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| 1  | Split-grain <sup>40</sup> Ar/ <sup>39</sup> Ar dating: integrating  |
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| 2  | temporal and geochemical data from crystal cargoes  |
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# 23 Abstract

24 Large sanidine crystals from the Mesa Falls Tuff (MFT), Yellowstone volcanic field, have been split and individually dated via high-precision <sup>40</sup>Ar/<sup>39</sup>Ar geochronology with the undated portions further 25 analysed for major elements, trace elements, Pb and Sr isotopes in the sanidine and trace elements in 26 27 the melt inclusions. This allows the geochemical and geochronological identity of an individual 28 sanidine to be combined. Our MFT sanidines return a preferred eruption age of  $1.3011 \pm$ 29 0.0015/0.0016 Ma (2-sigma, n=56, MSWD 0.8, analytical / full external) with a significant component of subtly older (up to ~2 Ma) crystals. Combined with recent results (Rivera et al., 2016, Journal of 30 Petrology 57, 9, 1677-1704) our data define a global mean sanidine <sup>40</sup>Ar/<sup>39</sup>Ar age for the MFT of 31 32  $1.3022 \pm 0.0006/0.0008$  Ma (2 sigma, analytical / full external) relative to Alder Creek sanidine at 1.1891 Ma and total  $\lambda_{40Ar}$  5.5305e-10, which gives  $R_{ACS}^{MFT}$ : 1.09542 ± 0.00050. The ability to couple 33 34 geochemistry and geochronology from a single grain allows us, for the first time, to evaluate the origin 35 of the subtly older sanidines present in the same pumices as juvenile sanidines. Melt inclusions from 36 all Mesa Falls sanidines represent extremely fractionated melts with low Sr contents (max. 12 ppm, 37 n=39), and rare earth element patterns which require that they be formed from an A-type magma 38 rather than the preceding subduction-related Eocene volcanism as previously suggested. <sup>87</sup>Sr/<sup>86</sup>Sr from 39 juvenile and subtly older sanidines shows the same range of 0.7073 to 0.7096 illustrating the 40 susceptibility of such low-Sr melts to slight degrees of assimilation. Pb isotopic compositions are more 41 restricted and identical between the juvenile and subtly older sanidines in the Mesa Falls Tuff (<sup>207</sup>Pb/<sup>206</sup>Pb 0.900-0.903, <sup>208</sup>Pb/<sup>206</sup>Pb 2.217-2.226, *n*=83) and these compositions rule out the 42 43 underlying Huckleberry Ridge Tuff member B as a potential source for the subtly older sanidine. LA-ICPMS <sup>206</sup>Pb/<sup>238</sup>U dating of Mesa Falls zircons supports no role for the Huckleberry Ridge Tuff. 44 45 Rather, these subtly older sanidines are interpreted as containing excess mantle-derived Ar. The ability to couple the geochemical and geochronological records within individual sanidine crystals that we 46 47 demonstrate here has potential to provide new insights for a variety of petrological studies such as 48 diffusional modelling.

49

50 Keywords: Ar/Ar geochronology, split grain, radiogenic isotopes, Yellowstone, crystal cargo

# 51 **1. Introduction**

52 With the increasing appreciation that most (if not all) magmatic systems represent open rather than 53 closed systems, the challenge for igneous petrology has become how to best exploit the complex 54 geochemical records found within a single volcanic deposit. Open system behaviour may be observed 55 in a variety of ways, most commonly it is exhibited in the complex crystal cargo evident in many 56 volcanic rocks. Complexity in crystals has been reported in major elements, trace elements, and in-situ 57 isotopic measurements (Davidson et al., 2007; Charlier et al., 2007; Stelten et al., 2015). With the 58 advent of isotopic micro-sampling and so called 'crystal isotope stratigraphy' (e.g. Davidson et al., 59 2007 and references therein), it appears that with rare exceptions isotopic heterogeneity is pervasive (e.g. Knesel et al., 1999). The ability to geochemically fingerprint individual crystals and zones within 60 61 crystals allows for different crystal populations to be discriminated. These populations are most 62 typically referred to as phenocrystic (i.e. that grew from the melt in which they are found), antecrystic (precursor grains which grew in an earlier part of the magmatic system) and xenocrystic (foreign 63 64 grains), yet even such distinctions are complicated as each term comes laden with implied meaning. Thus, the history of a population of grains within a single deposit is challenging to unravel. 65

