For example, at the Mistastin Lake impact

225 structure, PDFs are abundant in quartz, and completely absent in plagioclase ( $An_{31-55}$ )  
226 (Pickersgill et al., 2015b). Similarly, at the Lonar Crater, India, which is a basaltic target with  
227 highly calcic plagioclase, no PDFs have been identified despite abundant examples of diaplectic  
228 glass (Kieffer et al., 1976; Jaret et al., 2015). However, we are aware of two exceptions to the  
229 high-Ca plagioclase resistance to PDFs, both in lunar anorthite, from the meteorite Oued Awlitis  
230 001 (Wittmann et al., 2019), and Apollo sample 10046 (Dence et al., 1970). This indicates a  
231 strong compositional (and therefore crystallographic) effect on whether or not PDFs form in  
232 feldspars. The lack of observed planar features in experimentally shocked plagioclase ( $An_{63}$ ) led  
233 Gibbons and Ahrens (1977) to suggest that the presence of PDFs in plagioclase is not as  
234 diagnostic of shock as maskelynite, and that PDFs would represent a purely local effect.

### 235 **3.5. Shock darkening**

236 Shock darkening is a reduction in the transparency of a mineral in transmitted light (e.g.,  
237 Reimold and Miller, 1989; French et al., 1997; Jaret et al., 2018). This optical phenomenon is a  
238 common shock feature in mafic silicates, but has also been observed in feldspars (Fig. 6). Shock  
239 darkening of silicates is attributed to submicroscopic Fe-bearing droplets, that are mobilized and  
240 concentrated at shock conditions prior to melting (e.g., Rubin, 1992; Moreau et al., 2017). Shock  
241 darkening has been described in terrestrial shocked feldspars at the Gardnos (French et al.,  
242 1997), Manson (Koeberl et al., 1996), Roter Kamm (Reimold and Miller, 1989), and Sudbury  
243 (French, 1968) impact structures. Shock darkening has also been observed in experimentally  
244 shocked andesine, albite, and bytownite (Jaret et al., 2018). In plane-polarized light samples  
245 shocked between 17 and 50 GPa all show a pronounced darkening that is inherent to the samples  
246 and not a surface coating, nor does it correspond to elemental composition at the spatial  
247 resolution of the electron microprobe, which supports the attribution of darkening to sub-

248 micrometer Fe-droplets. To our knowledge no atomic level study of shock darkened feldspars  
249 has been conducted. This effect is superficially similar to clouded feldspar, but can be  
250 differentiated in that in clouded feldspar the features are usually resolvable with an optical  
251 microscope (e.g., Poldervaart and Gilkey, 1954; Whitney, 1972; Smith and Brown, 1988; Putnis  
252 et al., 2007).

### 253 **3.6. Alternate twin deformation**

254 An effect observed in some feldspars is the deformation of only alternate twins in, for  
255 example, a polysynthetically twinned plagioclase. This alternate twin deformation includes the  
256 formation of possible PDFs in every other twin (forming a ladder texture or alternate twin  
257 lamellae, Fig. 5a,b), and the formation of diaplectic glass (alternate twin isotropisation).  
258 Alternate twin effects are interpreted to occur in response to "favorable" crystallographic  
259 orientation of one twin set with respect to the local shock wave (Stöffler, 1966, 1967; Taylor and  
260 Dence, 1969; Robertson, 1973; Dressler, 1990; Jaret et al., 2014; Pickersgill et al., 2015b;  
261 Pittarello et al., 2020a). Jaret et al. (2014) interpreted alternate twin deformation to indicate  
262 relatively low pressure conditions where subtle orientation differences are enough to impede  
263 shock deformation, whereas at higher shock conditions there is so much energy that slight  
264 misorientations do not affect the deformation process. Likewise, Pickersgill et al., (2015a) used  
265 in situ micro X-ray diffraction ( $\mu$ XRD) to find that adjacent twin sets generally deform to the  
266 same degree, which they interpreted to mean that the difference in lattice orientation between  
267 twins relative to the shock wave that causes alternate twin deformation occurs over a very  
268 narrow range of orientations.

### 269 **3.7. Diaplectic glass**

270 Diaplectic glasses, sometimes (in older papers) referred to as *thetomorphic glasses*, are  
271 phases which are optically isotropic (i.e., extinct on rotation of the stage between crossed  
272 polarisers); amorphous to Raman, X-ray diffraction, or other techniques that measure  
273 crystallinity; but still retain the exact chemical composition, and morphology of the precursor  
274 mineral thereby preserving the original rock fabric (Fig. 7) (e.g., Engelhardt et al., 1967; Binns,  
275 1967; Engelhardt and Stöffler, 1968; Hörz and Quaide, 1973). Diaplectic glasses do not exhibit  
276 flow textures or vesiculation, even when observed using high resolution imaging techniques such  
277 as scanning electron microscopy (SEM). Preservation of the internal fabric is indicated by  
278 preservation of, for example, magmatic pyroxene lamellae in maskelynite from the Martian  
279 meteorite Shergotty (Fig. 7) and Fe-oxide “clouding” in diaplectic plagioclase glass from the  
280 Mistastin Lake impact structure (Pickersgill et al., 2015b). Even chemical zonation is preserved  
281 in diaplectic plagioclase glass from both terrestrial impact structures (Jaret et al., 2015) and  
282 meteorites (Treiman and Treado, 1998). The transition from crystalline (i.e., birefringent)  
283 feldspar to diaplectic feldspar glass is documented by grains in which some areas are amorphous,  
284 and some areas remain birefringent. This transition can manifest as alternate twin isotropisation  
285 (Stöffler, 1966, 1971; French, 1998), but is equally found unrelated to twins or other planar  
286 features (Fig. 7C,D, e.g., Pickersgill et al., 2015b; Pittarello et al., 2020a, 2020b). Both partially  
287 and fully isotropic diaplectic glasses are characterized by a lack of fracturing, even when  
288 surrounded by highly fractured minerals (Fig. 7 a,b).

### 289 **3.7.1. Identification**

290 In transmitted light, initial suspicion of diaplectic glass comes from noting a grain that is  
291 constantly extinct between crossed polarizers. In reflected light or during SEM of a polished  
292 surface (thin or thick section), initial inference of diaplectic glass comes from observing a

293 notable pattern of fractures: fractures are rare or absent in partly and fully isotropic feldspar and  
294 fractures in adjacent minerals such as olivine and pyroxene terminate at the interface with the  
295 diaplectic glass (see Fig. 7 a,b).

296 In order to definitively identify diaplectic glasses, it is crucial to ensure that the phase  
297 being examined is 1) amorphous, and 2) the same composition as feldspar. Permanent extinction  
298 under an optical microscope can indicate a) an amorphous phase; b) orientation of an optic axis  
299 of a crystal parallel to the line of sight of the microscope (perpendicular to the thin section  
300 surface); and c) naturally isotropic minerals (e.g., garnet). It is therefore necessary to confirm an  
301 amorphous state using additional techniques such as obtaining an interference figure using the  
302 Bertrand lens of the microscope (no interference figure should be present with an amorphous  
303 phase), or using a U-stage to rotate the sample out of extinction. If the phase is indeed  
304 amorphous, in situ  $\mu$ XRD will reveal a diffuse band across the detector rather than distinct  
305 diffraction spots. Likewise, Raman spectra of amorphous feldspar have two broad peaks of  
306 relatively low intensities at  $\sim 400\text{-}600 \text{ cm}^{-1}$  and  $\sim 1000\text{-}1100 \text{ cm}^{-1}$  (Fig 14). Composition  
307 should be confirmed using an analytical SEM or electron probe microanalysis (EPMA).

### 308 **3.7.2. Structure**

309 The amorphous nature of diaplectic glasses suggests that they have lost the ordered  
310 internal atomic arrangement of the precursor crystal. However, diaplectic feldspar glasses  
311 display several properties that indicate a transitional state between a crystal and a glass produced  
312 by quenching of a molten mineral (i.e., a monomineralic melt glass):

313 1) Diaplectic plagioclase glass has a density and refractive index (RI) intermediate between  
314 a plagioclase crystal and a melt glass of identical chemical composition (Engelhardt et al.

315 1967; Ostertag 1983; Fritz et al. 2019a). Both density and RI decrease with increasing  
316 shock level (Arndt et al., 1982, see Fig. 8);

317 2) Diaplectic glasses maintain substantial intermediate-range structural order, unlike melt  
318 glasses (e.g., Milton and de Carli, 1963; Bunch et al., 1967; Arndt et al., 1982; Ostertag,  
319 1983; Jaret et al. 2015).

320 3) Diaplectic plagioclase glass retains a structural memory of its former crystalline state  
321 (Milton and de Carli, 1963; Bunch et al., 1967; Hörz and Quaide, 1973; Arndt et al.,  
322 1982; Diemann and Arndt, 1984; Jaret et al., 2015).

323 When heated at ambient pressure, diaplectic glasses can recrystallize back into a single  
324 crystal, which is distinctly different from melt glass, which always reverts to a polycrystalline  
325 aggregate (Bunch et al., 1967; Arndt et al., 1982). Sometimes these recrystallized feldspars even  
326 regain undulatory extinction. Further investigations into the memory effects of diaplectic glass  
327 have been investigated using high-pressure experiments and are discussed in Section 4.

328 Structural differences between melt glass and diaplectic glass have been investigated by a  
329 number of techniques. Synchrotron X-ray total scattering experiments of crystalline plagioclase,  
330 natural diaplectic plagioclase glass, and fused plagioclase glass (all of labradorite composition)  
331 show that both glasses are characterized by lack of periodicity beyond distances of 10 Å, unlike  
332 crystalline plagioclase (Jaret et al., 2015). However, in the intermediate range (4–10 Å) the  
333 diaplectic glasses show a higher degree of atomic order than melt glasses (Jaret et al., 2015).  
334 Orientation sensitive measurements on diaplectic glass have suggested retention of orientation  
335 and position of some atoms despite loss of long-range order. Micro-FTIR spectroscopy of  
336 individual naturally and experimentally shocked plagioclase grains is also sensitive to

337 orientation, where peak positions and intensities of vibrational modes in plagioclase can shift by  
338 ~40 wavenumbers, making it difficult to correlate changes in peak intensities or positions with  
339 shock level (Fig. 9). However, the degree to which orientation affects the spectra decreases with  
340 increasing shock level (Jaret et al., 2015). Orientation effects have been found to be a useful tool  
341 for distinguishing between diaplectic and melt-produced glasses. For example, in infrared  
342 spectroscopy diaplectic labradorite glass exhibits one Si-O stretching band, but this peak position  
343 changes by 40 wavenumbers upon rotation of the grain (Fig. 9). Melt glass on the other hand  
344 does not show effects of orientation on spectral position (Jaret et al., 2015). Nuclear magnetic  
345 resonance (NMR) spectroscopy shows the overall spectral shape and pattern of diaplectic  
346 labradorite glass is similar to that of crystalline plagioclase, suggesting a lack of large-scale  
347 differences in structure or silica polymerization. Fused labradorite glass on the other hand has a  
348 larger difference in chemical shift in NMR spectra compared to crystalline labradorite, further  
349 supporting that diaplectic plagioclase glass and fused plagioclase glass are structurally different  
350 (Jaret et al., 2015).

### 351 **3.7.3. Alteration**

352 In a hydrous environment, feldspars and diaplectic glass are susceptible to alteration. As  
353 a result, recrystallized feldspars are typical in terrestrial impactites. The resultant textures are  
354 described as spherulitic or plumose microcrystals, interpreted as evidence of altered diaplectic  
355 glasses during post-shock heating (French, 1998). This effect occurs in the Manson impact  
356 structure, where shocked plagioclase appears to have highly altered alternating twins (Short and  
357 Gold, 1996). Likewise, at the Mistastin Lake impact structure, alternating twins in some samples  
358 have altered to zeolites (Pickersgill et al., 2015b). In both cases, these textures are interpreted as  
359 preferential alteration of diaplectic glass in grains that underwent alternate twin isotropization.

360 **3.7.4. Formation pressure**

361 Fully isotropic diaplectic feldspar glass forms by the passage of a shock wave with  
362 pressures of more than 24-30 GPa and less than ~47 GPa (Table 1, Ostertag, 1983; Jaret et al.,  
363 2018; Fritz et al., 2019a). Both meteorites and terrestrial impactites have documented a  
364 relationship between the formation of diaplectic plagioclase glass from plagioclase in contact  
365 with shock melt veins (Walton et al., 2016; Sharp et al., 2019). In these cases, plagioclase further  
366 from melt veins, or in contact with thin (1-4  $\mu\text{m}$ ) veins remains crystalline, but plagioclase in  
367 contact with melt veins has become isotropic. The formation of diaplectic glass is therefore  
368 attributed to the elevated temperature gradient caused by contact with shock melt, combined with  
369 the large deviatoric stresses experienced by minerals along vein margins. These natural  
370 observations agree with experimental results that demonstrate that both elevated temperature and  
371 shear stress reduce the pressure at which plagioclase transforms to diaplectic glass (e.g., Kubo et  
372 al., 2010; Daniel et al., 1997). There is a strong compositional effect on the pressures associated  
373 with amorphization, where more calcic plagioclase becomes amorphous at lower pressure  
374 conditions (Ostertag, 1983; Angel, 1994; Fritz et al., 2019a; Jaret et al., 2018). Most likely this  
375 trend reflects the Si-Al polymerization rather than Ca content directly (Angel, 1994; Jaret et al.,  
376 2018). There has been significantly less work on quantifying the differences in shock response  
377 between the alkali-series feldspars. Ostertag (1983) reported amorphization of experimentally  
378 shocked sanidine and orthoclase at pressures of ~30 and ~32 GPa, whereas microcline retained  
379 some birefringence up to pressures of 45 GPa.

380 The transitional regime between diaplectic and melt glasses results in shock-temperatures  
381 in plagioclase that are not high enough to cause melting of the whole grain, but are just high

382 enough to cause eutectic melting along plagioclase/pyroxene interfaces (Fig. 7) and assimilation  
383 of minerals with a lower melting point.

### 384 **3.8. Melting**

385 At higher shock pressure monomineralic melt glass forms with the same composition as  
386 the original feldspar (e.g., Stöffler, 1972, 1967). An example of fully melted plagioclase glass is  
387 shown in Fig. 10, in the strongly shocked Martian meteorite ALH 77005. Petrographically  
388 melted feldspar glass is significantly different from diaplectic feldspar glass in that melt glass  
389 displays flow textures and/or vesicles.

390 Upon cooling, if the pressure pulse is longer than the undercooling, high pressure phases  
391 can crystallize from the melt (e.g., Sharp and DeCarli, 2006; Walton et al., 2014). Small  
392 occurrences of plagioclase-composition melt glass can be identified by careful high-resolution  
393 backscatter electron imaging (El Goresy et al., 2013). Care must be taken to ensure that apparent  
394 feldspar-composition glasses are not whole-rock melts using EPMA or analytical SEM. Feldspar  
395 melting is associated with local peak shock pressure of  $\sim 45\text{-}50$  GPa, however melting can occur  
396 in porous targets shocked to lower pressures, where the shock kinetic energy is mostly converted  
397 into heat (e.g., Kieffer, 1971; Davison et al., 2010).

### 398 **3.9. High pressure polymorphs**

399 The temperatures during shock are such that weakly to strongly shocked rocks can serve  
400 as a heat sink allowing localized zones of shock melt to quickly cool by assimilation and thermal  
401 conduction during elevated shock pressure conditions. Therefore, localized zones of shock melt  
402 (10s to 100s of micrometers in size) provide the pressure-temperature-time path (P-T-t) allowing  
403 for the formation and preservation of a variety of high-pressure phases (Tomioka and Miyahara,

404 2017; Fritz et al., 2017). These local zones of shock melt serve as natural crucibles mimicking  
405 the conditions in the Earth's lower crust or mantle. High-pressure minerals are not common in  
406 Earth's upper crust, but they are produced endogenically and therefore can only be related to  
407 shock when observations are combined with relevant geologic context (French and Koeberl,  
408 2010). A list of high-pressure minerals related to shocked feldspar is given in Table 2, together  
409 with their chemical formulae, experimentally derived pressure and temperature (P-T) stability  
410 fields and references. High-pressure phases have been identified in most types of meteorites, in  
411 various rocks from terrestrial impact structures, and synthesized in high-pressure laboratory  
412 experiments. This indicates that formation and preservation of high-pressure phases is a typical  
413 result of impact cratering events.

### 414 **3.10. Sieve-texture and checkerboard feldspar**

415 "Checkerboard" feldspar is a checkered-looking pattern of subgrain domains found  
416 within feldspar clasts in impact melt rocks (Fig. 11). Checkerboard texture has the appearance of  
417 crystallographically controlled individual rectangular domains separated by melt products (e.g.,  
418 Grieve, 1975). Internal features such as twinning and extinction angle are preserved across the  
419 grains demonstrating that they were once a single crystal (Fig. 11). This texture was originally  
420 termed "checkerboard" by Grieve (1975), who described unshocked partly digested plagioclase  
421 inclusions in a melt rock with a sieved, "checker-board" appearance due to thermal effects of the  
422 surrounding melt. Bischoff and Stöffler (1984) first proposed a detailed scenario for  
423 development of sieve texture in clasts in impact melt rock from the Lappajärvi impact structure  
424 (Finland). Their scenario essentially regarded checkerboard texture as the result of rapid,  
425 crystallographically-controlled recrystallization from diaplectic feldspar glass, comparable to the

426 suggested formation of ballen quartz (e.g., Ferrière et al., 2010). As such, the texture would  
427 represent positive evidence of transient shock pressures exceeding 25-35 GPa.

428         Whitehead et al., (2002) revised the formation of checkerboard texture based on feldspar  
429 clasts in melt rocks from the Popigai impact structure. Their scenario more closely resembles the  
430 disequilibrium melting of plagioclase observed during the rapid decompression of intermediate  
431 composition magmatic systems, resulting in typical (i.e., igneous) “sieve-textured” feldspar  
432 (Nelson and Montana, 1992). Whitehead et al. (2002) noted similar textures in clinopyroxenes in  
433 the Popigai impact melt rocks, which are also commonly associated with so-called “sieve-  
434 textured” feldspars in volcanic rocks.

435         A texture similar to checkerboard feldspar occurs in volcanic rocks, and is also referred  
436 to as “sieve texture”. The images that accompany the igneous definition of “sieve texture” (Klein  
437 and Philpotts, 2016) imply crystallographic control of the melt channels. This can lead to  
438 confusion when the terms “sieve-texture” and “checkerboard texture” are used interchangeably,  
439 because there are two genetic interpretations for a similar appearance: i) volcanic chemically  
440 driven disequilibrium melting (Nelson and Montana, 1992) and ii) impact-induced  
441 crystallographically controlled disequilibrium melting (Whitehead et al., 2002).

442         Whitehead et al., (2002) attempted to distinguish impact melting from igneous melting  
443 based on the optical crystallography and X-ray element mapping of subgrains within the feldspar  
444 clasts. However, it is not clear from their observations that distinguishing these two processes is  
445 possible (Harris et al., 2019). Therefore, it has not been convincingly shown that the sieve  
446 texture of igneous petrologists and the sieve or checkerboard texture of impact petrologists are  
447 the result of different processes. Consequently checkerboard or sieve-texture feldspar, while  
448 common in both volcanic and impact melt rocks, should not be used as a unique diagnostic

449 criterion of either process unless significant progress is made in differentiating igneous sieve  
450 texture from impact-induced sieve or checkerboard texture.

#### 451 **4. Experimental Shock Calibrations**

452 High-pressure experiments enable formation conditions (i.e., pressure, temperature,  
453 timing) to be assigned to rock and mineral samples based on the microscopic textures produced.  
454 Such experiments are conducted using four main techniques: 1) shock experiments; 2) shock  
455 recovery experiments; 3) rapid compression experiments; 4) static compression experiments.  
456 These different techniques (1-4 in previous sentence, summarized in Table 3) create different  
457 physical conditions. As a result, the same high-pressure effects (e.g., amorphization) may  
458 develop at different pressure and temperature conditions depending on the experimental process  
459 and the characteristics of the test material (i.e., initial grain size, composition, porosity, presence  
460 of structural defects, etc.). Therefore, the reported experimental data need careful evaluation  
461 before comparing results between different experiments and their application to natural shock  
462 metamorphism. The experimental setups mentioned previously could influence the pressures at  
463 which specific microstructural deformation is observed in several ways:

464 1) By studying the samples in a compressed or decompressed state, i.e., in situ during static  
465 pressure experiments or after decompression to ambient pressure from static pressure  
466 experiments (Daniel et al., 1997; Johnson et al., 2002a, 2003; Kubo et al., 2010; Jaret et  
467 al., 2018).