Recent improvements in geochronological techniques, particularly in the realm of ID-TIMS 66 U/Pb and <sup>40</sup>Ar/<sup>39</sup>Ar geochronology are allowing the timescales of magmatic processes to be 67 68 investigated in unprecedented detail (e.g., Crowley et al., 2007; Schoene et al., 2010; Schmitz and 69 Kuiper, 2013). This improved precision on individual measurements is allowing subtle differences in 70 the ages of individual grains to be recognised and is serving to highlight the role of antecrystic 71 material during petrogenesis. Sluggish diffusion within zircon allows the growth of the crystal, 72 reflecting the magmatic history, to be tracked in both compositional and temporal space with either in-73 situ measurements (Wotzlaw et al., 2014, 2015) or a combination of in-situ and bulk isotope dilution 74 measurements (e.g. Rivera et al., 2014; Szymanowski et al., 2016). This coupling of compositional 75 information with high-precision geochronology provides a powerful toolkit with which to investigate 76 magmatic processes, yet hitherto it has been most fully utilised with zircon. To generate the highest

precision <sup>40</sup>Ar/<sup>39</sup>Ar ages involves the total fusion or step-heating of a potassium-rich mineral phase 77 78 (typically sanidine/anorthoclase or biotite). However, conventional approaches to dating using <sup>40</sup>Ar/<sup>39</sup>Ar methods have required that compositional and chronological information be retrieved from 79 80 different grains (i.e., data are disconnected). This disconnect can lead to a situation whereby the same 81 crystal could be considered phenocrystic or antecrystic depending on the analytical method employed. 82 The importance of this distinction is of particular importance for diffusion-based studies where zoning 83 patterns in crystals are used to infer timescales of processes immediately pre-eruption (e.g. Morgan 84 and Blake, 2006; Till et al., 2015).

This paper for the first time retrieves geochemical and textural information from a population of sanidine grains which are also dated via high precision <sup>40</sup>Ar/<sup>39</sup>Ar, and compares the chronological record in sanidine to that derived from zircons (via LA-ICPMS) to illuminate how large-volume magmas are generated using the Mesa Falls eruption from Yellowstone as a case study.

89

# 90 2. Geological Background

91 The Yellowstone volcanic field is the present-day manifestation of a long-lived thermal anomaly in the 92 inland northwest of North America which has produced the Earth's youngest large igneous province. 93 Volcanism began with the Columbia River flood basalts which began erupting  $\sim 16.5$  Ma from a series 94 of vents in northern Nevada and eastern Oregon (Hooper et al., 2007). Synchronously, rhyolitic 95 volcanism was occurring from widely spread sources across Oregon, Idaho and Nevada (Coble and 96 Mahood, 2012). At approximately 14.5 Ma, silicic volcanism became focussed along the track of the 97 Snake River Plain with a broadly time-transgressive evolution of volcanism from south-western Idaho 98 from the Bruneau-Jarbidge eruptive centre and surrounds through the central Snake River Plain to the 99 Heise and Yellowstone eruptive centres. This time-transgressive rhyolitic activity produced numerous 100 large-volume eruptions which are found as ignimbrite sheets on the margins of the plain (e.g. Ellis et 101 al., 2012a; Knott et al., 2016) and as tephra deposits which are found across much of the western 102 U.S.A. (Perkins and Nash, 2002; Nash et al., 2006). Presently, the hotspot lies directly under the

Yellowstone volcanic field as illustrated by the high heat flow and abundant geothermal activity (see
review by Lowenstern et al., 2006).