468 2) By applying hydrostatic or non-hydrostatic conditions to the sample (Daniel et al., 1997)

469 3) By changing the rate at which the sample is compressed (e.g., Sims et al., 2020)

- 470 4) By changing the duration of the high-pressure conditions (Huffman et al., 1993; Kubo et  
471 al., 2010; Ogilvie et al., 2011; Fritz et al., 2019a); and,
- 472 5) By selecting the pre-experiment or the syn-experiment temperature of the experimental  
473 setup (Gratz et al., 1992; Huffman and Reimold, 1996; Kubo et al., 2010; Tomioka et al.,  
474 2010; Fritz et al., 2011, 2019a).

475 In addition, experimental measurements of the shock wave and particle velocity  
476 parameters allow for calculation of the pressure, density, and energy of the compressed material  
477 using the Rankine–Hugoniot relations (Fig. 2). These different experimental approaches have  
478 provided a variety of complementary information on material properties during shock  
479 compression and after decompression to ambient pressures (see Section 3).

480 The wealth of experimental data provides space for a variety of opinions among shock  
481 researchers. For simplicity, the discussion can be separated into two end-member interpretations:

- 482 1) *The shock metamorphic record in naturally shocked rocks is in conflict with experimental*  
483 *results obtained by the different types of experiments* (Stöffler et al., 1991, 2018; Chen et  
484 al., 1996; Sharp and DeCarli, 2006; Gillet et al., 2007; El Goresy et al., 2013). Various  
485 researchers agree that there is a substantial difference between results from static pressure  
486 and dynamic (shock) pressure experiments, with different authors favoring results from  
487 either one or the other (Stöffler and Langenhorst, 1994; Sharp and DeCarli, 2006; Gillet  
488 et al., 2007; Kubo et al., 2010; Jaret et al., 2018; Sims et al., 2020). The argument is that  
489 static experiments are too slow and do not generate a shock wave, because the pressure  
490 increases stepwise and is sustained over relatively long time scales of seconds to minutes.  
491 The dynamic experiments occur over extremely short timescales ( $\sim 1$  ns –  $1$   $\mu$ s) and  
492 cannot sustain the shock conditions for the same duration as in a natural impact event. In  
493 the case of reverberation shock experiments, the maximum pressure is not achieved in a  
494 single shot as is the case in natural shock waves (Langenhorst, 2002). Thus, neither static

495 nor dynamic experiments reproduce the duration and loading path as inferred for the  
496 natural case (Stöffler and Langenhorst, 1994; Sharp and DeCarli, 2006; Gillet et al.,  
497 2007), which is typically in the temporal range of milliseconds to rarely a few seconds  
498 (Langenhorst, 2002; Sims et al., 2019; Bowling et al., 2020). These interpretations are  
499 due to a well-discussed discrepancy between the pressures associated with  
500 transformations observed in static experiments compared to those in dynamic  
501 experiments. For example, diamond anvil cell (DAC) experiments and shock experiments  
502 from the same starting material show a ~7-9 GPa difference in amorphization pressures  
503 (Jaret et al., 2018; Sims et al., 2020), consistent with interpretations compiled from  
504 literature of either isolated shock or static experiments (Daniel et al., 1997; Johnson et al.,  
505 2002a, 2003). Sims et al. (2020) recognized memory effects in the DAC experiments  
506 suggesting two amorphization points – one where the sample appears amorphous while  
507 still under compression, and a higher point where the sample remains amorphous after  
508 decompression. The higher pressure is closer to those of the shock experiments, which  
509 may in part help reconcile the discrepancies. Recent “rapid compression” experiments  
510 using membrane DAC setups that allow for control on strain-rate have shown that strain  
511 rate has a significant effect on the pressures at which plagioclase deforms (Sims et al.,  
512 2019).

513 2) *The shock metamorphic record in naturally shocked rocks agrees with the results*  
514 *obtained by both shock and static pressure experiments* (Fritz et al., 2017). Shock  
515 recovery experiments show that with increasing shock pressure different types of shock  
516 deformation effects develop in olivine and plagioclase, respectively. These two different  
517 thermo-barometers provide consistent results if studied in naturally shocked rocks such as  
518 shocked igneous Martian meteorites (Fritz et al., 2005a). These consistent results mean,  
519 that either both olivine and plagioclase are affected by potential differences between  
520 shock recovery experiments and the natural case to the same degree, or that shock  
521 recovery experiments accurately reproduce the conditions in a natural impact event.  
522 Shock reverberation experiments and hydrostatic pressure experiments produce an  
523 amorphous material from plagioclase and quartz as long as both types of experiments  
524 achieved the same pressure and temperature conditions and the samples are studied in a  
525 decompressed state (Fritz et al., 2019a). The shock-loading path achieved in shock

526 reverberation experiments is a reasonable analog for the diffuse shock front traveling  
527 through lithological units composed of different types of minerals, porosity and fractures  
528 during a natural impact event (Fritz et al., 2011). Shock reverberation experiments allow  
529 well-defined shock pressures to be applied to a relatively large sample volume (Müller  
530 and Hornemann, 1969; Fritz et al., 2011). Most problems regarding the shock thermo-  
531 barometry of meteorites are not related to experimental limitations but mainly result from  
532 inconsistent interpretation of the observational evidence (Fritz et al., 2017, 2019a). This  
533 includes: a) Misunderstandings regarding the duration of isobaric shock pressure  
534 conditions in naturally shocked rocks, which mostly implies that high-pressure phases  
535 form in (and near) local shock melt zones during declining shock pressures, b) A  
536 mineralogically inconsistent definition of the S6 shock level leading to an incorrect shock  
537 pressure assignment of all chondrites containing high-pressure phases (Fritz et al., 2017).

538 Reconciling the different interpretations of the shock thermo-barometry of rocks remains an  
539 ongoing area of active research and discussion. For example, a recent study of the shock stage  
540 distribution of 2280 ordinary chondrites confirmed that all S6 chondrites need reclassification  
541 (Bischoff et al., 2019). Despite these discussions, there is agreement on the types of shock  
542 deformation effects in feldspar and the order in which these shock effects occur with increasing  
543 pressure (see Section 3).

## 544 **5. Quantifying shock features in feldspars**

545 Multiple techniques have been applied to begin to quantify shock effects in feldspars.  
546 These techniques are summarized in Table 4, and discussed in the following sections.

### 547 **5.1. Measuring planar microstructures in feldspars**

548 In quartz, the number and orientation of PDF sets have been correlated with specific  
549 pressures through shock loading experiments allowing for refined pressure estimates from  
550 measurement of PDF orientations (e.g., Hörz, 1968; Huffman and Reimold, 1996; Ferrière et al.,

551 2009). Such a correlation has not been developed for feldspars due to the complexities involved  
552 in PDF formation, identification, and orientation measurement in this mineral group.

553         The universal stage (U-stage) is the most common method of determining PDF  
554 orientations in quartz, and this method has been used with feldspars, though it is significantly  
555 more challenging (Stöffler, 1967; Dworak, 1969; Pittarello et al., 2020a). Measuring the  
556 crystallographic orientation of any planar element in feldspars is complicated compared to  
557 measuring the same features in quartz because feldspar is optically biaxial, and the crystal  
558 structure (and therefore optical properties) of feldspars varies with composition through the solid  
559 solution from Ca-rich to Na-rich to K-rich end members (Doeglas, 1940; Haff, 1940; Fairbairn  
560 and Podolsky, 1951). Therefore in addition to the measurements themselves being complex,  
561 accurately indexing the measured planes requires use of the specific indexing template for the  
562 composition of feldspar under investigation; examples of such indexing templates can be found  
563 in Stöffler (1967) and Dworak (1969). These factors result in orientation measurements of planar  
564 features in feldspars using solely a U-stage to be both time-consuming and error-prone. For  
565 example, Dworak (1969) measured 113 lamellae in 51 crystals of labradorite using the U-stage,  
566 but due to the inherent difficulty in using a U-stage with feldspars, the lowered birefringence of  
567 some host crystals, and the changes in An content (even within a single zoned crystal) the  
568 assignment of the orientation of these lamellae had a large error. The most common PDF  
569 orientations measured in feldspars are summarized in Table 5. Both Stöffler (1967) and  
570 Robertson (1966) came to the conclusion that planar deformation features in feldspar form over  
571 such a narrow range of pressures that it would not make sense to correlate orientations with peak  
572 pressures. By implication, the mere presence of planar deformation features in feldspar,  
573 regardless of their orientation, would be a sensitive shock barometer.

574 An additional issue in investigating PDFs in feldspar with the U-stage is optimizing the  
575 choice of glass hemispheres. The glass hemispheres each have a fixed refractive index, and for  
576 accurate measurements the refractive index must be the same as the host crystal. However, the  
577 refractive index of plagioclase ranges from 1.53 (albite) to 1.59 (anorthite) and for K-rich  
578 feldspar, it ranges from 1.518 (orthoclase) and 1.539 (microcline). The hemispheres used for  
579 measurements in quartz have refractive index 1.554, which is between the end member values  
580 for plagioclase, but is too high for K-feldspar. Hemispheres with closer refractive indices exist,  
581 but are not readily available to order and are difficult to match to the grain of interest. However,  
582 the error induced by the difference in refractive index between the investigated grain and that of  
583 the glass hemisphere pair can be calculated, and used to correct the data (e.g., Reinhard, 1931).

584 To circumvent some of the difficulties encountered by using the U-stage alone, EBSD  
585 has recently been combined with U-stage or focused ion beam (FIB) milling and BSE imaging to  
586 measure orientation of PDFs. EBSD provides crystal orientation of the host mineral  
587 circumventing some of the challenges of defining a unique orientation to a biaxial mineral using  
588 the U-stage alone. EBSD can also be used to determine a family of planes to which the trace of a  
589 plane most likely belongs. However, to complete the orientation measurement in 3D, a “dip”  
590 measurement is acquired using the U-stage (Pittarello and Koeberl, 2017; Pittarello et al., 2020a,  
591 2020c) or by making a FIB trench perpendicular to the trace of the plane and measuring the dip  
592 off the resultant image (Pickersgill and Lee, 2015). Thus far, using the FIB technique for this  
593 purpose has been demonstrated only for quartz, however the same principle should work for  
594 feldspars. Currently, transmission electron microscopy (TEM) is necessary for a definitive  
595 identification of the crystallographic orientation along which PDFs develop in feldspar.

## 596 **5.2. Density, refractive index, and birefringence**

597 Density, refractive index (RI), and birefringence decrease as pressure increases until the  
598 mineral becomes amorphous (Stöffler, 1974; Stöffler and Langenhorst, 1994). At shock  
599 pressures <20 GPa the RI of shocked crystalline (birefringent) feldspar is only slightly lower  
600 than the RI of unshocked feldspar (Kleeman, 1971; Stöffler, 1974; Robertson, 1975). At  
601 pressures greater than ~24 to 35 GPa (depending on chemical composition), feldspars are  
602 completely converted to an amorphous state (diaplectic glass). Once this has occurred, and for  
603 each type of feldspar over a very narrow pressure range, the RI of diaplectic feldspar glass drops  
604 sharply (Ostertag, 1983). Shock pressures of ~45 GPa produce an amorphous material with a RI  
605 matching that of a melt glass of identical chemical composition to feldspar. As an exception,  
606 diaplectic anorthite glass produced in 41.5 GPa shock experiments displayed a RI that was  
607 higher than that of unshocked crystalline anorthite (Fritz et al., 2019a). The well-tested  
608 relationship between the refractive index of diaplectic feldspar glass, the An content, and the  
609 shock pressure is illustrated in Figure 9 (e.g., Binns, 1967; Engelhardt et al., 1970; Stöffler et al.,  
610 1975; Robertson, 1975; Ostertag, 1982; Fritz et al., 2019a).

### 611 **5.3. X-Ray Diffraction (XRD)**

612 Two types of X-ray scattering experiments have been applied to shocked feldspars: X-ray  
613 diffraction (XRD) experiments by powder or single-crystal XRD, in situ  $\mu$ XRD (in thin section  
614 and hand sample), and high-energy total X-ray scattering experiments using a synchrotron.  
615 Contributions by total-scattering experiments are discussed in section 3 as they relate to  
616 amorphous phase studies. At pressures too low to induce amorphization or phase-change, XRD  
617 patterns show peak broadening. There are two types of peak that can broaden in this fashion,  
618 peaks in the chi direction (visible with single-crystal or  $\mu$ XRD) and peaks in the  $2\theta$  direction  
619 (visible with single-crystal,  $\mu$ XRD, and powder diffraction).

620 Shock experiments on andesine (4-10 GPa) and oligoclase (3-34 GPa) by Hörz and  
621 Quaide (1973), showed that the degree of crystal lattice damage is correlated with shock  
622 pressure. Furthermore, by using single crystal XRD to measure the length of streaks in the chi  
623 direction (i.e., strain-related mosaicity) they could start to quantify the amount of strain recorded  
624 by the crystals and use that as a proxy to quantify shock level. A correlation between streak  
625 length and shock level was also documented in a single-crystal XRD study of plagioclase from  
626 the Charlevoix impact structure (Walawender, 1977). Pickersgill et al. (2015a) built upon this  
627 work by using in situ  $\mu$ XRD to quantify strain-related mosaicity by measuring the full width at  
628 half maximum (FWHM) of the intensity vs. chi angle pattern, and compared those measurements  
629 to optical signs of shock in samples of labradorite and andesine from the Mistastin Lake impact  
630 structure, and lunar anorthite collected during the Apollo program. As with the single-crystal  
631 studies discussed previously, peak broadening in the chi dimension increased with increased  
632 optical signs of shock, up to amorphization of the sample (i.e., formation of diaplectic glass) – at  
633 which point the X-rays produced a diffuse band on the detector. Powder XRD analyses of  
634 shocked and unshocked sanidine by Kayama et al. (2012) reported almost no change in  
635 diffraction peak intensity between samples. However, with increased shock pressure they did  
636 find slight peak broadening in  $2\theta$ , followed by a sudden change to no detectable peaks (i.e., an  
637 amorphous pattern) for amorphous samples.

638 Strain-related mosaicity, as measured using in situ  $\mu$ XRD or single-crystal XRD, shows  
639 more variation in shock level of feldspars prior to amorphization than optical determinations, and  
640 likely could be useful in subdividing the lower end of the shock scale (Pickersgill et al., 2015a).  
641 However, in order to properly quantify shock via strain-related mosaicity, more work needs to be  
642 done such as correlating results with those from other techniques (Raman for example) and via

643 application to a wider variety of feldspar compositions and, ideally, also to experimentally  
644 shocked samples. Caution must be applied when using XRD to measure shock, as the features  
645 described in the previous paragraphs indicate strain, and therefore, without contextual  
646 information, it is impossible to differentiate the cause of the strain (i.e., shock or tectonism).

#### 647 **5.4. Thermal Infrared Absorption Spectroscopy**

648 Thermal infrared absorption spectra of naturally shocked feldspars show lower intensity  
649 and less spectral detail than unshocked feldspar. Both detail and intensity continue to decrease as  
650 pressure increases as a result of increased glass content and progressive lattice disordering  
651 (Lyon, 1963; Bunch et al., 1967; Stöffler and Hornemann, 1972; Stöffler, 1974; Arndt et al.,  
652 1982; Ostertag, 1983; Johnson et al., 2002b, 2003; Jaret et al., 2015, 2018). Crystalline feldspars  
653 exhibit infrared reflectance peaks between 950 and 1150  $\text{cm}^{-1}$  (reststrahlen bands) that  
654 correspond to Si-O stretching modes of the  $\text{SiO}_4$  tetrahedra. Additional peaks occur between 700  
655 and 850  $\text{cm}^{-1}$  corresponding to Si-bridging oxygen modes, and between 400 and 600  $\text{cm}^{-1}$   
656 corresponding to O-Si-O bending modes (Iiishi et al., 1971; Okuno, 2003). Bulk IR emissivity of  
657 experimentally shocked albite shows a progression of change to the IR spectra in the form of  
658 decrease in overall band intensity relative to background, as well as loss of the low wavenumber  
659 peaks. Spectroscopically, plagioclase that has been transformed to amorphous material (either  
660 by melting or as diaplectic glass) exhibits one broad peak reflecting the Si-O stretching vibration.

661 Micro-FTIR spectroscopy of individual plagioclase grains, both naturally and  
662 experimentally shocked, has an added complexity in that infrared spectra are orientation-  
663 sensitive, where peak positions and intensities of vibrational modes in plagioclase can shift by  
664  $\sim 40$  wavenumbers, depending on orientation, making it difficult to correlate changes in peak  
665 intensities or positions with shock level (Fig. 9). However, the degree to which orientation

666 affects the spectra decreases with increasing shock level (Jaret et al., 2018). Furthermore,  
667 orientation effects have been found to be a useful tool for distinguishing between diaplectic and  
668 melt-produced glasses. Diaplectic glass retains orientation effects in IR, despite its amorphous  
669 nature. Melt glass on the other hand does not show effects of orientation on spectral position.

## 670 **5.5. Cathodoluminescence (CL)**

671 The main effects of shock metamorphism on cathodoluminescence (CL) spectra of  
672 plagioclase are 1) decreasing luminescence with increasing amorphization (e.g., Fig. 12, Kaus  
673 and Bischoff, 2000; Götte, 2009; Götze, 2009; Pittarello et al., 2015; Kayama et al., 2018;  
674 Pittarello et al., 2020b); 2) the normal band emission at ~ 550-570 nm is often shifted to ~ 630  
675 nm (Kayama et al., 2009a); 3) diaplectic plagioclase glass produces a broad band around 350 nm  
676 (Götte, 2009; Pittarello et al., 2015; Gucsik et al., 2004). Intriguingly, and in contrast to  
677 plagioclase, alkali feldspar shows increased CL intensity with increasing shock pressure  
678 (Kayama et al., 2012).

679 CL spectral analysis has been proposed to quantify shock effects and correlate with  
680 pressure (e.g., Kayama et al., 2012, 2018). However, the correlation between shock and signal  
681 intensity is complex because the intensity of the luminescent centers is regulated by numerous  
682 variables, including the abundance of activators, which can be trace elements, point defects, or  
683 both, and the abundance of quenchers, which would reduce the effect of the activators. Even in  
684 unshocked feldspars, the nature of activators is not always clear, so imaging studies must be  
685 accompanied by detailed back-scattered electron imaging in order to properly interpret the  
686 microstructures (e.g., Finch and Klein, 1999; Götze et al., 2000; Parsons et al., 2015). As a result  
687 of the complex mechanisms involved in luminescence, and therefore the relationship between  
688 luminescence and shock intensity, CL is generally viewed as less reliable than other techniques

689 (e.g., XRD or Raman) for quantifying shock intensity, though it remains one of the most useful  
690 tools for investigating microstructures.

## 691 **5.6. Raman spectroscopy**

692 Raman spectroscopy allows identification of high-pressure phases (Section 3.9) and  
693 spectra show that increased shock corresponds to a progressive decrease in peak intensity and  
694 increase in peak width compared to unshocked feldspar (Fig. 13, e.g., Fritz et al., 2005b; Jaret et  
695 al., 2014). Peak positions and width are somewhat dependent on composition, so it is challenging  
696 to identify quantitative metrics in Raman spectra that uniquely identify shock.

697 At pressures < 29 GPa, shocked plagioclase has lower intensity and broader FWHM of  
698 characteristic Raman bands compared to unshocked plagioclase. At higher pressure (>~40 GPa)  
699 Raman spectra exhibit a broad plateau as a result of amorphization (Fig. 13, Jaret et al., 2018). At  
700 much higher pressure (>45 GPa) high post-shock temperatures lead to recrystallization, and  
701 reappearance of the characteristic Raman spectra (Fritz et al., 2005a,b).

## 702 **5.7. Photoluminescence spectroscopy**

703 Photoluminescence spectroscopy is a technique, complementary to Raman spectroscopy,  
704 which has only been applied to shocked plagioclase by Pittarello et al. (2020b).

705 Photoluminescence allows quantification of the “damage” in the lattice (i.e., amorphization) as  
706 percent of disordered structure within the investigated volume (a few  $\mu\text{m}^3$ ). Similar to peak  
707 broadening in Raman spectra,  $\text{Nd}^{3+}$  luminescence bands get wider and are less defined with  
708 increased levels of amorphization.

## 709 **5.8. Nuclear Magnetic Resonance (NMR) Spectroscopy**

710 NMR spectroscopy uses nuclear spin transitions as a measure of the local chemical  
711 bonding and coordination environment (Stebbins and Xue, 2014). Because NMR is sensitive to  
712 the entire local chemical environment and not just long-range ordering, NMR is particularly  
713 useful for measuring the structure of amorphous materials. In amorphous silicates, NMR spectra  
714 of  $^{29}\text{Si}$  (“ $^{29}\text{Si}$  NMR”) are sensitive to the number and length of Si-O bonds, types of nearest  
715 neighbor atoms, and coordination of Si – all of which can change during shock deformation or if  
716 high-pressure phases are produced (Jaret et al., 2015).

## 717 **6. Ongoing debates**

### 718 **6.1. Formation mechanism of diaplectic glass**

719 The details of diaplectic glass formation are the subject of much debate. There are, in  
720 general, two main hypotheses: 1) rapid quenching of monomineralic melt at high pressure, or 2)  
721 solid-state structural collapse/destruction of atomic ordering of the mineral. Some details of these  
722 two formation hypotheses are explained here.

#### 723 **6.1.1. Melt formation**

724 Grady (1977) interpreted diaplectic glass to form by quenching of a high-density melt  
725 phase upon pressure release. Such a phase would have solidified so rapidly that flow, which  
726 would normally erase the morphology and texture of the precursor minerals, did not have time to  
727 occur. In this scenario, local temperature spikes due to heterogeneous shock compression, cause  
728 local melting despite the temperature in the majority of the rock remaining below the liquidus  
729 (Grady et al., 1975; Grady, 1980). A key aspect of Grady’s scenario for formation of diaplectic  
730 glass is the dissipation of strain energy by the development of shear bands. Microstructures  
731 similar to shear bands have been observed in recovered samples from rapid compression

732 experiments (Sims et al., 2019). Similarly, Arndt et al. (1982) suggested that diaplectic glass  
733 forms from shock-induced melting, but that the duration of high temperature is short enough that  
734 the liquid transition is incomplete and the disordered transitional state is ‘locked in’. Stöffler and  
735 Langenhorst (1994) explained the formation of diaplectic quartz glass by shock melting and  
736 quenching during decompression. The concept of diaplectic glass being a high-pressure melt was  
737 later adopted to explain the formation of maskelynite by Chen and El Goresy (2000) and El  
738 Goresy et al. (2013).

### 739 ***6.1.2. Solid state formation***

740 The amorphous structure, and absence of flow textures, suggests a near instantaneous  
741 change from a crystalline structure to glass without melting (e.g., De Carli and Jamieson, 1959;  
742 Engelhardt and Stöffler, 1968). Ahrens et al. (1969) used shock experiments on plagioclase to  
743 suggest that diaplectic glass forms by solid-state release from a high-pressure phase during  
744 decompression. Williams and Jeanloz (1988) also favored a scenario where diaplectic glass is  
745 produced by reversion from a higher-pressure phase, which, similar to the scenario of Hemley et  
746 al. (1988), calls for high pressure-induced high-coordinated glasses. Hemley et al.’s assumption  
747 of a high-coordinated phase during shock compression rests on the experimentally determined  
748 Hugoniot curve where the steeper high-pressure branch (region III in Fig. 2) indicates a density  
749 during shock compression similar to the density expected of a compressed hollandite-like  
750 structured phase (Ahrens et al., 1969). A hollandite-like structure scenario would be supported  
751 by the existence of high-pressure phases and/or high-coordinated glasses. However, evidence of  
752 high-pressure phases was not observed in static experimentally produced diaplectic glass (Daniel  
753 et al., 1997; Sims et al., 2020). Likewise, high-coordinated glasses were not observed in in situ  
754 compression experiments or naturally shocked samples (Kubo et al., 2010; Jaret et al., 2015;

755 Sims et al., 2019). Furthermore, recent  $^{29}\text{Si}$  NMR studies of natural diaplectic plagioclase, did  
756 not detect structural remnants of high-pressure high-coordination glass, even though such  
757 remnants would be resolvable with NMR analyses (Jaret et al., 2015). Ashworth and Schneider  
758 (1985) proposed a mechanism for diaplectic  $\text{SiO}_2$  glass that is similar to metamictization – a  
759 process in which alpha particles physically displace atoms within the unit cell. In a diaplectic  
760 glass, physical displacement of atoms would be caused by the shock wave, rather than alpha  
761 particles.

762 Shock recovery experiments at starting temperatures of 77-293 K showed that in this  
763 temperature range the formation of diaplectic plagioclase glass is controlled by shock pressure  
764 only, and not temperature (Fritz et al., 2019a). This observation, together with the structural  
765 properties of diaplectic plagioclase glass (Jaret et al., 2015) and the recognition that diaplectic  
766 glass forms via a gradual collapse of the lattice (Hörz and Quaide, 1973, Fritz et al., 2019a), and  
767 the observation of isotropic twins alternating with crystalline twins in plagioclase (e.g., Stöffler,  
768 1971; Pickersgill et al., 2015a,b; Pittarello et al., 2020a) further advocates for a solid-state  
769 amorphization mechanism.

## 770 **6.2. Maskelynite and diaplectic glass**

771 The first description of maskelynite was by Tschermak (1872) in the Shergotty martian  
772 meteorite, described as a previously unrecognised isotropic phase of near labradorite  
773 composition. This new phase was then named after M. H. N. Story-Maskelyne, a famous English  
774 mineralogist of the time. Tschermak (1883) then found the same phase in chondrites and, upon  
775 realizing that it was pseudomorphous with plagioclase, changed his interpretation to a melted or  
776 otherwise vitrified glass of plagioclase composition. A shock origin for maskelynite was  
777 suggested by Binns (1967) based on differences in the RI of maskelynite and melt-glass. At the

778 same time, the term *diaplectic glass* was proposed to refer to “amorphous phases produced by  
779 shock waves without melting, and [which] are distinguishable from ordinary molten glasses”  
780 (Engelhardt et al., 1967; and Engelhardt and Stöffler, 1968). Since that time, *maskelynite* and  
781 *diaplectic plagioclase glass* have often been used interchangeably, which attached a genetic  
782 connotation to maskelynite. Historically, most isotropic plagioclase-composition material was  
783 called maskelynite, but modern observations have enabled us to distinguish between phases that  
784 have flowed (e.g., Chen and El Goresy, 2000) and “true” diaplectic glasses that show no  
785 evidence of melting (e.g., Jaret et al., 2015, Diemann and Arndt, 1984), even within a single thin  
786 section. Because there is confusion over how the term maskelynite is interpreted, it is imperative  
787 that clear observations regarding the nature of the isotropic phase be included in any description  
788 so that it is clear whether the authors are using “maskelynite” to refer to an apparently solid-state  
789 glass or a melt product.

## 790 **7. Concluding remarks & remaining questions**

791 Feldspars are typically difficult to study due to the optical complexity of their crystal  
792 structure and, when in the presence of water, their relatively rapid weathering rate.  
793 Consequently, the feldspar group is often neglected in favor of quartz or olivine for use in shock  
794 barometry. As a result, the shock scale for feldspar is limited and, essentially, qualitative.  
795 However, feldspars will have particular utility when studying quartz-poor rocks such as basalt, a  
796 dominant rock type on Mars and Earth’s moon, and an important surficial lithology on most  
797 terrestrial bodies. There have been many advances on the effects of shock metamorphism on  
798 feldspars over the past five decades but there are several major questions remain open, in  
799 particular:

- 800 1. Determining the exact nature and formation mechanism of diaplectic glass. Recent work  
801 has made significant strides, but confirmation and further studies would help to bolster  
802 these results (Section 6.1);
- 803 2. Understanding the formation of PDFs in feldspars, including exactly why the An content  
804 appears to be a controlling factor; the relationship between thin closely spaced shock-  
805 induced microtwins and PDFs; whether FFs form in feldspars; and the development of  
806 apparently shock-induced planar features that do not appear to be strict PDFs when  
807 investigated using high resolution TEM and EBSD (Section 3.4);
- 808 3. Whether or not measuring the crystallographic orientations of shock-induced planar  
809 features in feldspars can be improved upon and developed into a useful tool for shock  
810 barometry (Section 5.1);
- 811 4. Understanding the exact mechanism by which only alternate twins deform in  
812 polysynthetically twinned plagioclase, leaving the other set apparently undamaged, and  
813 whether this could be used to determine the local orientation of shock wave propagation  
814 (Section 3.6);
- 815 5. How, or if, the intensity of shock effects is affected by the presence of pre-existing  
816 microstructures (e.g., cleavage, twinning, exsolution), the crystallographic orientation  
817 with respect to the shock wave, and the presence of other mineral phases (Sections 3.4,  
818 3.6);
- 819 6. How best to utilize and integrate multiple analytical techniques, which probe slightly  
820 different aspects of feldspar composition/structure and deformation (Section 5);

821 7. How to reconcile the pressure conditions associated with crystallographic transformations  
822 between various experimental techniques and naturally shocked materials, which is an  
823 ongoing challenge in many fields of geoscience: do all types of minerals behave in the  
824 same way during shock recovery experiments and natural impacts or do some mineral  
825 develop the same shock effects at lower pressures? (Section 4)

826 8. How does shock influence radioisotopic age determination? A topic not touched upon in  
827 this paper, because it is worth an entire volume on its own, but the interested reader can  
828 examine, for example, Jessberger and Ostertag (1982) and Fernandes et al. (2009).

829 Because shocked feldspars are a significant constituent of most planetary materials, the  
830 importance of understanding their shock behavior and ability to inform shock barometry will  
831 only become more relevant as we gain increasing numbers of samples from other planetary  
832 bodies.

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1440

## 1441 **10. FIGURE CAPTIONS**

1442 Figure 1: Plot showing the relationship between anorthite content and pressure needed to  
1443 form partly to fully isotropic plagioclase. The overall trend demonstrates that plagioclase of  
1444 higher mole % An content results in complete amorphization at lower pressures than plagioclase  
1445 of lower mole % An content. The graph is based on shock recovery experimental data  
1446 summarized by Fritz et al. (2019a).

1447

1448 Figure 2: Shock pressure versus specific volume Hugoniot plot of a representative  
1449 plagioclase ( $An_{-50}$ ) illustrating the three distinct phase regimes (I, low pressure; II, mixed; and

1450 III, high pressure) and the range of the Hugoniot elastic limit (HEL). Data from Marsh (1980)  
1451 and van Thiel (1977).

1452

1453 Figure 3: A) Transmitted light photomicrograph of heavily fractured anorthite in Apollo  
1454 sample 15415,90; fractures are highlighted by offset twins when viewed between crossed  
1455 polarizers. B) Backscattered electron image of fractured plagioclase (albite, dark grey) from the  
1456 El'gygytgyn impact structure, with planar features (black arrows) in the lower right corner.

1457

1458 Figure 4: Undulatory extinction in Apollo samples 76335,55 (A) and 60215,13 (B)  
1459 between cross-polarized light, as the stage is rotated a wave of extinction passes through the  
1460 crystal.

1461

1462 Figure 5: Planar microstructures in oligoclase from the Tenoumer impact structure (A,  
1463 B); note the planar features in (B) only exist in alternating twins; after Jaret et al. (2014). C)  
1464 Andesine ( $An_{23}$ ) from the Manicouagan impact structure showing multiple sets of planar  
1465 features. D) K-rich feldspar with albite exsolution (beige, labeled Ab) and tartan twins spread  
1466 over the whole grain from the Gardnos impact structure; planar features on the right side of the  
1467 image (parallel to white lines) with the same orientation as polysynthetic twins.

1468

1469 Figure 6: Plane-polarised photomicrographs of andesine experimentally shocked to 25  
1470 GPa (A) and 56 GPa (B) showing progressive degrees of shock-darkening. Electron probe  
1471 microanalysis, Raman, and FTIR analyses suggest this darkening is purely optical, and does not  
1472 correspond with observable chemical or structural changes. After Jaret et al. (2018).

1473

1474           Figure 7: Diaplectic plagioclase glass. A) Illustration of maskelynite in the Shergotty  
1475 martian meteorite, highlighting preserved external crystal morphology and irregular fractures  
1476 extending from maskelynite grain margins into surrounding igneous pyroxene (after Tschermak  
1477 1872, drawn at 75x magnification, estimated grain size is  $\sim 400 \mu\text{m}$ ). B) Composite reflected light  
1478 (left) and transmitted light between crossed polarizers (right) image showing a partly isotropic  
1479 plagioclase crystal enclosed by olivine in martian meteorite Chassigny, though the plagioclase is  
1480 only partly isotropic, the same fracture pattern is observed as in the illustration by Tschermak  
1481 (A). C,D) Partially isotropic plagioclase in PPL (C) and XPL (D) from Apollo 15 sample  
1482 15684,4. Black areas in XPL remain extinct on rotation of the stage. E,F) Completely isotropic  
1483 plagioclase in basalt from Lonar crater, India (after Jaret et al. 2015). G) Plane-polarised  
1484 transmitted light image of maskelynite from the Shergotty meteorite ( $\sim 30 \text{ GPa}$ ) which includes a  
1485 perfectly preserved magmatic pyroxene. H) Plane-polarised transmitted light image of  
1486 maskelynite from the Los Angeles Martian meteorite ( $\sim 45 \text{ GPa}$ ) with a pyroxene fragment  
1487 partially affected by eutectic melting.

1488

1489           Figure 8: Plots demonstrating the relationship between refractive index, pressure, and  
1490 composition of diaplectic feldspar glasses. A) Refractive index vs. An content. Displayed are  
1491 isobars indicating the shock pressures that result in a plagioclase normative phase with a given  
1492 index of refraction. Fields indicating the optical properties, birefringent or isotropic, are  
1493 indicated by grey lines; after Fritz et al. (2005b) using data from Stöffler et al. (1986). B)  
1494 Refractive indices of diaplectic feldspar glasses from experiments in the pressure range from 28-  
1495 45 GPa. After Ostertag (1983).

1496

1497           Figure 9: IR spectra of 2 orientations from (A) crystalline, (B) diaplectic glass, and (C)  
1498 melt glass of labradorite composition. Crystalline labradorite exhibits multiple peaks, all  
1499 showing shifts with orientation indicating the anisotropic nature of the material. Both melt glass  
1500 and diaplectic glass are amorphous (one broad peak) but diaplectic glass retains orientation  
1501 effects not observed in the melt glass sample. After Jaret et al. (2015).

1502

1503           Figure 10: Plane-polarised light photomicrograph of plagioclase glass displaying bubbles  
1504 and flow textures surrounded by brown stained olivine, in some areas the plagioclase glass has  
1505 developed a rim of recrystallized plagioclase. From the strongly shocked ALH (Allan Hills)  
1506 77005 martian meteorite.

1507

1508           Figure 11: Photomicrographs of sieve or checkerboard feldspar. A, B) Plagioclase from  
1509 Gow Lake impact structure, showing preserved polysynthetic twins between crossed polarizers  
1510 (B). C) Plane-polarized light photomicrographs showing magnified view of checkerboard  
1511 feldspar in a clast from Gow Lake impact melt rock; D) sieve-textured feldspar from a volcanic  
1512 andesite (image credit: Dr Alessandro Mommio).

1513

1514           Figure 12: Cathodoluminescence images and spectra of plagioclase ( $An_{55}$ ) experimentally  
1515 shocked to 28 GPa by Fritz et al., (2019a). A) Transmitted light photomicrograph between  
1516 crossed polarizers (XPL) showing heterogeneous distribution of isotropic domains (black). B)  
1517 Optical microscope cathodoluminescence (OM-CL) image of the same grain as in (A), showing  
1518 three spectra locations (1-3). Areas of higher crystallinity (point 1) have higher luminescence

1519 than isotropic domains (points 2 and 3). C) CL spectra corresponding to points 1-3 in (B), further  
1520 showing that the domains with remaining crystallinity have higher luminescence than those that  
1521 are isotropic, after Pittarello et al. (2020b).

1522

1523 Figure 13: Raman spectra of shocked feldspar. A) Raman spectra of andesine ( $An_{43}$ ):  
1524 unshocked (0 GPa), shocked (up to 29 GPa), and diaplectic glass (>29 GPa). For clarity, three  
1525 spectra of 0 GPa (top line), 26 GPa (middle line), and 56 GPa (bottom line), representing the  
1526 aforementioned categories (unshocked, shocked, diaplectic glass) are shown here. Up to  
1527 pressures of 29 GPa (middle line), increasing shock corresponds to an overall decrease in peak  
1528 intensity and causes the  $480/510 \Delta \text{cm}^{-1}$  peak ratio to approach 1. Above 29 GPa (bottom line),  
1529 the sample appears amorphous, and exhibits a different Raman pattern, with two broad peaks at  
1530  $\sim 500 \Delta \text{cm}^{-1}$  and  $580 \Delta \text{cm}^{-1}$ . Data from Jaret et al. (2018). B) Raman spectra of diaplectic  
1531 feldspar glass with three different An contents (top:  $An_{15}$  – naturally shocked L6 chondrite;  
1532 middle:  $An_{50}$  – experimentally shocked troctolite; bottom:  $An_{94}$  – experimentally shocked  
1533 gabbro). The spectral position of the hump in the region of  $950$  to  $1150 \text{ cm}^{-1}$  varies with An  
1534 content. Data from Fritz et al. (2019a).

1535

Table 1: Summary table of impact-related metamorphism in feldspar group minerals, with key to the section of the paper that discusses them in more detail (Column 1: “Sctn.”) and relevant figures (Fig.). Reported pressures are in GPa and are derived from <sup>[1]</sup>Ostertag (1983) and <sup>[2]</sup>Jaret et al. (2018), both of which conducted shock recovery (gas gun) experiments. These pressures should act as guides and are indicative of trends in the order of formation of shock features, however the exact values may not be directly applicable to natural samples for reasons discussed in Section 4.

Sctn.	Feature	Fig.	Orthoclase	Sanidine	Microcline	Albite	Oligoclase	Andesine	Labradorite	Bytownite		
			Single crystal <sup>[1]</sup>	Single crystal <sup>[1]</sup>	Single crystal <sup>[1]</sup>	An <sub>02</sub> , rock <sup>[2]</sup>	Single crystal <sup>[1]</sup>	An <sub>43</sub> , rock <sup>[2]</sup>	Single crystal <sup>[1]</sup>	An <sub>77</sub> <sup>b</sup> , rock <sup>[1]</sup> , [2]		
3.1	Irregular fracturing <sup>a</sup>	3	<10.5 to 32	<10.5 to 30	<10.5 to 34	<i>N.R.</i>	<10.5 to 29	<i>N.R.</i>	10.5 to 27	10.5 to 26 <sup>a</sup>		
3.2	Undulatory extinction Wave of extinction passes through the grain	4	<i>No limits explicitly stated</i>									
3.3	Mosaicism Patchwork of slightly different extinction angles within a grain	<i>b</i>	10.5 to 28	14 to 30	14 to 45	<i>Ambiguous</i>	~18 to 31	<i>Ambiguous</i>	18 to 28	22 to 27 <sup>[1]</sup>		
3.4	Planar features/microstructures Open parallel planar fractures (PFs) and/or Closely spaced parallel planar lamellae (PDFs)	5	~10.5 to 36	10.5 to 31	~10.5 to 42	<i>N.O.</i>	~14 to 33	<i>N.O.</i>	14 to 23	14 to 27 <sup>[1]</sup> <i>N.O.</i> <sup>[2]</sup>		
3.5	Shock darkening Darkening of mineral in plane-polarised transmitted light	6	<i>N.R.</i>	<i>N.R.</i>	<i>N.R.</i>	>~24	<i>N.R.</i>	>~24	<i>N.R.</i>	>~22 <sup>[1]</sup>		
3.7	Diaplectic glass Amorphous phase that maintains external crystal form and internal features such as chemical zoning	7		Partial		>26-28	>26-28	>26-28	~26-28	~28-30	>26-28	~25-27 <sup>[2]</sup>
				Complete		>32-34	>32-34	>45 <sup>c</sup>	>~55 <sup>d</sup>	>32-34	>~47 <sup>d</sup>	>28-30 <sup>[1]</sup> , >~38 <sup>[2]</sup> <sub>d</sub>
3.8	Melt glass Amorphous phase that shows flow textures and/or vesiculation	10	45-50 <sup>e</sup>	45-50 <sup>e</sup>	45-50 <sup>e</sup>	<i>N.O.</i>	>42 <sup>e</sup>	<i>N.O.</i>	>42 <sup>e</sup>	>45-50 <sup>[1]</sup> <sub>e</sub> , <i>N.O.</i> <sup>[2]</sup>		
3.9	High pressure phases Phases related to feldspar composition that are stable at higher pressure conditions than found on Earth’s surface		<i>See Table 2</i>									
3.10	Sieve and checkerboard feldspar Feldspar clasts in melt rocks made up of individual rectangular domains separated by melt products	11	<i>Temperature rather than pressure is the controlling factor</i>									
3.6	Alternate twin deformation Special cases of planar features or diaplectic glass confined to only alternate twins	5a,b	<i>This is a broad category of effects that each seem to occur over a narrow range of pressure conditions, somewhat lower than the equivalent full grain effect. e.g., alternate diaplectic glass twins fall into the category above of “partial isotropization”.</i>									

**N. R.** = Not reported – features were not explicitly discussed in the manuscript so presence in those samples is ambiguous. **N.O.** = None observed - explicitly stated that the author looked but did not find that feature in the sample. **Ambiguous** = Jaret et al. (2018) found a texture reminiscent of mosaicism, but were ambiguous in their interpretation, so did not explicitly report pressure conditions for mosaicism in those samples. <sup>a</sup>Ostertag (1983) found that fracturing decreased over 22 GPa. <sup>b</sup>Note: no images of mosaicism are included because it is challenging to accurately capture that feature in a still image, and no high quality images were found. <sup>c</sup>Ostertag (1983) found that microcline maintained reduced birefringence up to 45 GPa, where their experiments stopped. <sup>d</sup>Note: Different pressures for the amorphisation of plagioclase are reported in the literature, the table shows data from Ostertag (1983) and Jaret et al. (2018) and Figure 1 uses the data compiled by Fritz et al. (2019a), as a result the values reported in this table for complete amorphisation from Jaret et al. (2018) are higher (~55 GPa, ~47 GPa, ~38 GPa) than those shown in

Figure 1 (<35 GPa), likely due to differences in experimental setup and/or plagioclase composition. °Ostertag (1983) did not detect melt glass unambiguously, so assumed values to be in accordance with Stöffler and Hornemann (1972).