105

106 ----INSERT FIGURE 1 location map with chemistry inset----

107

108 The Yellowstone volcanic field has evolved episodically over the past ~ 2 Ma with a number 109 of large-volume eruptions separated by periods of relative quiescence during which rhyolitic and basaltic lavas were erupted (Christiansen, 2001). The Mesa Falls Tuff (MFT) covers ~2,700 km<sup>2</sup> of 110 111 south-eastern Idaho and western Wyoming (Fig. 1) with an estimated bulk volume of 280 km<sup>3</sup> 112 (Christiansen, 2001). It consists of a series of parallel-bedded fallout deposits of well-sorted, angular, 113 framework-supported pumice and coarse ash beds overlain by a non to weakly-welded ignimbrite 114 containing a variety of lithic clasts. The MFT is a high-silica rhyolite (75.8-77.8 wt.% SiO<sub>2</sub>; Fig. 1) with a mineral assemblage containing sanidine + quartz + plagioclase + augite + fayalite + ilmenite + 115 116 magnetite and accessory zircon and apatite as is typical for Yellowstone rhyolites. Isotopically, the MFT has bulk <sup>87</sup>Sr/<sup>86</sup>Sr<sub>i</sub> of 0.7084-0.7088 (Doe et al., 1982), ɛNd of -10.1 (Nash et al., 2006), and 117 <sup>206</sup>Pb/<sup>204</sup>Pb of 17.26, <sup>207</sup>Pb/<sup>204</sup>Pb of 15.56, and <sup>208</sup>Pb/<sup>204</sup>Pb of 38.25 (Doe et al., 1982). In terms of stable 118 isotopes, the MFT is normal in terms of  $\delta^{18}$ O, as is usual for the larger eruptions at Yellowstone 119 120 (Bindeman and Valley, 2001).

# 121 **3. Methods**

The samples used for this study were collected from a classic locality of the MFT exposed on the west side of highway 20, approximately 6 km north of the town of Ashton, Idaho (Fig. 1). Large single sanidine crystals up to 1 cm in diameter were separated from larger fragments of pumice with crystals hand-picked under a binocular microscope. Crystal surfaces show no evidence of alteration and were typically transparent. Selected sanidines were then cleaned using dilute HNO<sub>3</sub>. Grains without adhering glass and the fewest melt inclusions were selected for further work. Sanidines were split into

| 128 | smaller fragments (typically two, but occasionally three or four) using a clean razor blade and a                              |
|-----|--|
| 129 | fragment less than 2 mm in diameter was selected for <sup>40</sup> Ar/ <sup>39</sup> Ar dating with the other fragments placed |
| 130 | into separate labelled vials for geochemical analyses. The grains for <sup>40</sup> Ar/ <sup>39</sup> Ar dating were then      |
| 131 | individually re-cleaned in methanol to remove any debris from the cutting process (Fig. 2).                                    |
| 100 |  |

132

133 ----INSERT FIGURE 2 of grains cut----

134

# 135 **3.1**<sup>40</sup>Ar/<sup>39</sup>Ar geochronology

<sup>40</sup>Ar/<sup>39</sup>Ar analyses were conducted at the NERC Argon Isotope Facility, Scottish Universities

137 Environmental Research Centre (SUERC). Details of irradiation durations, irradiation correction

138 factors and discrimination corrections are provided in appendix file SF#1 (.xls).

139 Two dating experiments were conducted, which required samples to be irradiated in separate 140 batches. Experiment 1: bulk sanidine separates harvested from the MFT ignimbrite (90118) and MFT 141 fallout (90116) were loaded in two large wells in an Al disc for irradiation. Alder Creek sanidine was 142 loaded in to 4 wells surrounding the two samples to allow accurate determination of the J-parameter. 143 These samples were irradiated in the CLICIT facility of the OSU reactor for 120 minutes. These 144 samples were analysed prior to progressing to phase II (split grain approach). Experiment 2: 145 approximately 12 months later the grain fragments produced from cutting the large sanidine crystals 146 (90118) were loaded in to a 21 pit Al disc for irradiation (5 discs in total). Each single fragment had its 147 own well position (17 fragments in each Al disc). For J determinations Alder Creek sanidine was 148 loaded into 5 different positions (centre, top, bottom, left, right - see SF#1) in each Al disc. This 149 allowed for determination of an accurate J parameter for each individual disc. These samples were 150 irradiated in the CLICIT facility of the OSU reactor for 120 minutes.

151 Single crystals of Alder Creek sanidine (ACs) were fused using a CO<sub>2</sub> laser and the argon 152 isotope composition of cleaned extracted gases analysed using a MAP 215-50 noble gas mass 153 spectrometer using methods described by Mark et al., (2008) and Smith et al., (2011). For determination of the J parameter for each Al disc the weighted average  ${}^{40}\text{Ar}*/{}^{39}\text{Ar}_{K}$  was calculated for each well, and the arithmetic mean and standard deviation of the  ${}^{40}\text{Ar}*/{}^{39}\text{Ar}_{K}$  for all standard wells from each different Al disc were used to characterise the neutron fluence for the unknowns. This approach was deemed sufficient, as due to the relatively short irradiation durations there was no significant horizontal variation in J parameter across the irradiation holder.

159 The unknown samples were analysed in exactly the same way as the standards. The standard 160 and unknowns were measured within a couple of days of each other. Background and mass 161 discrimination measurements (via automated analysis of multiple air pipettes) specific to each batch 162 are summarised in appendix file SF#1 (.xls). Air pipettes were run (on average) after every 5 analyses. 163 Backgrounds were measured following every two measurements and subtracted from ion beam measurements were arithmetic averages and standard deviations. Mass discrimination was computed 164 165 based on a power law relationship (Renne et al., 2009) using the isotopic composition of atmospheric Ar (Lee et al., 2006) that has been independently confirmed (Mark et al., 2011). Corrections for 166 radioactive decay of <sup>39</sup>Ar and <sup>37</sup>Ar were made using the decay constants reported by Stoener et al., 167 (1965) and Renne and Norman (2001), respectively. Ingrowth of <sup>36</sup>Ar from decay of <sup>36</sup>Cl was 168 corrected using the <sup>36</sup>Cl/<sup>38</sup>Cl production ratio and methods of Renne et al., (2008) and was determined 169 to be negligible. Argon isotope data corrected for backgrounds, mass discrimination, and radioactive 170 171 decay and ingrowth are given in the appendix file SF#1 (.xls).

172 Ages were computed from the blank-, discrimination- and decay-corrected Ar isotope data 173 after correction for interfering isotopes based on the following production ratios, determined from fluorite and Fe-doped KAlSiO<sub>4</sub> glass:  $({}^{36}Ar/{}^{37}Ar)Ca = (2.650 \pm 0.022) \times 10^{-4}$ ;  $({}^{38}Ar/{}^{37}Ar)Ca = (1.96 \pm 0.022) \times 10^{-4}$ ;  $({}^{3$ 174 0.08) x 10<sup>-5</sup>;  $({}^{39}\text{Ar}/{}^{37}\text{Ar})\text{Ca} = (6.95 \pm 0.09) \text{ x } 10^{-4}$ ;  $({}^{40}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})\text{K} = (7.3 \pm 0.9) \text{ x } 10^{-4}$ ;  $({}^{38}\text{Ar}/{}^{39}\text{Ar})$ ;  $({}^{38}\text{Ar}/{}^{$ 175  $(1.215 \pm 0.003) \times 10^{-2}$ ;  $({}^{37}\text{Ar}/{}^{39}\text{Ar})\text{K} = (2.24 \pm 0.16) \times 10^{-4}$ . Ages and their uncertainties are calculated 176 using the methods of Renne et al., (2010), the decay constant parameters of Renne et al., (2011) and 177 178 the ACs age of Niespolo et al.,  $2016 (1.1891 \pm 0.0008 \text{ Ma})$ , except where noted. For some of the age comparisons made herein, contributions from sources of systematic uncertainty (i.e., uncertainties in 179 <sup>40</sup>Ar/<sup>40</sup>K of the standard and <sup>40</sup>K decay constants) are neglected and only analytical uncertainties in 180