Table 2: High-pressure phases commonly associated with feldspathic compositions in highly shocked lithologies, including name, chemical composition, type locality, and the experimentally derived stability fields. The pressure [P] stability field of high-pressure phases strongly depends on chemical composition of the environment, and the prevailing temperatures [T], so the presented P-T conditions serve only as a guide and further details can be found in the cited references.

Name	Formula	Type Locality	Synthesized	Meteorites	Terrestrial impact structures	P-T	Refs.
Donwilhelmsite	CaAl <sub>4</sub> Si <sub>2</sub> O <sub>11</sub>	Lunar meteorite Oued Awlitis 001	Yes	Yes	No	13 – 36 GPa, >1200 °C	1-3
Jadeite	NaAlSi <sub>2</sub> O <sub>6</sub>	Myanmar		Yes	Yes	2.5 GPa; 1000 °C	4-9
Liebermannite (K-Hollandite)	KAlSi <sub>3</sub> O <sub>8</sub>	Martian meteorite Zagami	Yes			12 GPa, 1026 °C	10-16
Lingunite (Na-Hollandite)	NaAlSi <sub>3</sub> O <sub>8</sub>	Sixiangkou L6 chondrite	Yes	Yes	Yes	Stable at 22-23 GPa; >1200 °C Metastable at ~20-24 GPa; ~1100-1300 °C	11-15, 17-22
Tissintite	(Ca,Na,□)AlSi <sub>2</sub> O <sub>6</sub>	Martian meteorite Tissint	Yes			6 - 8.5 GPa; 1000-1350 °C	23, 24
Zagamiite	CaAl <sub>2</sub> Si <sub>3.5</sub> O <sub>11</sub>	Martian meteorite Zagami		Yes			23-26

[1] Fritz et al. (2019b), [2] Irifune et al. (1994), [3] Beck et al. (2004), [4] James (1969), [5] Ohtani et al. (2004), [6] Tomioka et al. (2007), [7] Knight and Price (2008), [8] Kubo et al. (2010), [9] Stähle et al. (2011), [10] Ringwood et al. (1967), [11] El Goresy et al. (2000), [12] Gillet et al. (2000), [13] Langenhorst and Poirier (2000), [14] Tomioka et al. (2000), [15] Langenhorst and Dressler (2003), [16] Ma et al. (2018), [17] Liu (1978), [18] Yagi et al. (1994), [19] Liu (2006), [20] Liu and El Goresy (2007), [21] Agarwal et al. (2016), [22] Kubo et al. (2017), [23] Ma et al. (2015), [24] Rucks et al. (2018), [25] Walton et al. (2014), [26] Ma et al. (2017).

Table 3: Summary of different experimental techniques for calibrating shock features and parameters.

Technique	Timescale	Method	Purpose	References
Shock	Nanoseconds	In situ measurements of shockwaves	Determination of the equation of state (EoS) using Rankine-Hugoniot relations (Fig. 2)	Wackerle (1962)
Shock recovery	Shock pulses of microsecond duration	Shock impedance or reverberation technique	Investigations of shock effects in materials	Müller and Hornemann (1969); Hörz and Quaide (1973); Fritz et al. (2011)
Rapid compression	Seconds – 10s mins	Membrane and dynamic diamond anvil cell	Determination of phase stability Typically collect simultaneous XRD during compression and decompression	Letoulliec et al. (1988); Evans et al. (2007); Sims et al. (2019)
Static compression	Minutes to hours	Diamond anvil cell	Determination of phase stability Typically coupled with either XRD or Raman spectroscopy after each compression step	Angel (1988); Williams and Jeanloz (1988); Daniel et al. (1997); Sims et al. (2020)

Table 4: Summary of analytical techniques used for assessing shock in feldspars. Optical microscopy is not included as it is assumed that primary characterization has already been conducted.

Technique	Sample requirements	General technique use	Limitations	Application to studies of shock metamorphism	Refs.
Universal-stage (U-stage)	Thin section	Measuring 3D orientation of planar and linear elements (i.e., twins, PDFs, optic axes)	Complex operation for biaxial solid solution minerals like feldspars (see Section 5.1) Only possible to view about 2/3 of a rectangular thin section U-stage does not fit on most modern microscopes	Indexing planar microstructures	(1-4)
Electron Backscattered Diffraction (EBSD)	Highly polished surface (i.e., thin or thick section, polished mount) Crystalline material Sample must fit in SEM	Analysing the distribution of orientations within or between minerals Measuring deformation within a crystal Identifying phases	Cannot uniquely index planar features without the addition of U-stage measurements or FIB cross-sectioning (see Section 5.1) Only applicable to crystalline material (no patterns provided by amorphous material)	Indexing planar microstructures (Section 5.1, complementary to other techniques)  Increasing internal misorientation with increasing shock pressure  No diffraction from diaplectic and melt glasses	(5-8)
X-ray diffraction (XRD)	Powder Single-crystal In situ $\mu$ xRD	Powder (whole rock or mineral separates) Individual mineral separate  Thin sections Polished mount Hand sample	Identifying phases Quantifying strain Refining crystal structure Identifying lattice breakdown (e.g., evaluating degree of shock)	Cannot distinguish amorphous phases from each other  Peaks broaden in chi direction (single-crystal or in situ $\mu$ XRD only) at relatively low pressure  Peaks broaden in the 2 $\theta$ direction at higher pressure  No diffraction from diaplectic and melt glasses (they produce a broad diffuse band)	(9-12)
High-energy total X-ray scattering	Powdered mineral (~mg) Single crystal	Quantifying atomic distances and positions (e.g., evaluating degree of ordering)	Requires synchrotron beamtime	Diaplectic glasses show intermediate order but no long-range order and orientation effects  Melt glasses show no ordering >10 Å and no orientation effects	(13-14)
Cathodoluminescence (CL)	Thin section Thick section Polished mount	Imaging microstructures and defects	Luminescence is controlled by multiple variables, such as the abundance of activators (trace elements, point defects, both), and the abundance of quenchers	Plagioclase: Normal band emission at ~550-570 nm commonly shifted to ~630 nm with increasing pressure Diaplectic glass results in a broad band ~ 350 nm Increasing amorphisation causes decreasing luminescence  Alkali-feldspar: Increasing pressure results in increasing luminescence	(15-31)
Raman spectroscopy	Thin section Thick section Polished mount	Identifying phases (i.e., high-pressure polymorphs; see Section 3.9) Evaluating crystallinity	Libraries and databases for mineral identification are limited Difficult to assign individual peaks to specific modes	Bands progressively broaden and decrease in intensity resulting in increased FWHM with increasing pressure  Diaplectic and melt glasses result in broad peaks about 500 and 1000 $\text{cm}^{-1}$ (dependent on composition, see Fig. 14)	(31-36)
Photoluminescence spectroscopy	Thin section Thick section Polished mount	Revealing lattice defects	Need to have endmembers for quantitative damage calibration (unshocked to completely amorphous)	$\text{Nd}^{3+}$ luminescence bands broaden with increasing amorphization	(27)

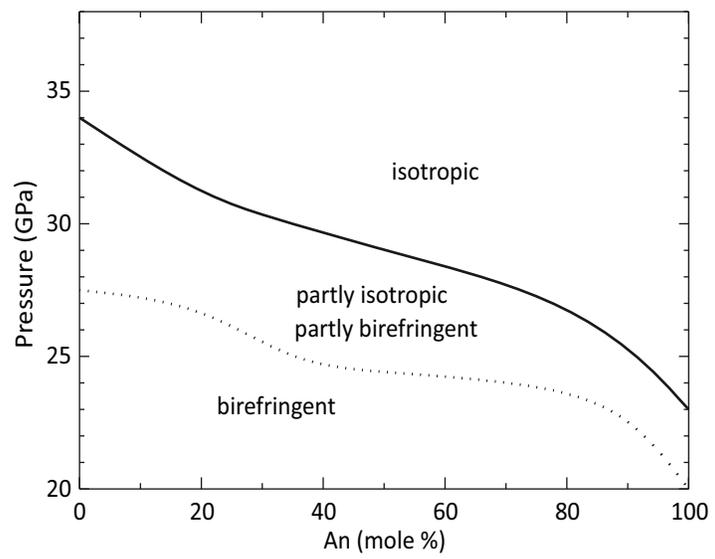
Thermal infrared absorption spectroscopy	Powder (whole rock or mineral separates) (mgs – 10 mgs)	Identifying phases Measuring crystal lattice structure	Averages grain orientations, so no orientation effects are detectable Limited sensitivity to minor components of the mixture	Decreased spectral detail and intensity with increasing pressure Loss of low wavenumber peaks (as a result of lattice disordering and increased glass content) with increasing pressure	(14, 35, 37–44)
Micro FTIR spectroscopy	Thin section Thick section Polished mount		Orientation sensitive measurements (e.g., peak height) can be hard to match to databases of non-oriented measurements	Diaplectic and melt glass result in one broad peak Decreased intensity of peaks due to loss of crystallinity with increasing pressure  Diaplectic glass preserves orientation effects Melt glass does not preserve orientation effects	(14, 35)
Nuclear magnetic resonance (NMR) spectroscopy	Powder (whole rock or mineral separates) (mgs – 10 mgs)	Identifying phases Determining local chemical environment	Sample must include an NMR active isotope (e.g., <sup>29</sup> Si in feldspars)	Diaplectic glass lacks long range order, results in a spectral shape similar to melt glass, but peak centre and FWHM are more similar to crystalline plagioclase  Melt glass lacks long range order, and results in broader spectral peak than crystalline plagioclase or diaplectic plagioclase glass	(14, 45)

[1] Reinhard (1931), [2] Turner (1947), [3] Stöffler (1967), [4] Dworak (1969), [5] Prior et al. (1999), [6] Pickersgill et al. (2017), [7] Pickersgill and Lee (2015), [8] Pittarello et al. (2020a), [9] Hörz and Quaide (1973), [10] Flemming (2007), [11] Vinet et al. (2011), [12] Pickersgill et al. (2015a), [13] Reeder and Michel (2013), [14] Jaret et al. (2015), [15] Lee et al. (2007), [16] Parsons et al. (2008), [17] Parsons and Lee (2009), [18] Finch and Klein (1999), [19] Götze et al. (2000), [20] Parsons et al. (2015), [21] Kaus and Bischoff (2000), [22] Götze (2009), [23] Gucsik et al. (2004), [24] Götze (2009), [25] Pittarello et al. (2015), [26] Kayama et al. (2018), [27] Pittarello et al. (2020b), [28] Kayama et al. (2012), [29] Kayama et al. (2009a), [30] Kayama et al. (2009b), [31] Kayama et al. (2009c), [32] Fritz et al. (2005a), [33] Fritz et al. (2019a), [34] Jaret et al. (2014), [35] Jaret et al. (2018), [36] Reynard et al. (1999), [37] Lyon (1963), [38] Bunch et al. (1967), [39] Stöffler and Hornemann (1972), [40] Stöffler (1974), [41] Arndt et al. (1982), [42] Ostertag (1983), [43] Johnson et al. (2002a), [44] Johnson (2003), [45] Myers et al. (1998).

Table 5: Summary table of orientations of shock-induced planar microstructures in feldspars measured by U-stage.

Feldspar composition	Reported orientation(s)	Crater	References
Labradorite	(001), (010), (111), (203), (101)	Manicouagan	Dworak (1969)
Andesine ( $\sim$ An <sub>30</sub> )	(001) 25%, (010) 11%, (100) 10%, ( $\bar{1}\bar{2}0$ ) 10%, (012) 7%, (130) 6%, others occurring at less than 2% frequency	Ries	Stöffler (1967)
Plagioclase	(001), (010), ( $\bar{1}02$ ), and ( $\bar{1}\bar{2}1$ )	Several Canadian craters	Robertson et al. (1968)
An <sub>98</sub>	(101), (010), (021)	Apollo breccia 10046	Dence et al. (1970)

# Figure 1



# Figure 2

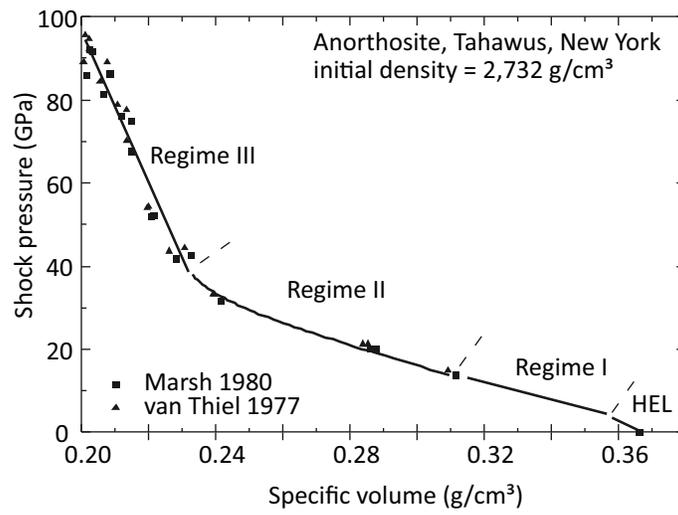
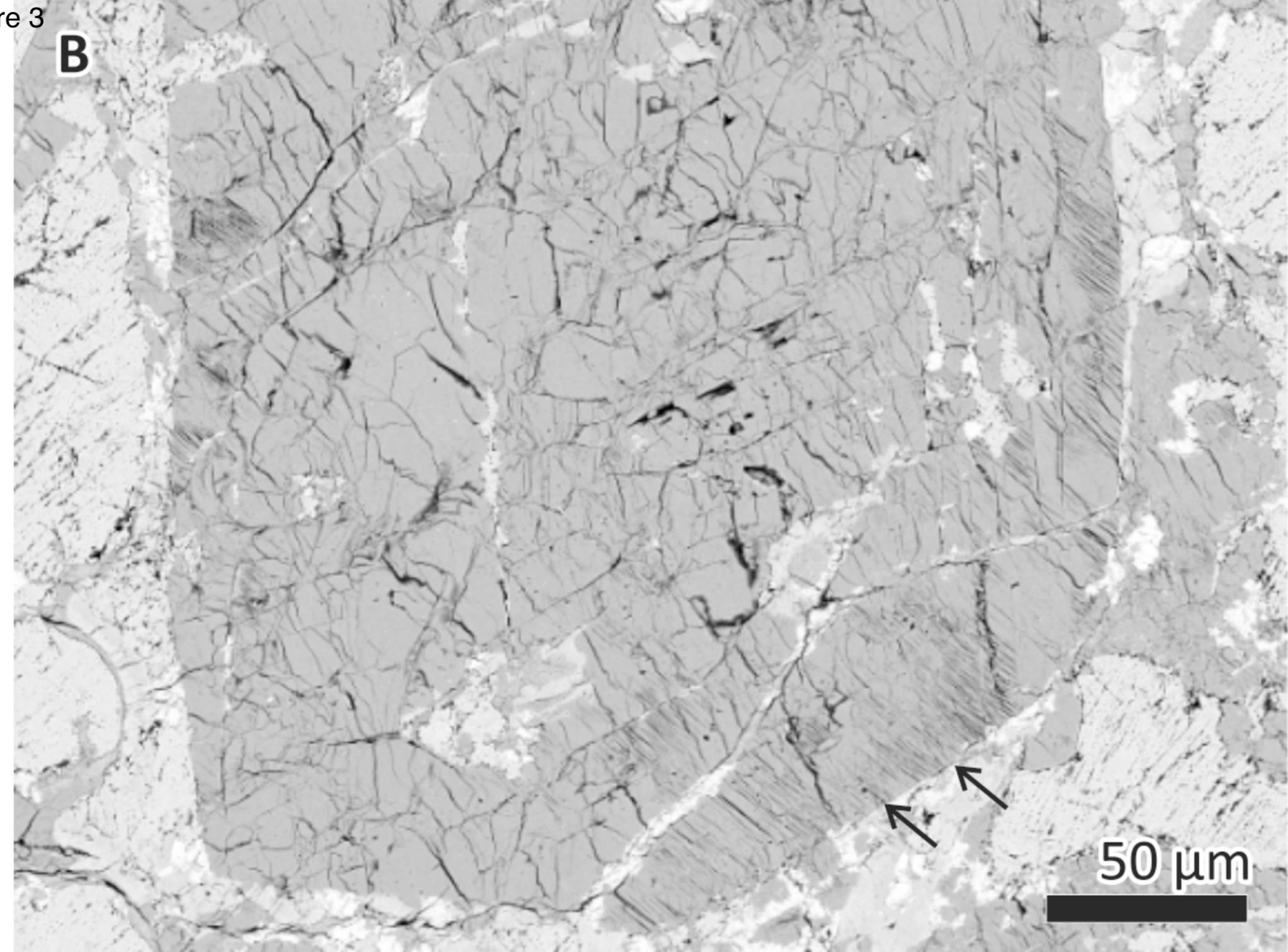
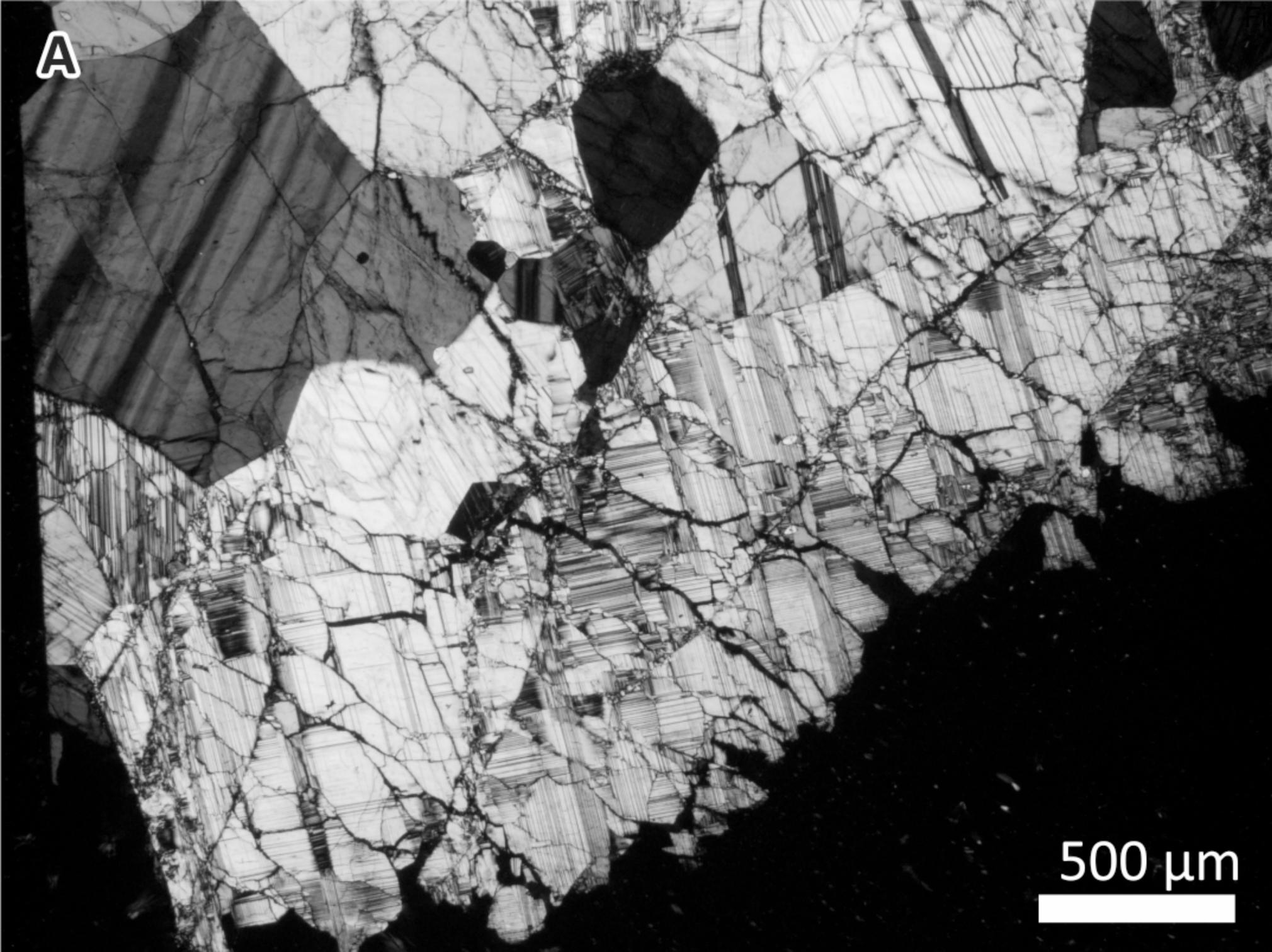


Figure 3



A

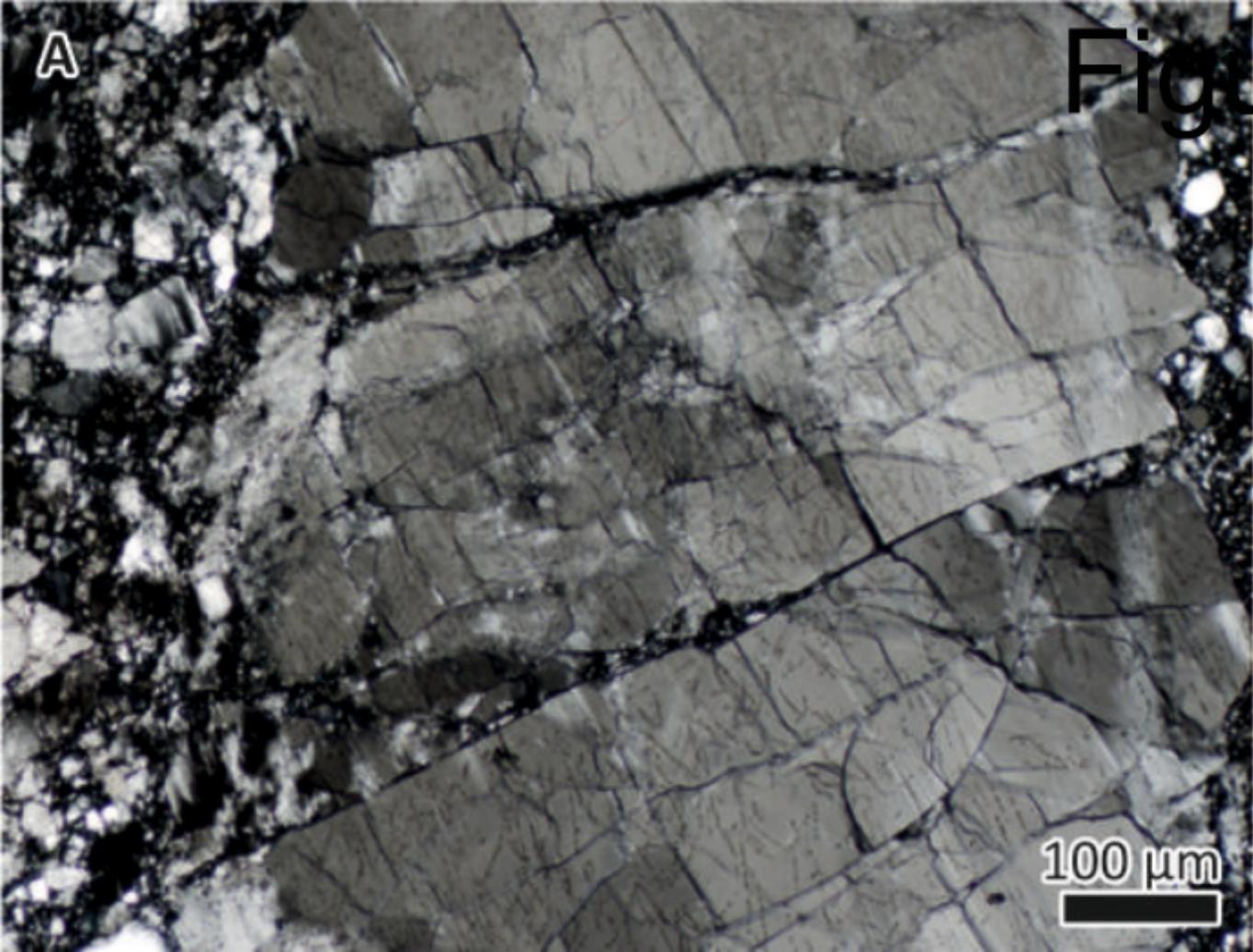


Figure 4

B

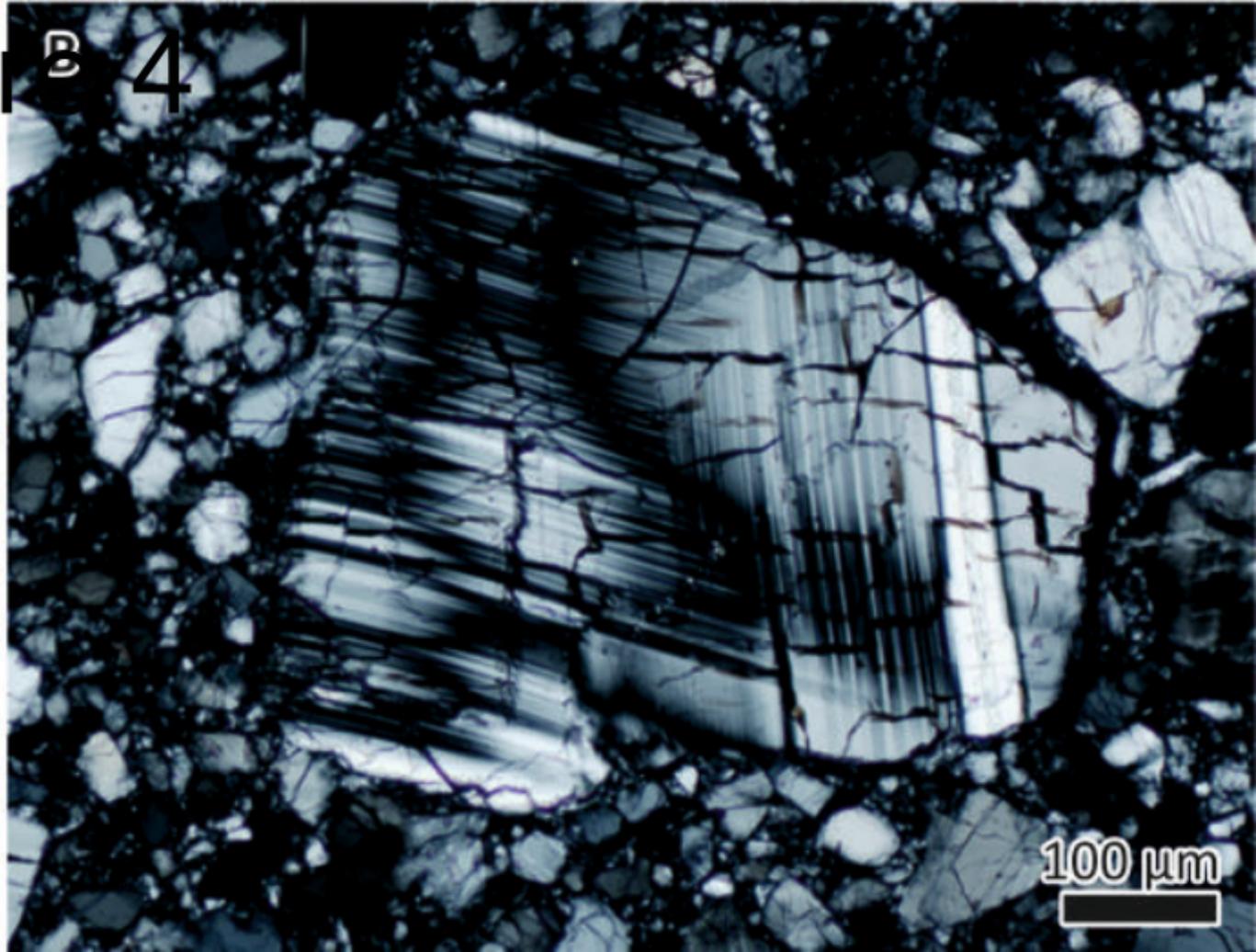


Figure 5

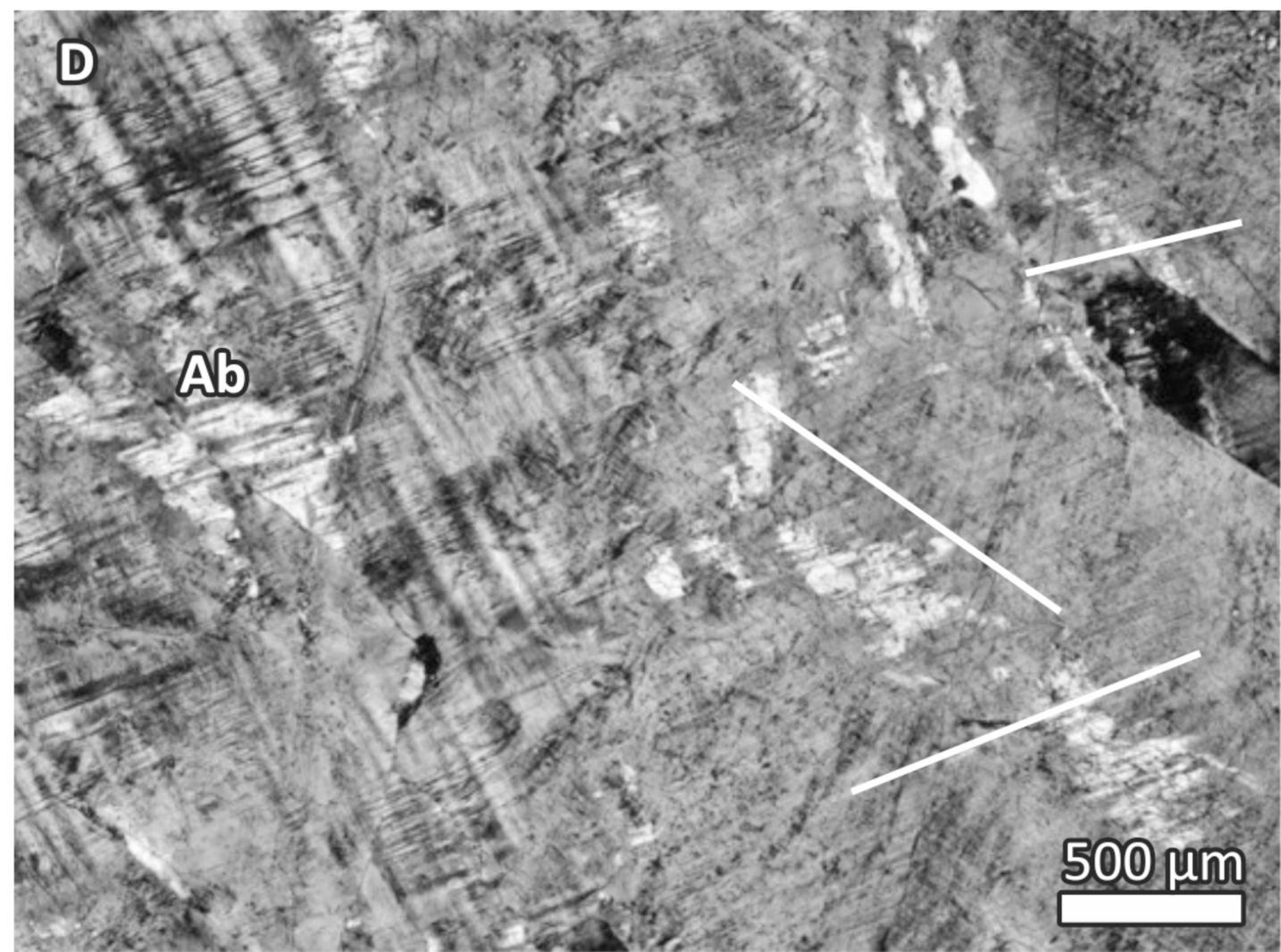
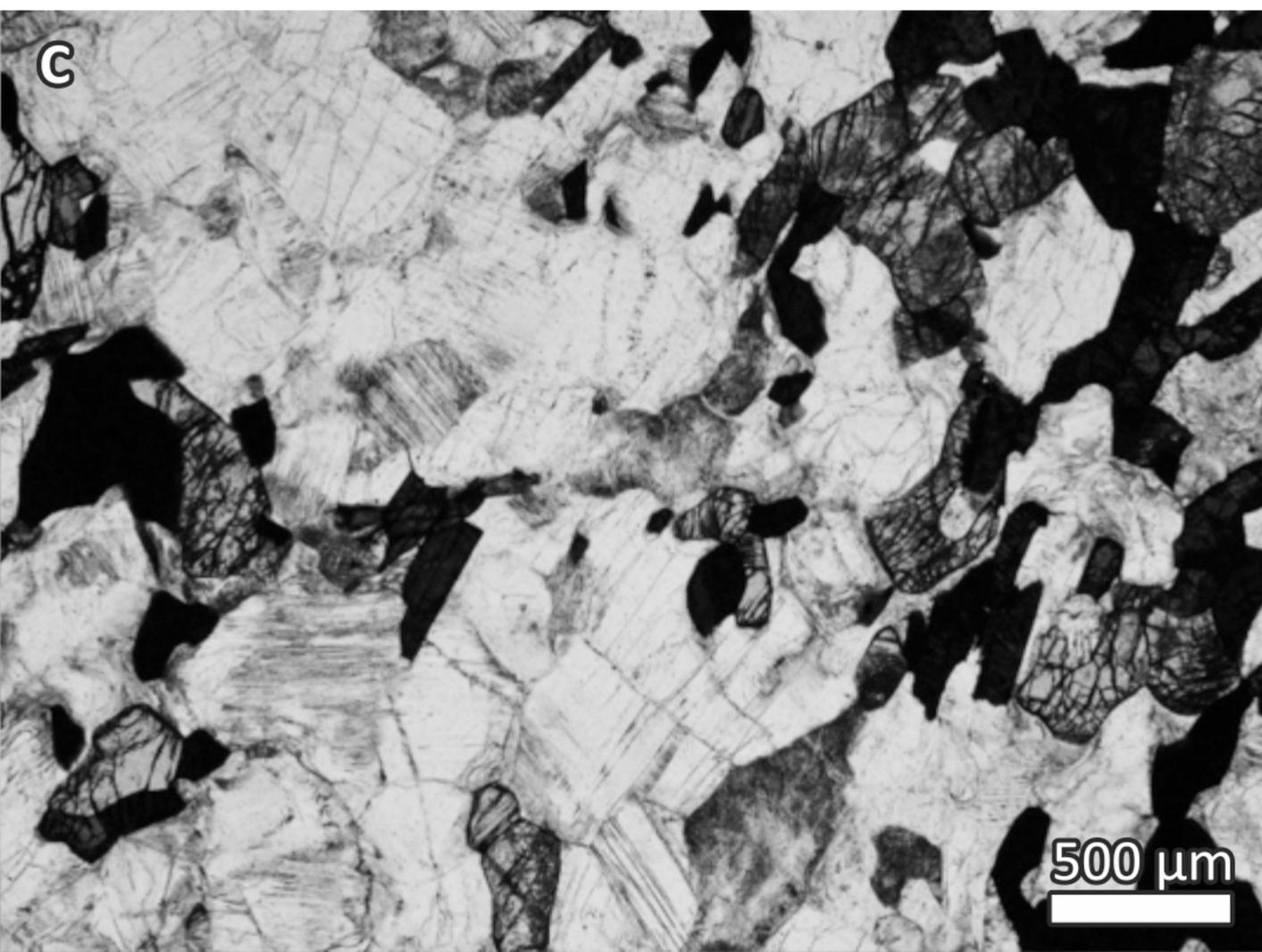
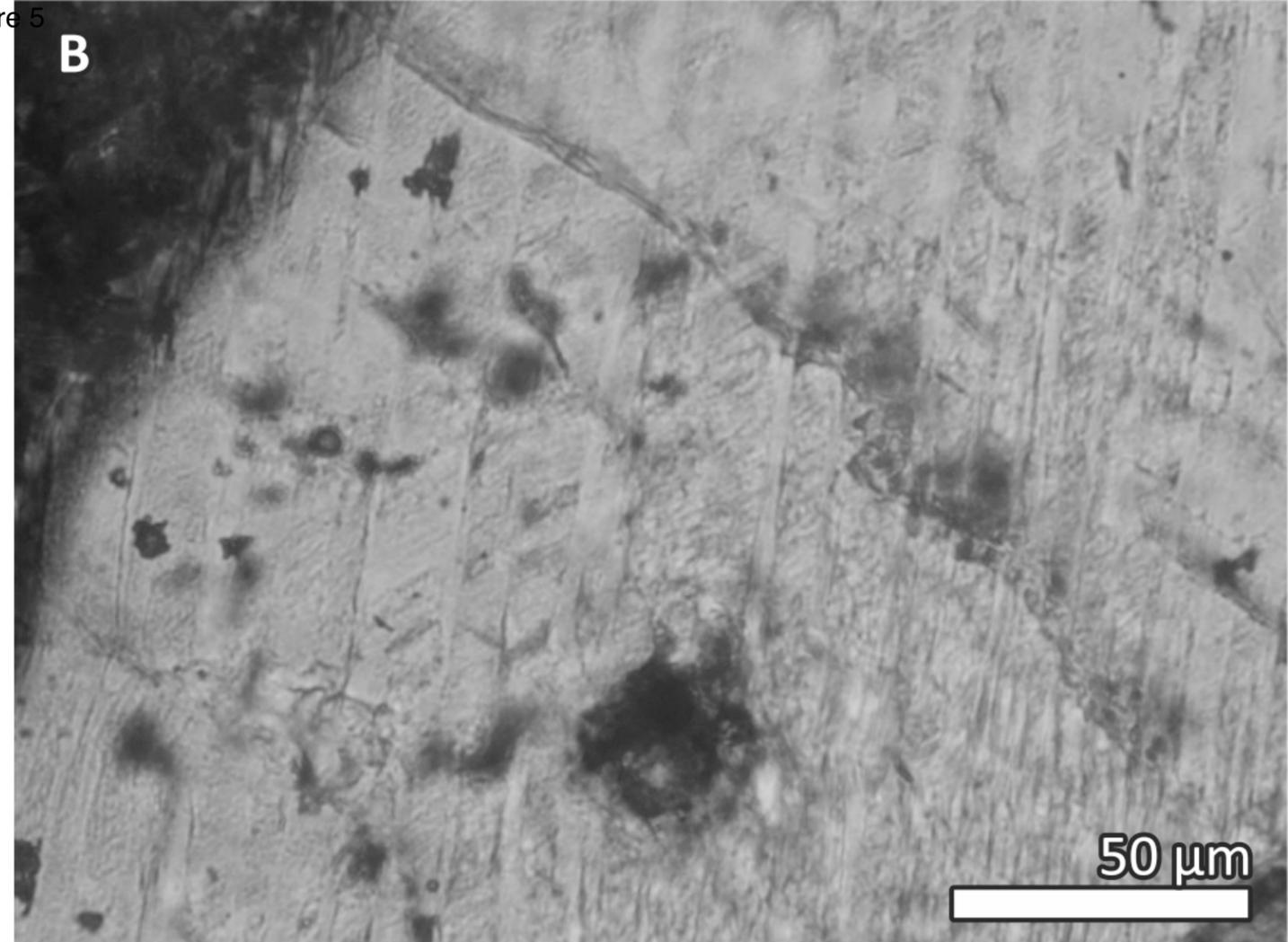
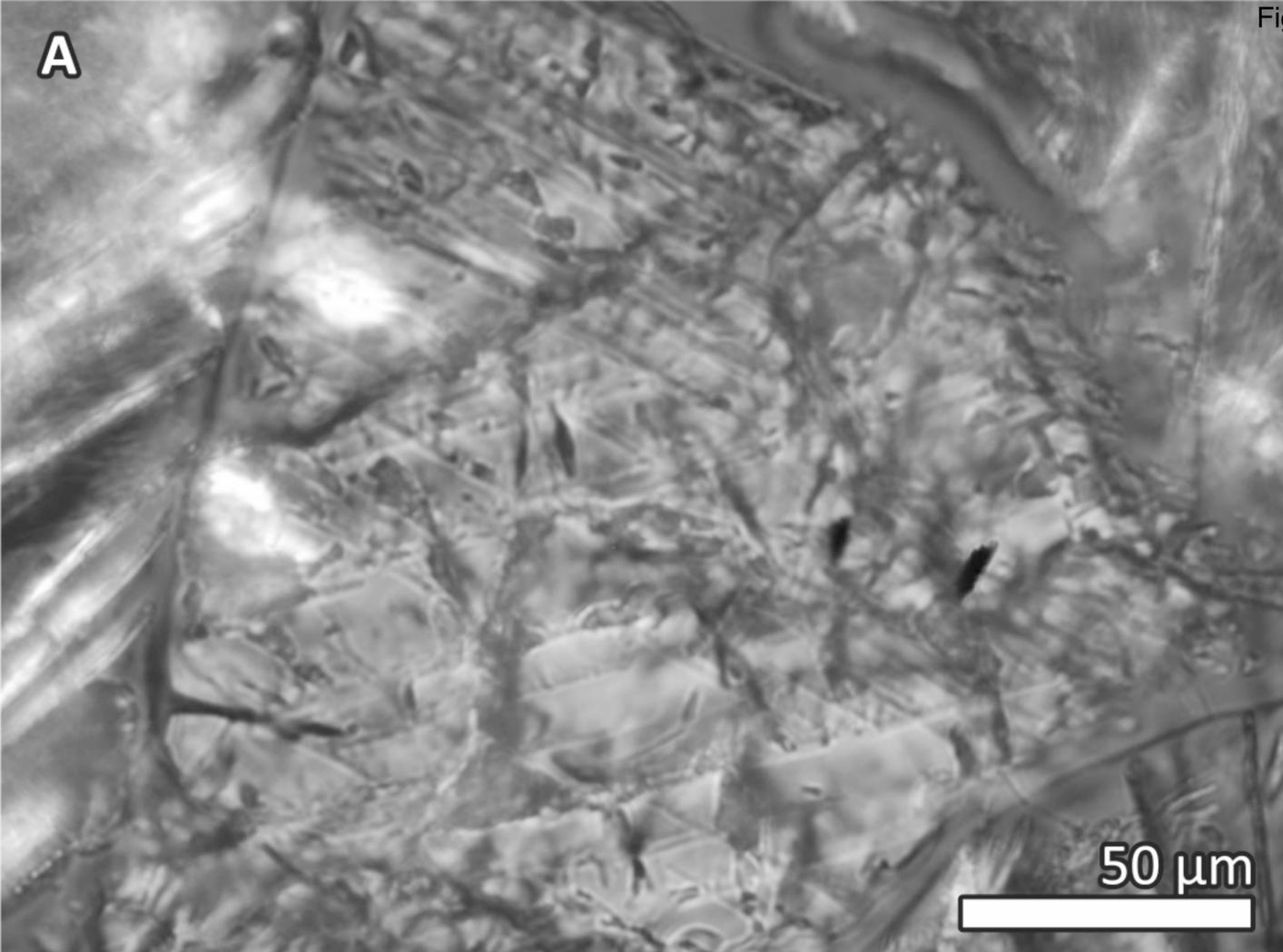