isotope measurements of samples and standards are included. These uncertainties are referred to herein as "analytical precision". For the purposes of this study analytical uncertainties include contributions from uncertainties in the interference corrections because these interference corrections have variable effects due to the slightly variable chemistry of the samples considered. Where not otherwise distinguished, uncertainties are stated as  $X \pm Y/Z$ , where Y is the analytical uncertainty as defined above, and Z is the full external precision considering both analytical and systematic sources of uncertainty (e.g., decay constant).

188

# 189 **3.2 Geochemistry**

190 The sub-grains for geochemistry were imaged on JEOL JSM-6390 SEM at ETH Zurich (full collection 191 of images in supplementary data) prior to analysis. Major elements were determined via electron 192 microprobe at ETH Zurich using analytical conditions of 15 kV and 15 nA with a beam diameter of 10 193 µm and counting times were shortened for Na and K to avoid element mobility. Trace elements in 194 sanidines, zircons and melt inclusions were determined via LA-ICPMS using a 193 nm ArF Excimer 195 laser from Resonetics coupled to a Thermo Element XR ICPMS within the Institute for Geochemistry and Petrology, ETH, Zurich (full data in supplementary materials). For trace elements spot sizes were 196 197 either 67 microns for sanidines or 30 microns for zircons and melt inclusions. NIST-612 was used as a 198 primary standard and GRD-1 as the secondary standard. Trace element data were reduced following 199 procedures described in Szymanowski et al., (2015) and are considered to be precise to better than 5% 200 of the reported value.

Lead isotope compositions in individual sanidine crystals were measured using laser ablation
multicollector inductively-couple plasma mass spectrometry (LA-MC-ICP-MS), using a Photon
Machines G2 excimer laser ablation system and NuPlasma MC-ICP-MS at Oregon State University.
Analytical techniques followed those given in Kent (2008) using Faraday Cup detectors to measure all
required isotopes. Analyses were made using a 85 µm laser spot translated at 5 µm/sec and using a
pulse frequency of 7 Hz. Measured ratios were corrected for mass bias based on measurement of
NIST-612 glass at similar ablation conditions throughout the analysis session, and an accepted value

of  ${}^{208}\text{Pb}/{}^{206}\text{Pb} = 2.1694$ ). Multiple replicate analyses of the NIST-612 and BCR-2G standard glasses (full data in supplementary materials) show that measured  ${}^{208}\text{Pb}/{}^{206}\text{Pb}$  and  ${}^{207}\text{Pb}/{}^{206}\text{Pb}$  ratios are within 0.1% of the accepted values.

211 Strontium isotope analyses on grain portions were performed at the Institute of Geochemistry 212 and Petrology, ETH Zurich. Samples were digested with concentrated HF/HNO<sub>3</sub> in Teflon beakers, 213 then evaporated and redissolved in 6 N HCl. All solutions were evaporated and the dry residues were 214 converted in 2.5 N HNO<sub>3</sub>, followed by Sr separation in ion exchange columns. The chromatography 215 technique used PP columns with Sr- and TRU-spec resin. Strontium isotope ratios were measured on a 216 ThermoFisher multi-collector TritonPlus thermal ionization mass spectrometer (TIMS) using single 217 outgassed Re-filaments. Strontium isotopic analysis was conducted in a static mode. The Sr isotope ratios are mass fractionation corrected to  ${}^{88}$ Sr/ ${}^{86}$ Sr = 8.375209; the NBS-987 standard measurements 218 219 returned  ${}^{87}$ Sr/ ${}^{86}$ Sr of 0.710246 ± 0.0000034 (2 SE; n=14) during the period of analysis. In most cases 220 Rb and Sr concentrations were determined by isotope dilution using a highly enriched <sup>87</sup>Rb/<sup>84</sup>Sr spike, 221 while in other cases the Rb and Sr contents derived from LA-ICPMS determinations were used. 222  $^{87}$ Sr/ $^{86}$ Sr ratios were age corrected using either the precise age of the individual grain, or in rare cases 223 an assumed age of 1.3 Ma. All results are provided in supplementary material.

224

3.3 Zircon U/Pb geochronology

226 Zircons for U-Pb geochronology were separated from pumices of the MFT via classical magnetic and 227 density separation techniques. A subset of these were prepared via chemical abrasion following the 228 procedure of Mattinson (2005) and mounted along with untreated zircons from the MFT to assess the 229 effects of chemical abrasion on zircons of this age. Grains were imaged via cathode luminescence 230 (CL) at the ScopeM facility of ETH Zurich using a FEI Quanta 200 FEG scanning electron 231 microscope. Following CL imaging both chemically abraded and untreated zircons were analysed via 232 LA-ICPMS with the whole procedure of analysis and data reduction following that described by 233 Guillong et al., (2014).

# 235 **4. Results**

# 236 **4.1**<sup>40</sup>Ar/<sup>39</sup>Ar sanidine geochronology

The two samples selected for conventional <sup>40</sup>Ar/<sup>39</sup>Ar dating, the MFT ignimbrite (#90118) and the 237 238 MFT fallout (#90116) yield ages of  $1.3017 \pm 0.0040$  Ma and  $1.2993 \pm 0.0033$  Ma, respectively (2-239 sigma, analytical precision). Both samples showed extensive contamination with crystals older than 240 the defined juvenile population with large numbers of individual data points (ages) rejected from 241 calculation of the weighted mean eruption age, 74% for the ignimbrite and 25% for the fallout. The 242 occurrence of subtly older sanidine crystals has been observed in a number of other studies of Yellowstone rhyolites including deposits erupted both explosively and effusively (Gansecki et al., 243 244 1996; Ellis et al., 2012b; Rivera et al., 2014; Singer et al., 2014; Stelten et al., 2015; Matthews et al., 245 2015; Troch et al., 2017). Owing to the large degree of contamination and the data not defining clear 246 juvenile age populations, the Mass Spec programme was used to screen the data to provide an age 247 using the 'youngest gaussian' filter with a MSWD probability cutoff of 0.05.

Owing to the large degree of antecrystic contamination in the ignimbrite we focussed on this sample for the follow-up study – the split grain work (#91235-#91309) to investigate the linking of geochemical-geochronological data for the same grain. Taking this group of split grains as a whole sample population they define an  $^{40}$ Ar/ $^{39}$ Ar age of 1.3015 ± 0.0022 Ma (2 sigma, analytical precision). The same statistical filter was used for these data as above. As expected, a large percentage (67%) of the data were rejected from calculation of the mean eruption age for the split grains, attesting to the presence of a significant older population.

Because of the unique way the split grains were irradiated and given the range of sanidine ages, the potential for grain to grain variability in J-value to be affecting the calculated age should be addressed. Even following splitting, the grains within this study remain relatively large (Fig. 2) and so it is conceivable that some grains may be acting to shield others from the fast neutron bombardment during irradiation. We can evaluate the potential for this in two different ways. Firstly, when the calculated ages of grains are plotted against pan position during irradiation (supplementary figure 4),
 no relationship is apparent. Secondly, the amount of <sup>39</sup>Ar can be plotted against age of the grain to test
 whether shielding of the interior of large grains is occurring. The lack of correlation between the
 amount of <sup>39</sup>Ar and the age of the crystal suggests that this is not an issue.