Figure 6

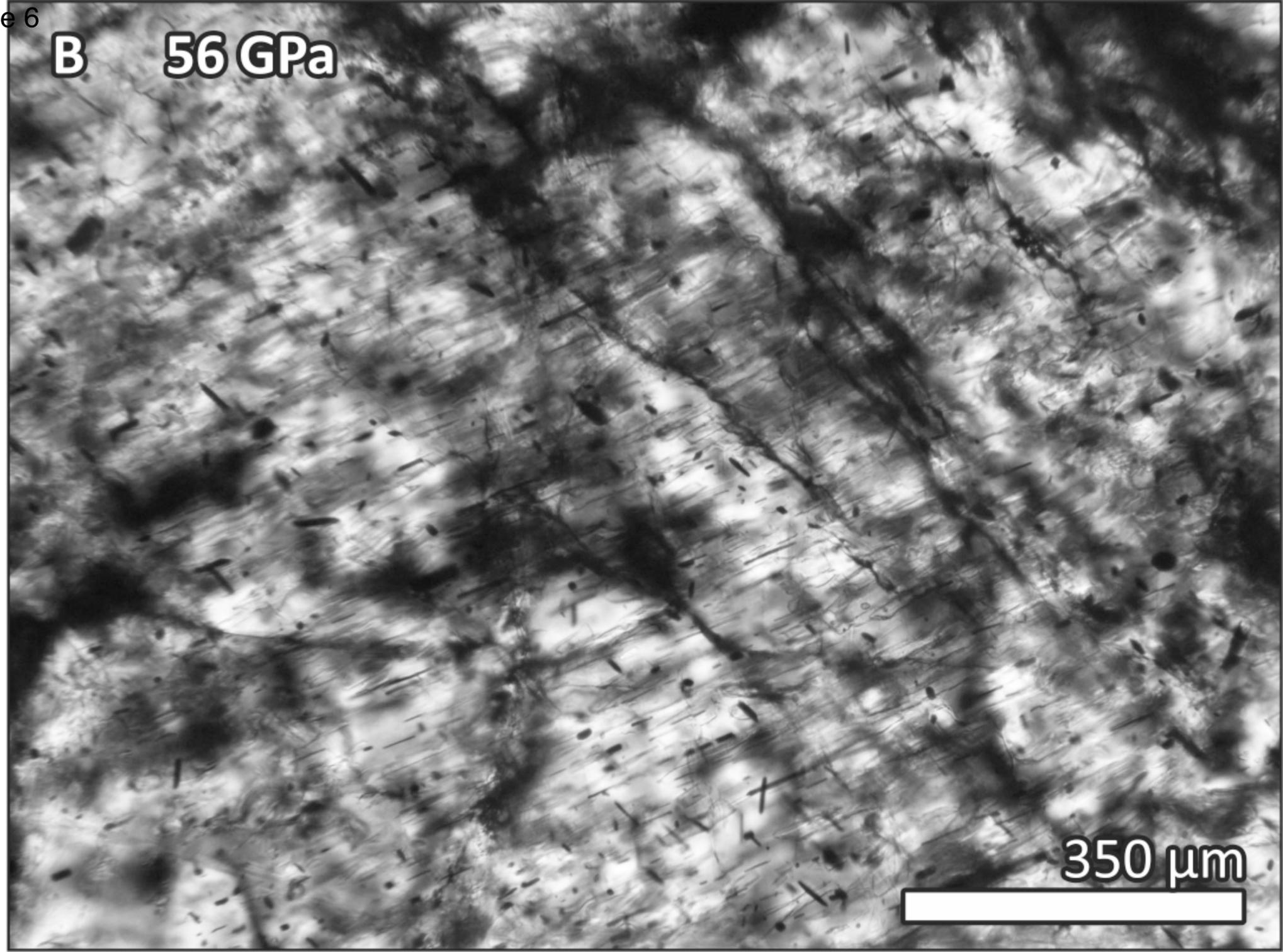
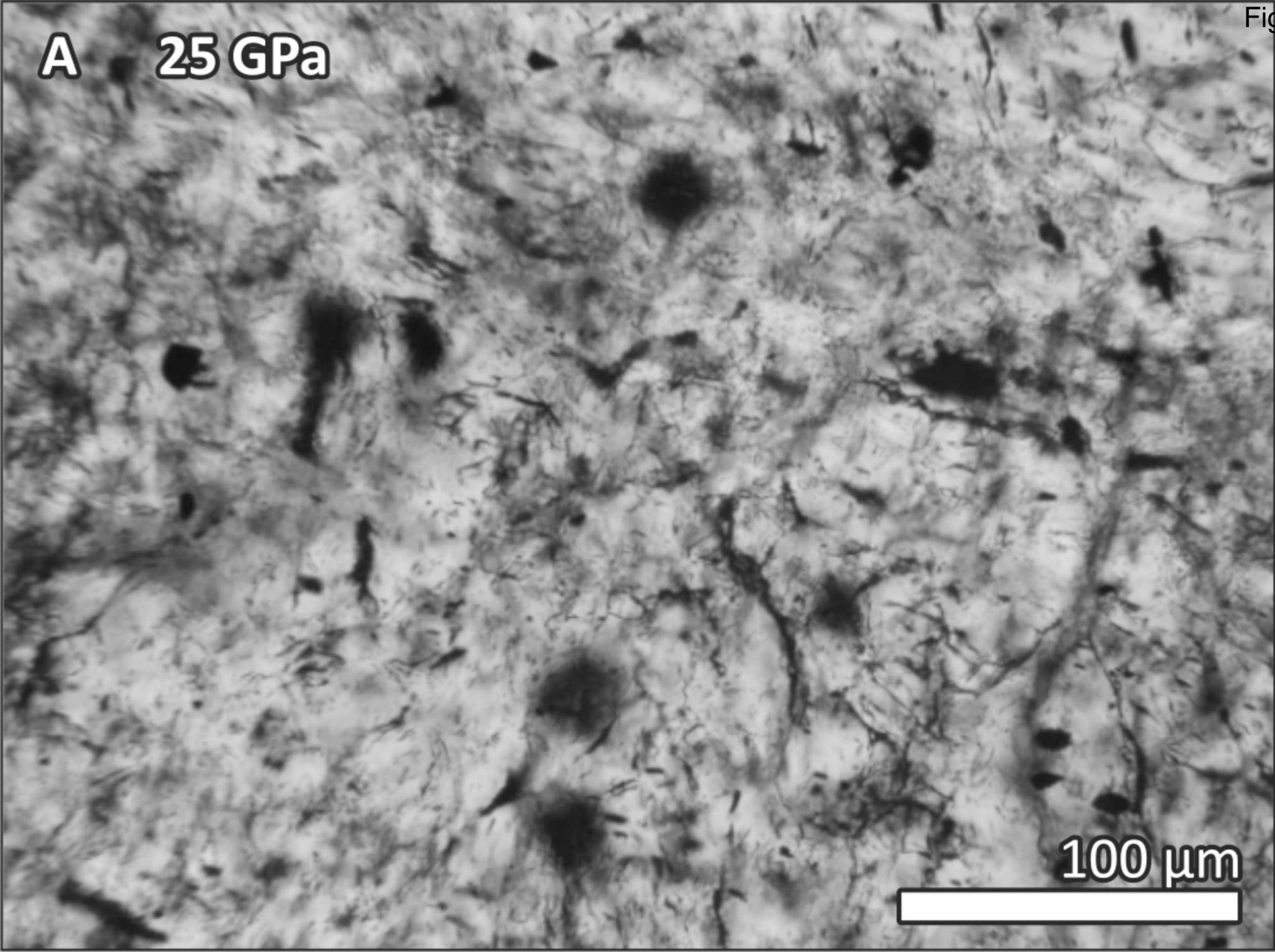
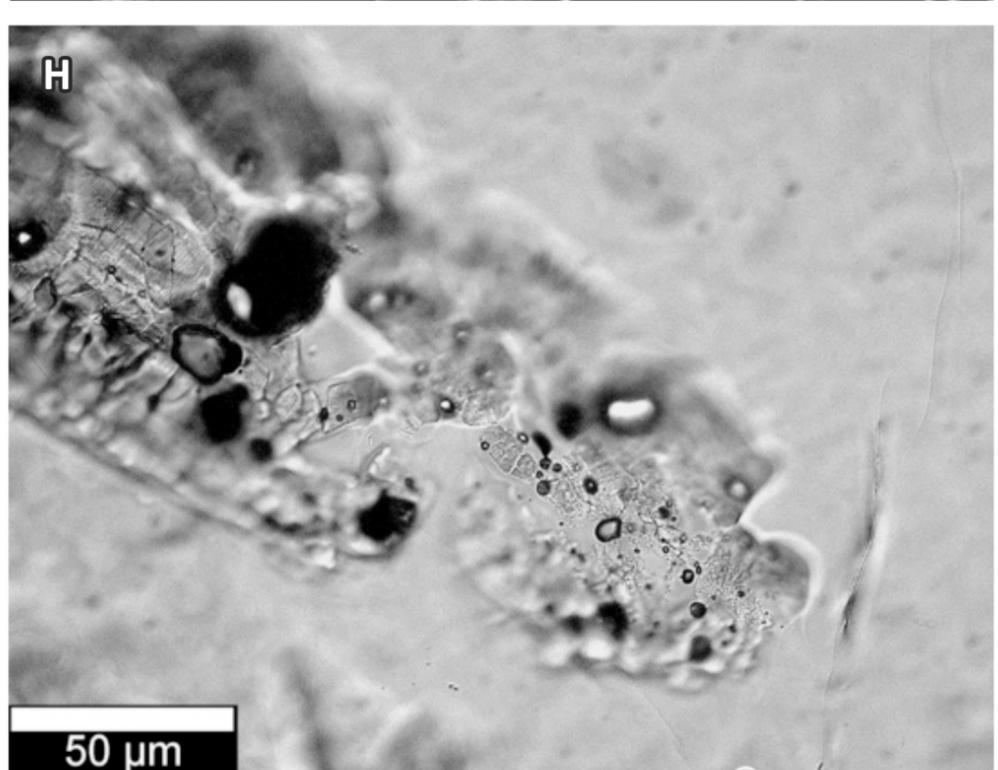
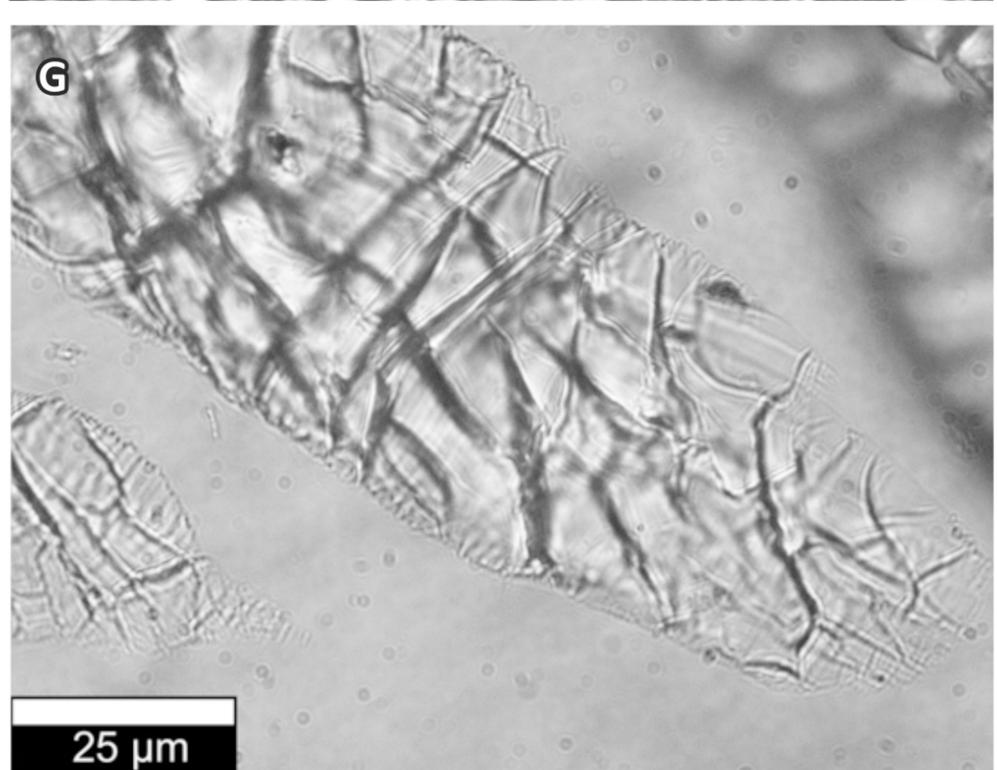
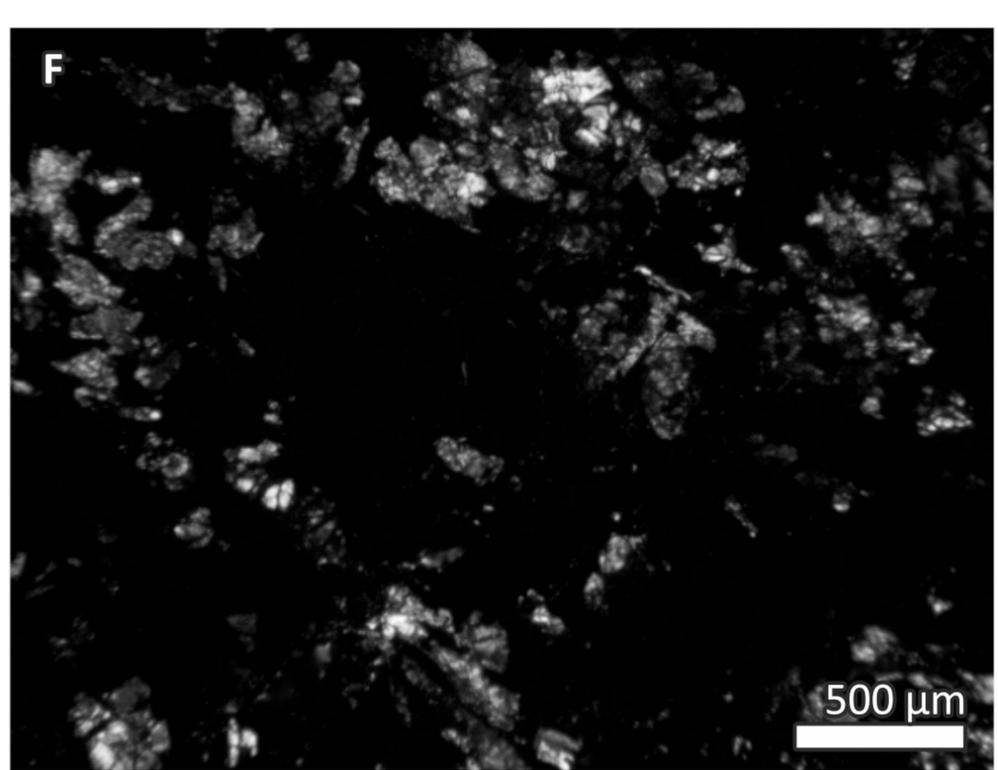
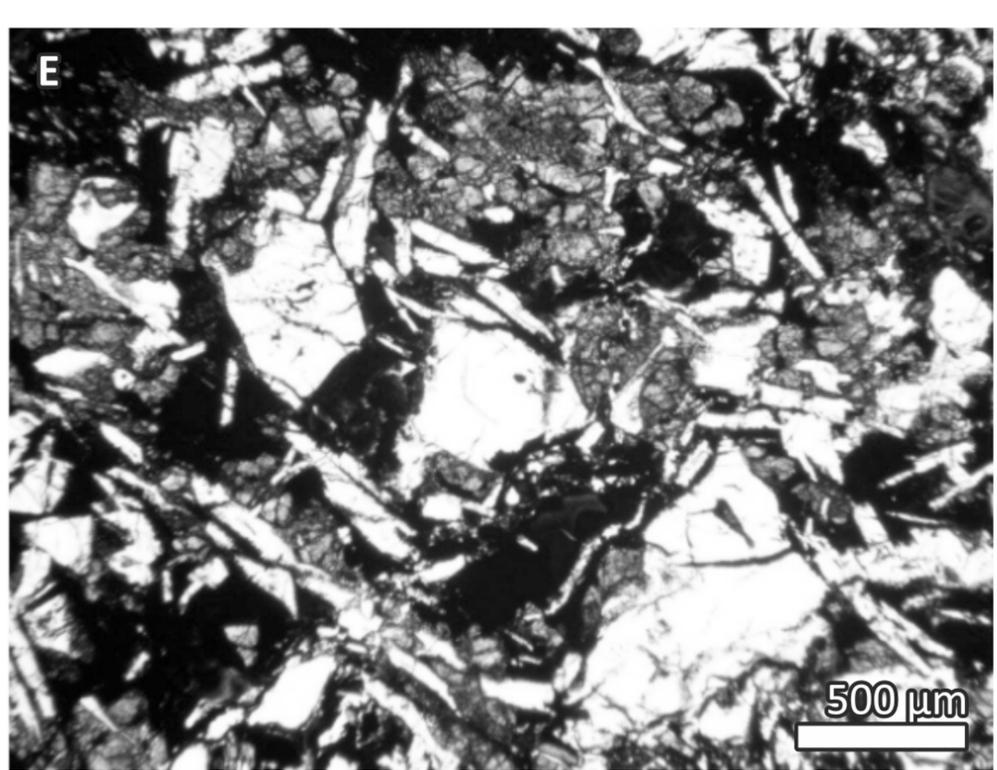
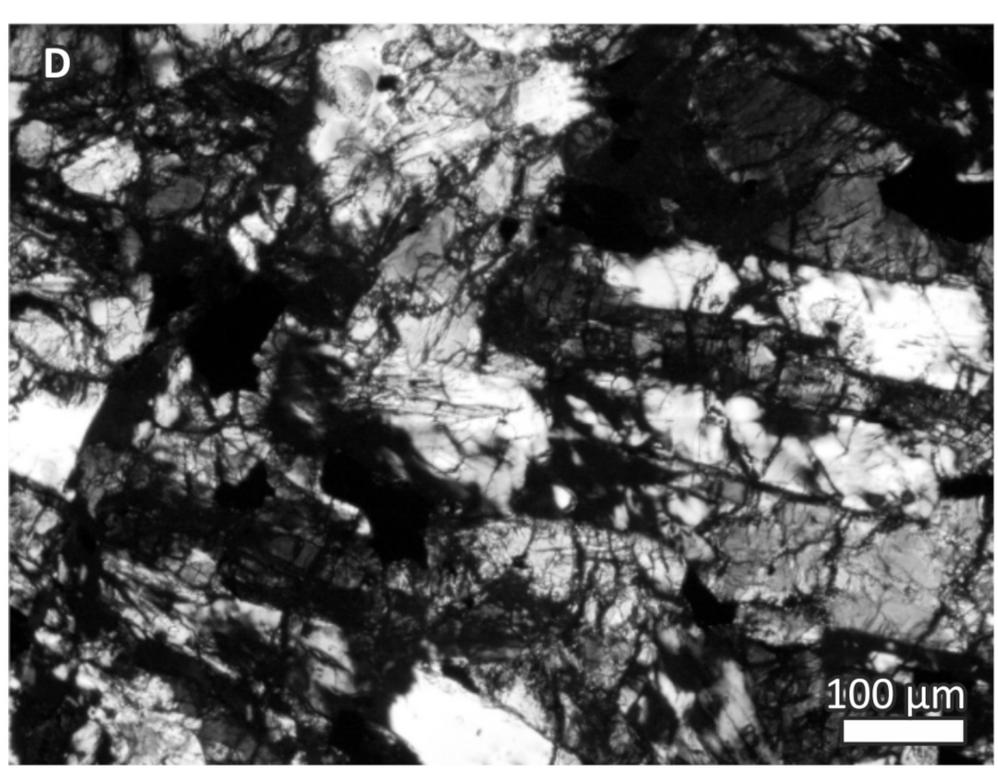
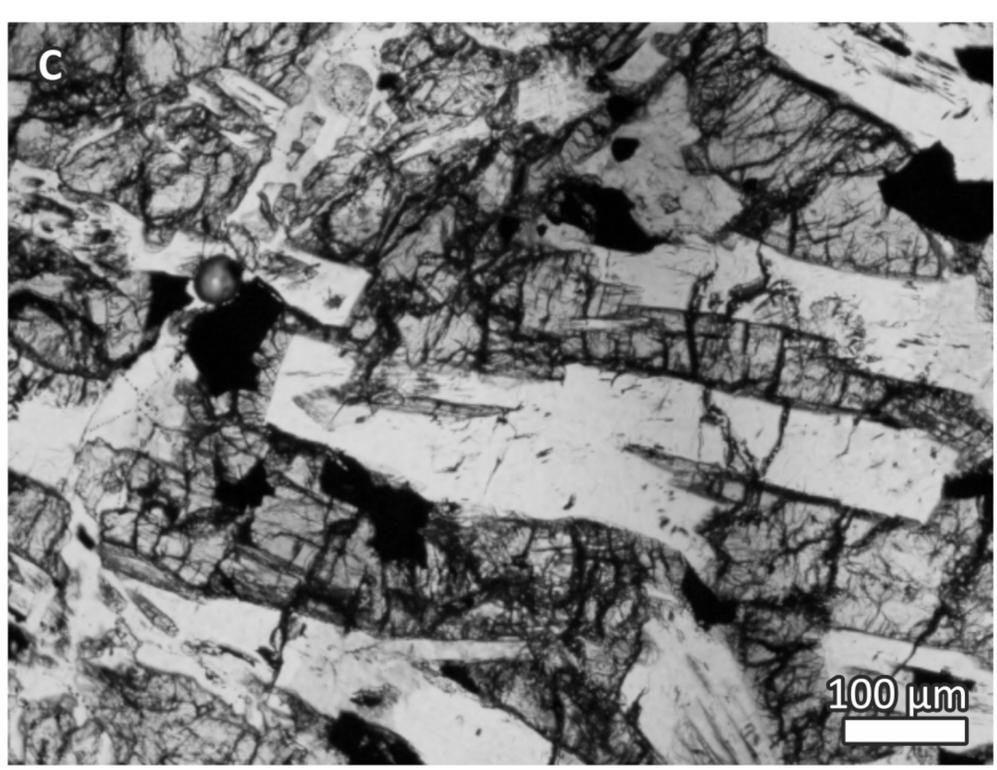
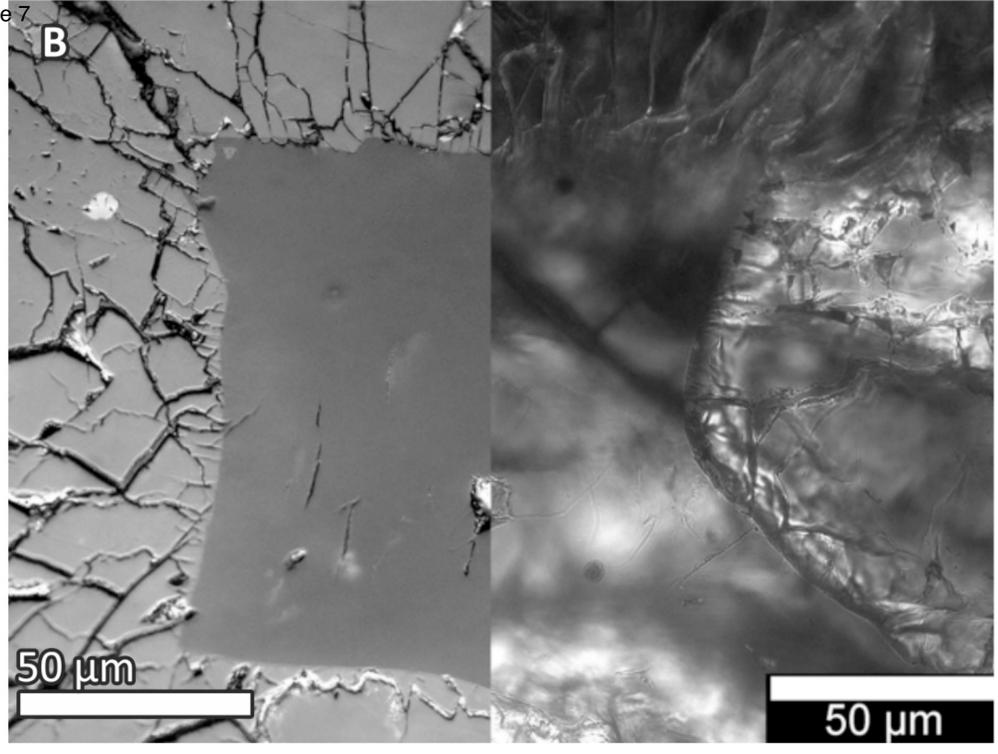
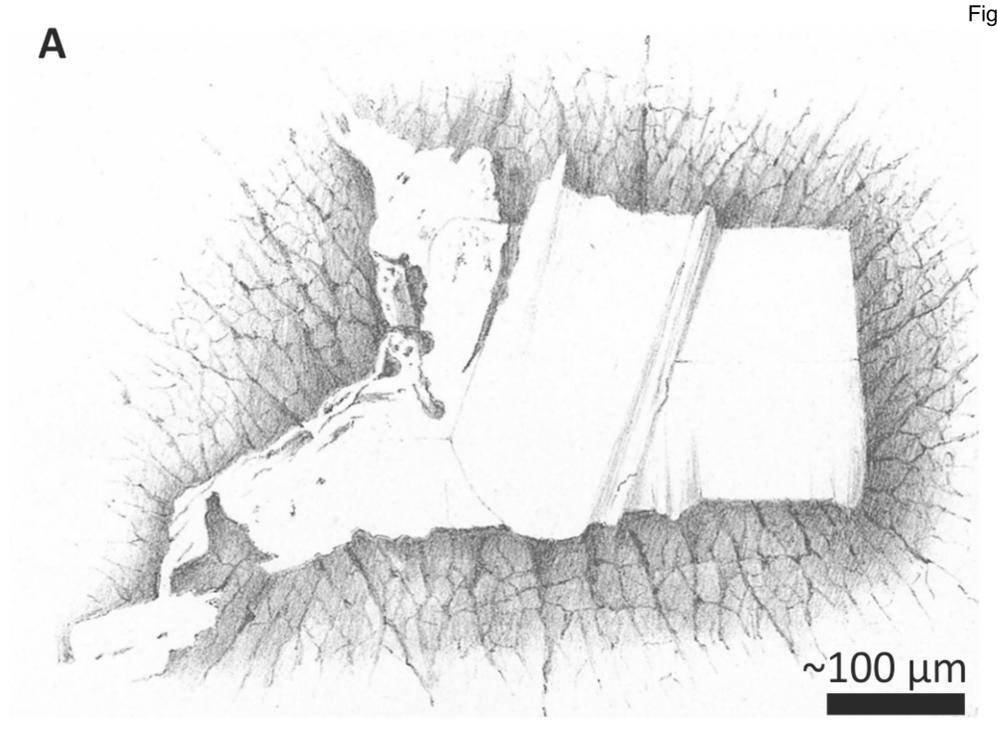
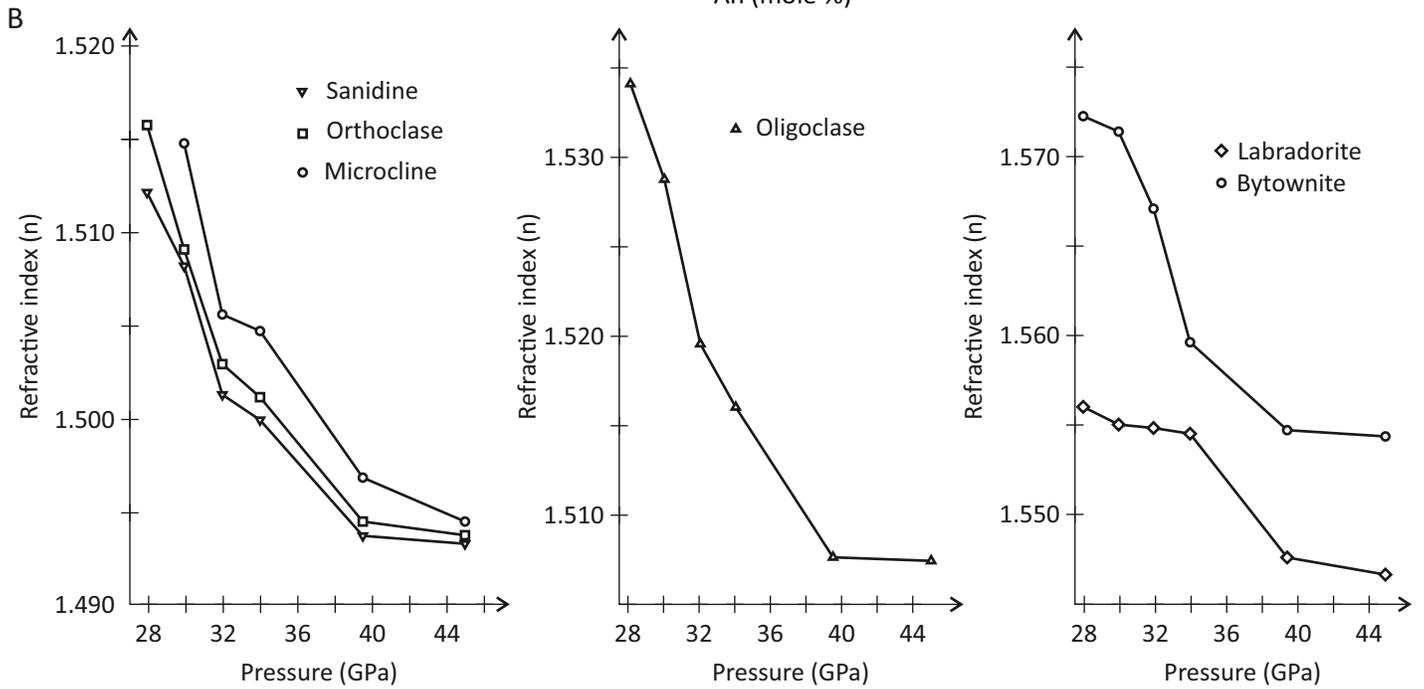
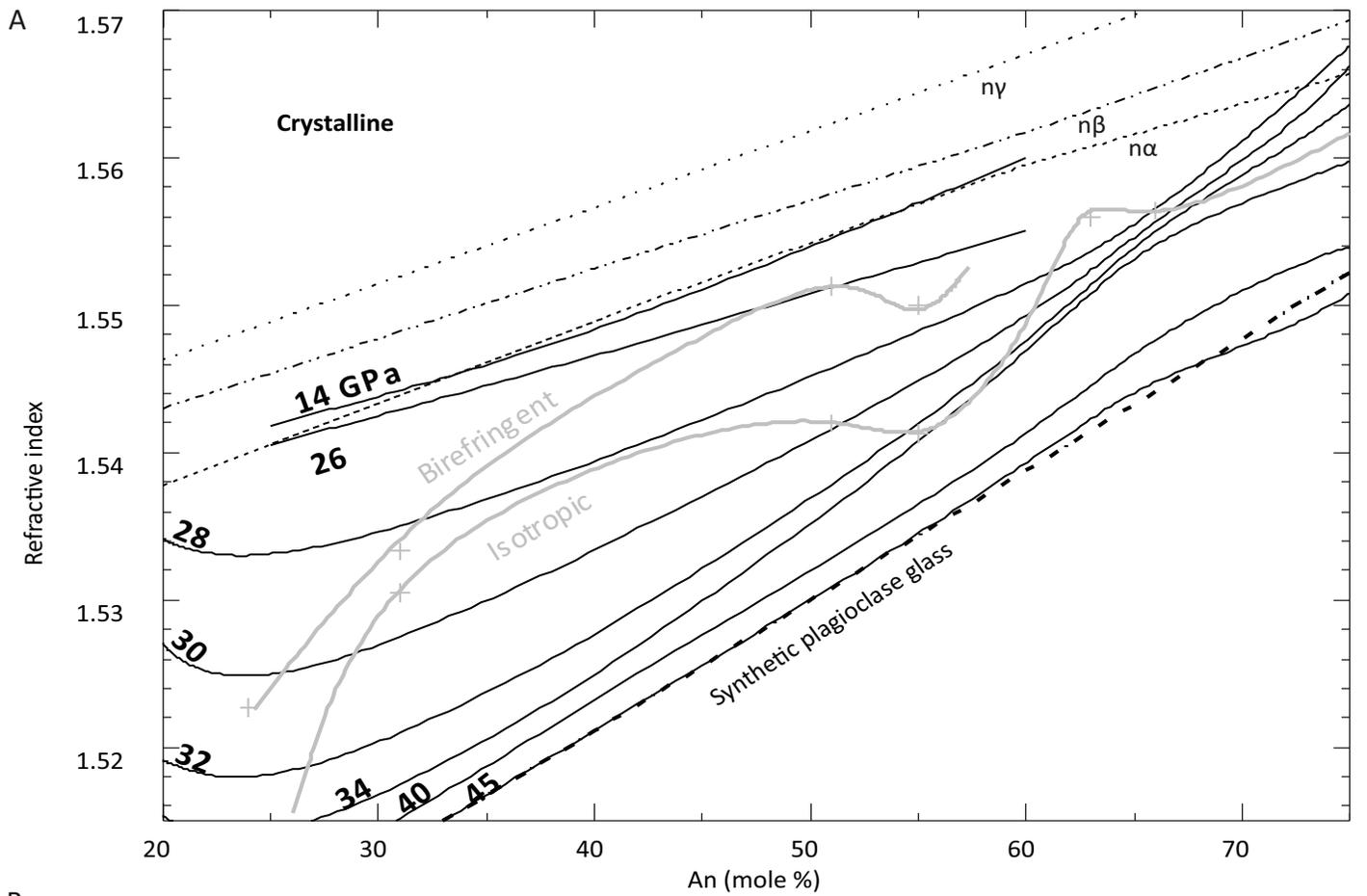