The <sup>40</sup>Ar/<sup>39</sup>Ar ages for samples from the two experiments are indistinguishable. We have thus 264 determined a global weighted average  ${}^{40}\text{Ar}/{}^{39}\text{Ar}$  age for the MFT:  $1.3011 \pm 0.0015/0.0016$  Ma (2-265 266 sigma, n56, MSWD 0.8). This eruption age for the Mesa Falls Tuff is in excellent agreement with the 267 published ID-TIMS zircon ages (Wotzlaw et al., 2015; Rivera et al., 2016) and with the most recent 268  $^{40}$ Ar/ $^{39}$ Ar age of Rivera et al., (2016). We note that the  $^{40}$ Ar/ $^{39}$ Ar of Rivera et al., (2016) is calculated 269 using a different Alder Creek sanidine age (Rivera et al., 2013) and a different decay constant (Min et 270 al., 2000). Despite the use of different calibrations the data are indistinguishable. The data of Rivera et 271 al., (2016) recalculated relative to the calibration employed by this study becomes  $1.3027 \pm 0.0009$  Ma 272 (analytical precision), which is indistinguishable  $(1.6 \pm 1.3 \text{ ka})$  from our age for MFT. Taking both 273 datasets (this study and Rivera et al., 2016) we can calculate a global mean age for the MFT of 1.3022  $\pm 0.0006/0.0008$  Ma (2 sigma), which defines  $R_{ACs}^{MFT}$ : 1.09542  $\pm 0.00050$ . 274

275

276 ---- INSERT FIGURE 3 new Ar/Ar geochronology----

277

# **4.2 Sanidine textural and compositional data**

Individual grain sizes for the 'geochemistry sub-grains' were typically in the range of a few hundred microns (images of all grains used for in-situ geochemistry in supplementary materials). Backscattered electron images indicate limited compositional zonation and that about 60% of the crystals contain melt inclusions which make up less than 5% of the observed section (e.g. grains 2, 5, 42, and 66 in supplementary materials). The frequency of melt inclusions in the geochemistry sub-grains was heightened by them being excluded from the portions of crystals sent for dating.

In terms of major element compositions, the Mesa Falls sanidines (n=166) show little variation 285 with compositions of Or<sub>59-63</sub> and no coherent sense of zonation within a crystal. These compositions 286 287 and the occurrence of major elemental homogeneity within a crystal agree well with previously 288 reported MFT sanidine data (Gansecki et al., 1998). Unlike the major elements, trace elements in the 289 Mesa Falls sanidines do show large compositional ranges which in the most extreme case, Ba, span 290 two orders of magnitude (69-8721 ppm). Large ranges are also observed in Ti (34-104 ppm), Sr (26-291 209 ppm) and Rb (91-152 ppm). Interestingly, despite the large range in overall compositions, no 292 coherent sense of zonation is observed when plotting rims against cores of individual grains, nor is 293 there compositional variability as a function of crystal age. While our analyses are best efforts at 294 analysing rims and cores of grains we note with the split grain method the discrimination of rims and 295 core of grains without concentric zonation remains problematic. Nevertheless, given the number of 296 grains (n=84) in this study, it is likely that the true compositional range of Mesa Falls sanidine has 297 been characterised.

298

299 ----INSERT FIGURE 4 Compositional data ----

300

301 Melt inclusions hosted in Mesa Falls sanidines are large and homogeneous (as shown in 302 supplementary material) allowing for trace elemental determinations to be carried out via LA-ICPMS. 303 The melt inclusions within the sanidines are similar to other Yellowstone glasses (e.g. Vazquez et al., 304 2009) and represent extremely fractionated rhyolite liquids as illustrated by the low Sr contents of 2-305 12 ppm and strong negative Eu anomaly (Fig. 5). Overall, the relatively flat normalised rare earth 306 element (REE) patterns of the MFT melt inclusions are identical to those observed in other rhyolites of 307 the province (e.g. Szymanowski et al., 2015) and in contrast to the steeper patterns observed from 308 subduction zone rhyolites. Notably, REE patterns in glassy melt inclusions are identical between 309 juvenile and subtly older sanidine grains (Fig. 5).

310

311 ----INSERT FIGURE 5 REE in melt inclusions----

312

| 313 | 4.3 Radiogenic isotopes  |
|-----|--|
| 314 | Sanidine within the MFT exhibits only slight variability between grains, with <sup>208</sup> Pb/ <sup>206</sup> Pb ratios        |
| 315 | spanning 2.217-2.226 and $^{207}$ Pb/ $^{206}$ Pb ratios spanning 0.900-0.903 (full data in supplementary                        |
| 316 | materials). The majority of this variability is contained within the precision of the measurements                               |
| 317 | (taken as 2 standard errors) and sanidine LA-ICPMS values are in good agreement with the bulk Pb                                 |
| 318 | isotopic data for the Mesa Falls Tuff published by Doe et al., (1982). When compared to all Pb                                   |
| 319 | isotopic data from the Yellowstone volcanic field (Fig. 6), it is clear that the Mesa Falls Tuff exhibits                        |
| 320 | relatively little variability. In Pb isotopic space all sanidines appear identical within analytical                             |
| 321 | precision, regardless of the age of the crystal (Fig. 6).  |
| 322 |  |
| 323 | INSERT FIGURE 6 Pb isotope data for MFT and all Yellowstone  |
| 324 |  |
| 325 | Mesa Falls sanidines show a large degree of isotopic variability in <sup>87</sup> Sr/ <sup>86</sup> Sr with 35 grains            |
| 326 | covering a range from 0.707355 $\pm$ 0.000167 to 0.709558 $\pm$ 0.000012 (Fig. 7, full data in                                   |
| 327 | supplementary material). That the range in <sup>87</sup> Sr/ <sup>86</sup> Sr is significantly larger than that displayed via Pb |

328 isotopes (above) is unsurprising given the relatively low Sr contents of the Yellowstone rhyolitic 329 liquid (2-11 ppm in the melt inclusions) rendering them extremely susceptible to small degrees of 330 contamination. Indeed, such Sr isotopic variability appears relatively common in rhyolites (Charlier et al., 2007; Davidson et al., 2007). The <sup>87</sup>Sr/<sup>86</sup>Sr does not appear to be related to the trace elemental 331 332 identity of the crystal with variations in Ba contents of more than 5,000 ppm observed between crystals with identical isotope ratios. Such decoupling of trace elements from isotopes may be a result 333 334 of the variable effects of cumulate remelting which readily returns elements sequestered in cumulates (such as Ba and Sr) to the melt without significantly altering isotopic compositions (e.g. Wolff et al., 335 2015). As might be expected when comparing sub-crystal scale compositions to averages of whole 336

pumice clasts, the range of <sup>87</sup>Sr/<sup>86</sup>Sr returned from the split grains is significantly greater than the
values of bulk Mesa Falls Tuff of 0.7084 and 0.7088 (Doe et al., 1982) and the 0.70868 of Hildreth et
al., (1991).

340

341 ---- INSERT FIGURE 7 Sr variability in here----

342

## 343 **4.4 The record from zircon**

344 Zircons from the MFT commonly show oscillatory zonation in CL images with a lack of notably 345 bright zones or dark cores (images in supplementary material). In terms of trace element composition, 346 Ti contents vary from below detection limit (in two cases) to 16 ppm which is similar to the range 347 observed previously for