Figure 7



# Figure 8



# Figure 9

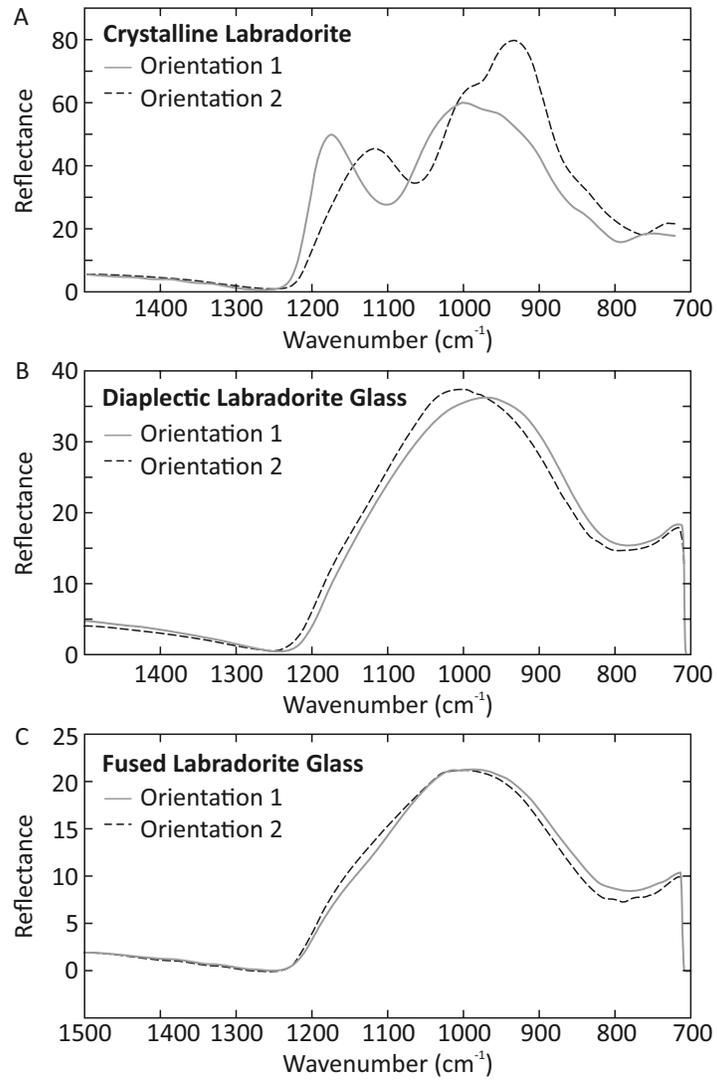
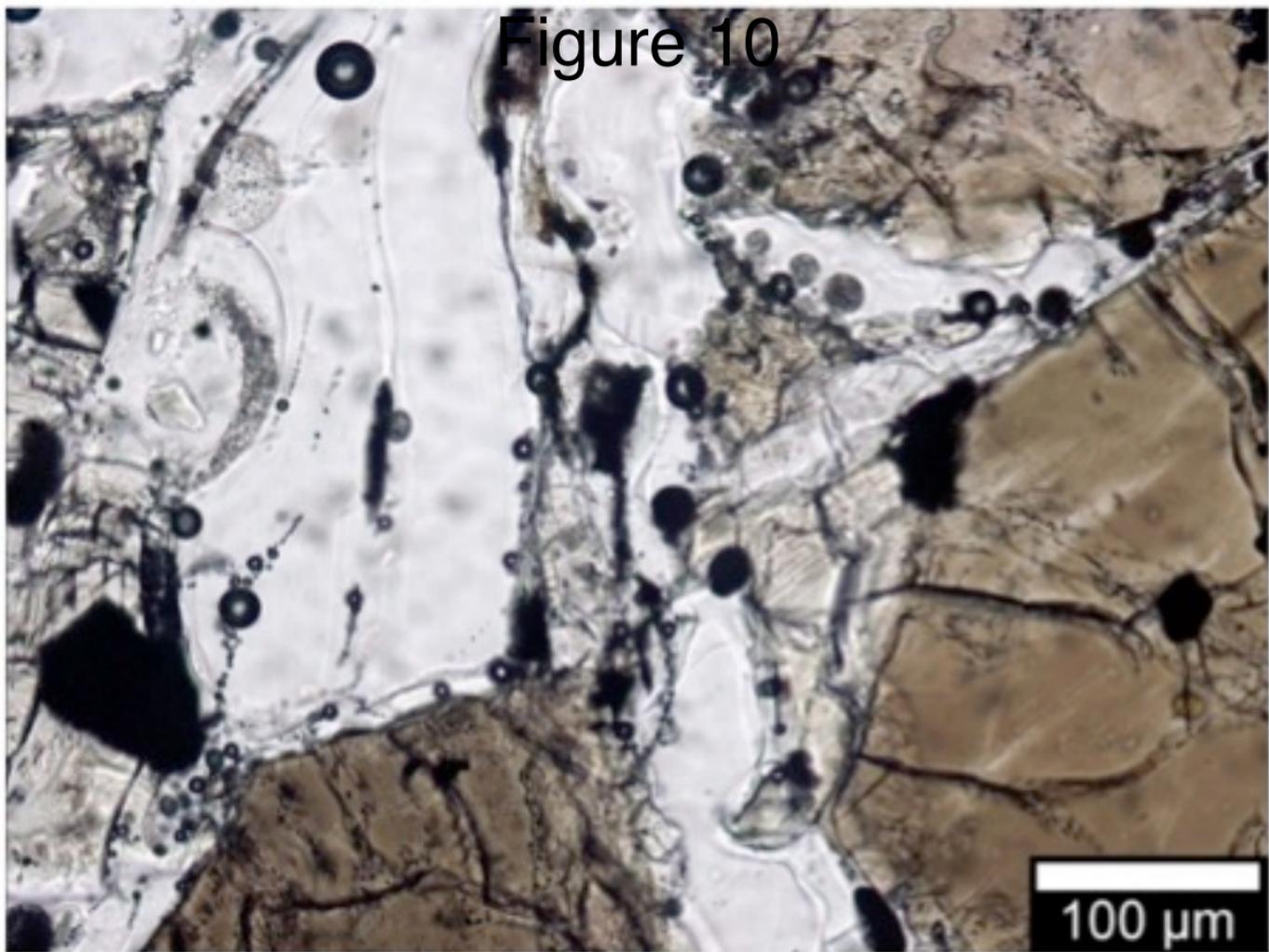
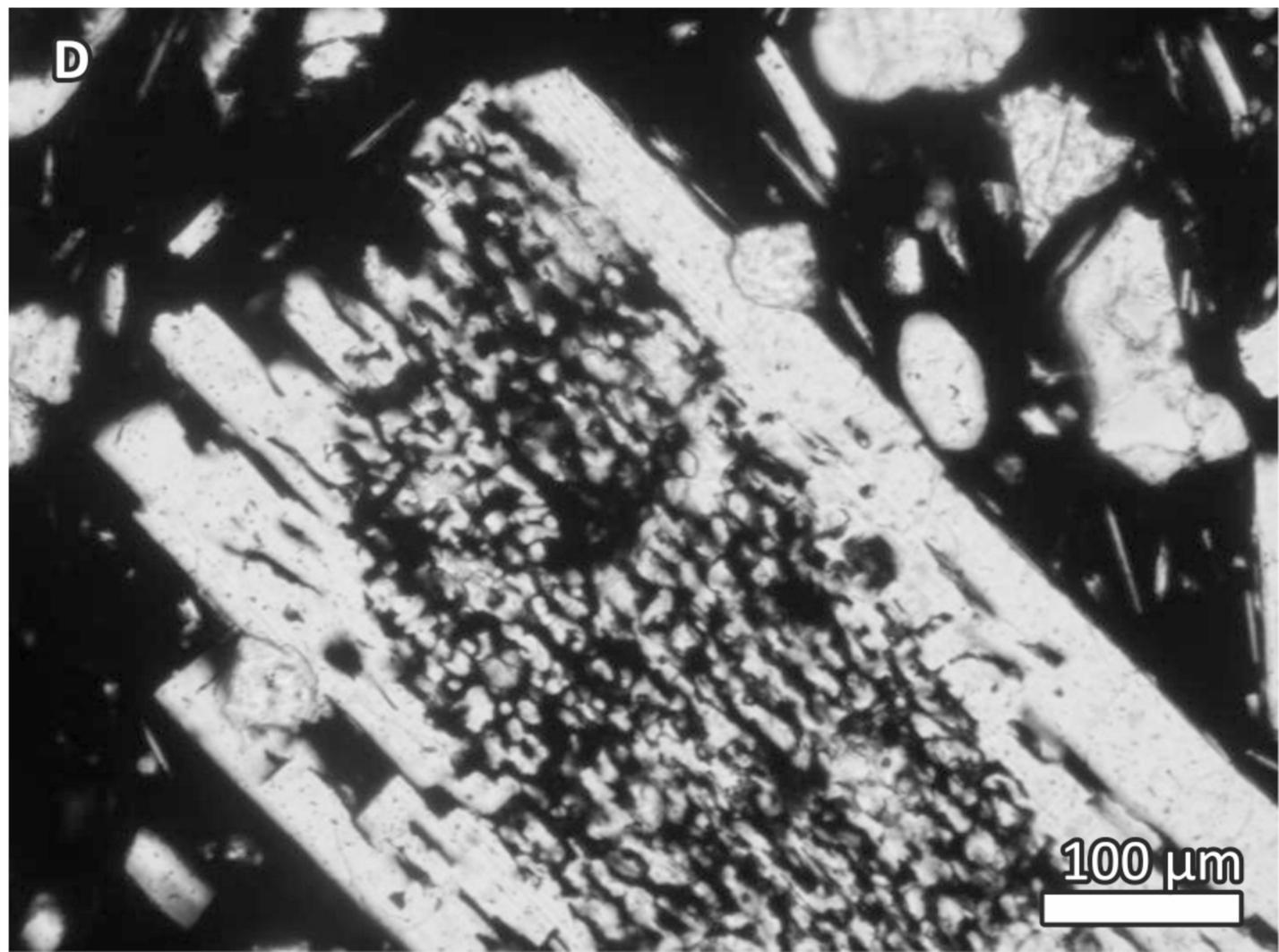
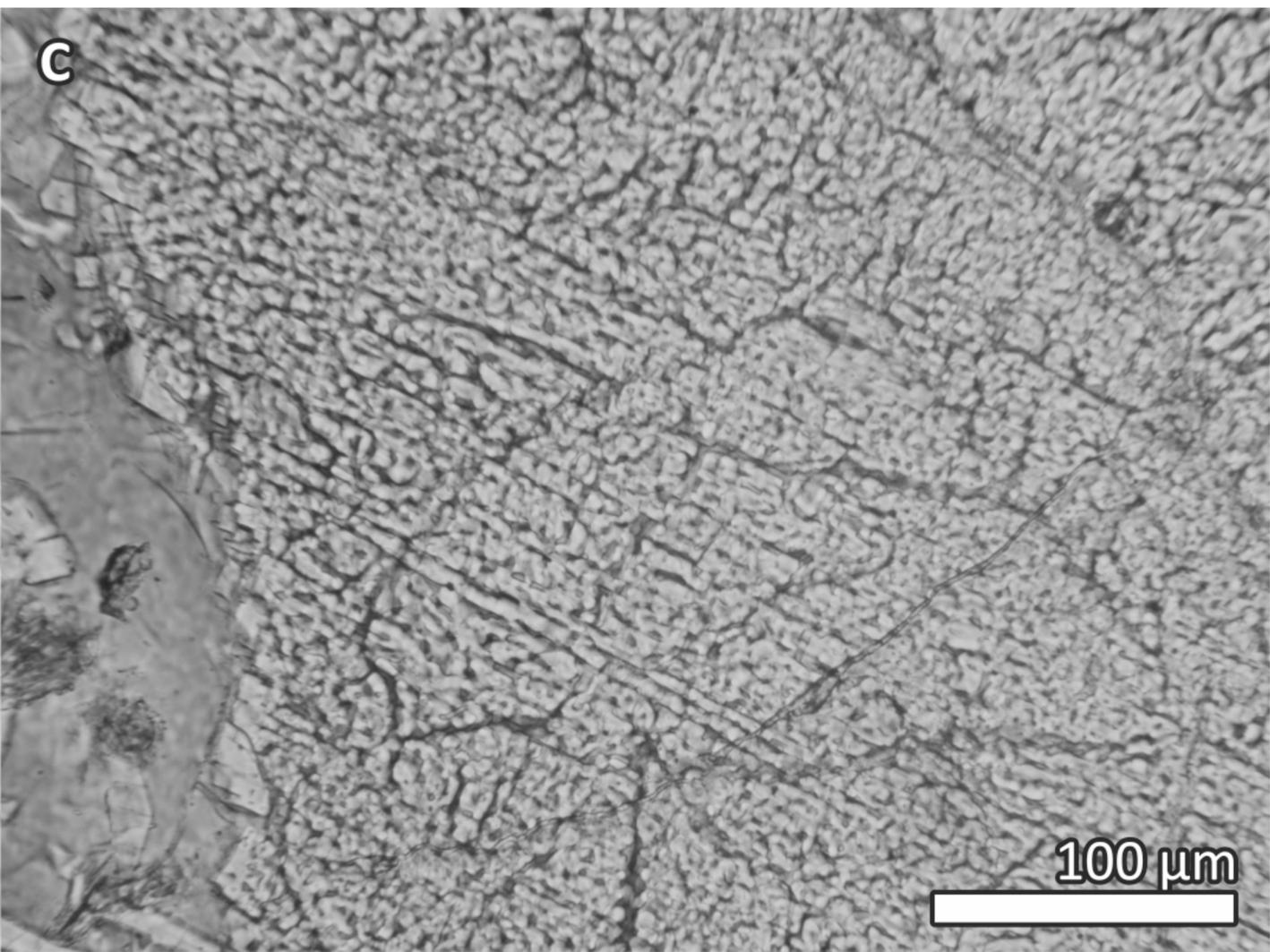
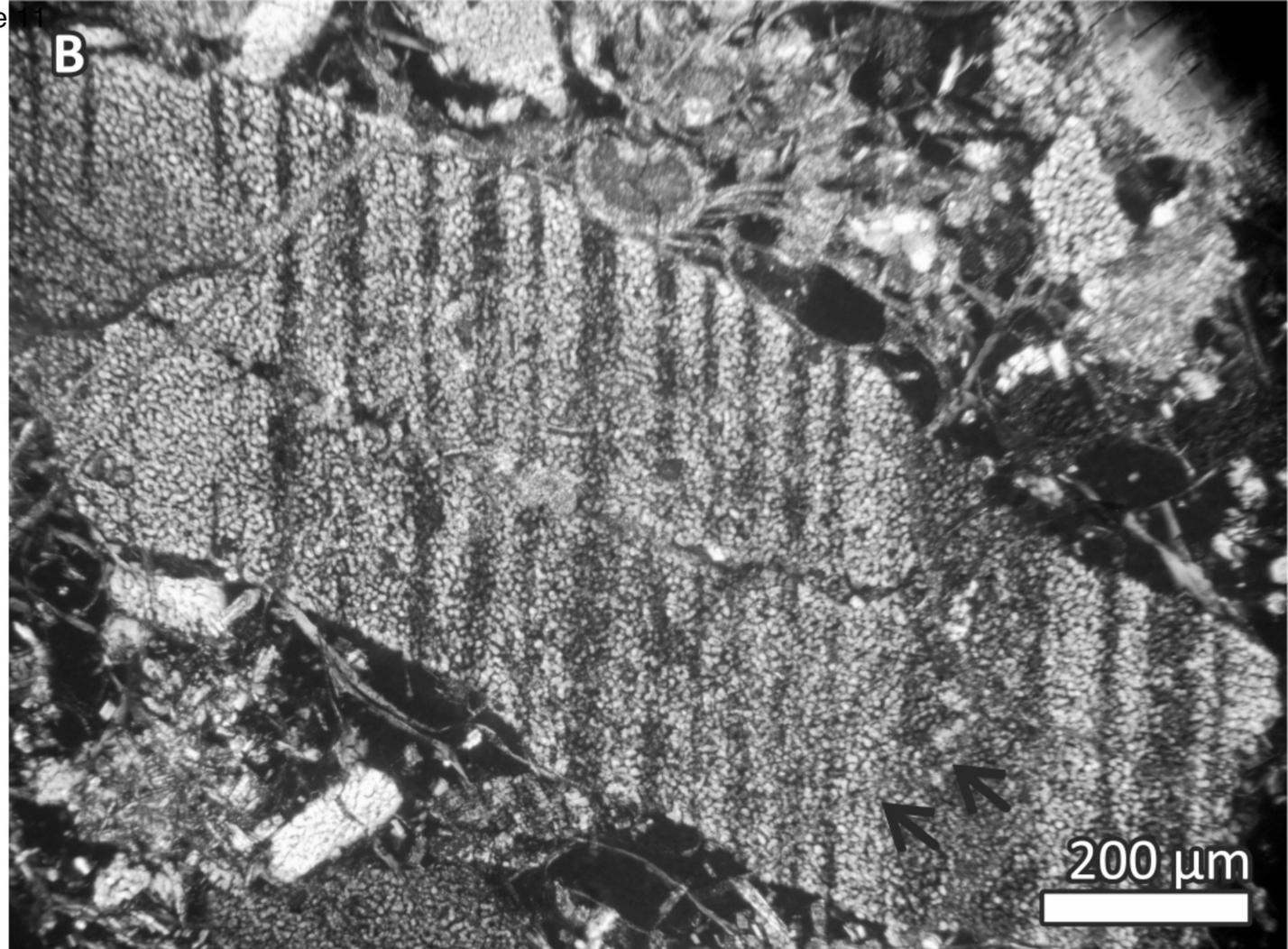
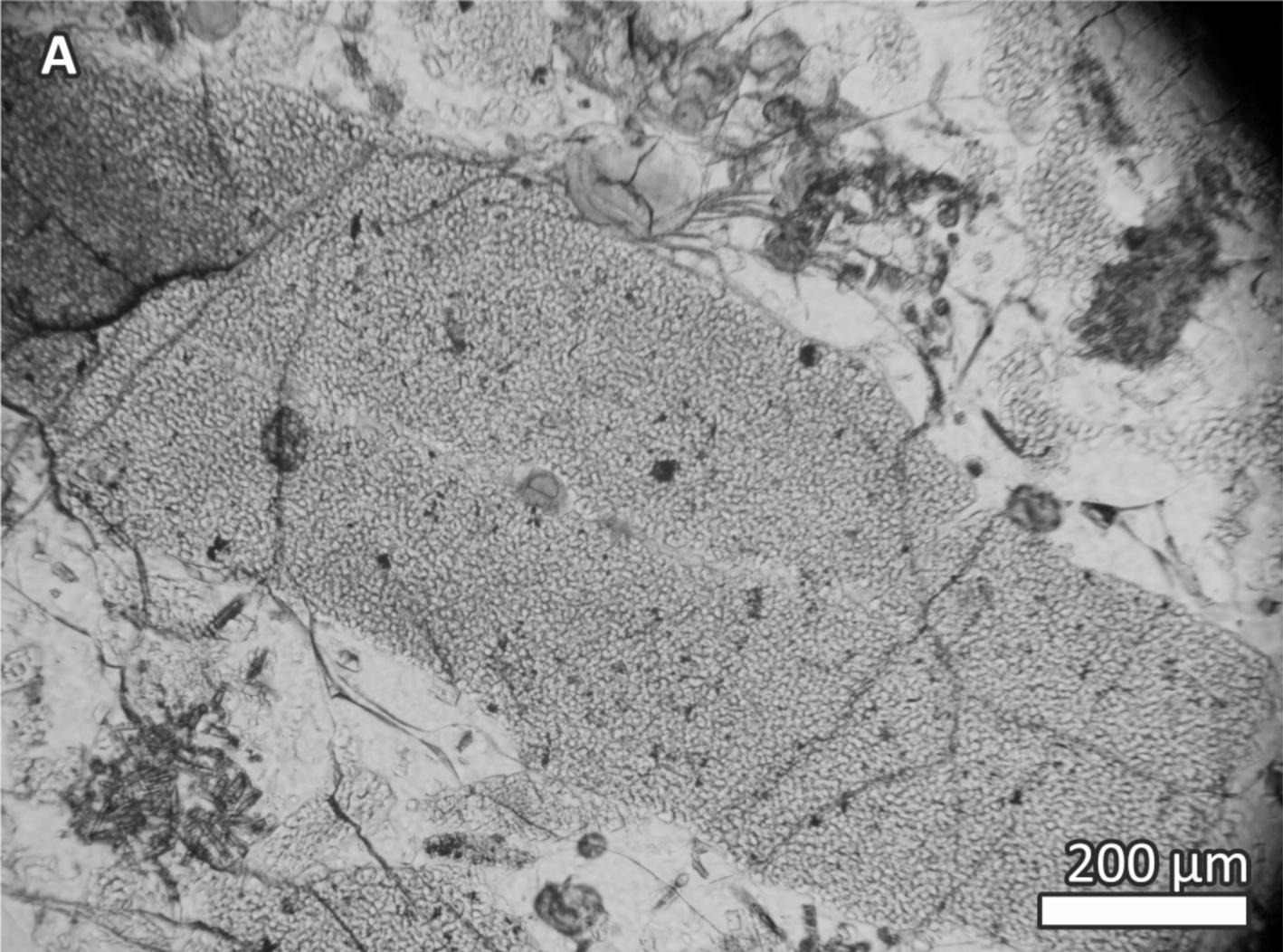
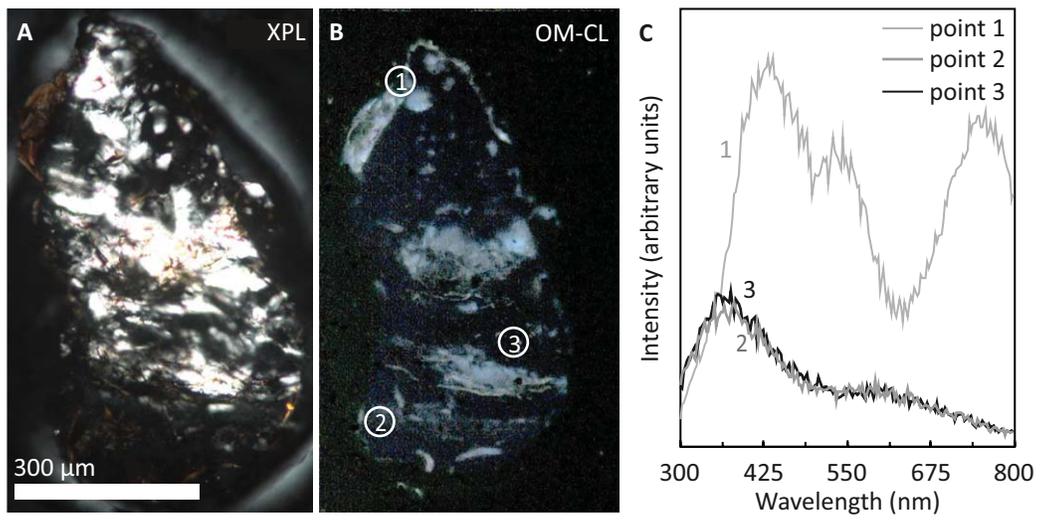


Figure 10





# Figure 12



# Figure 13

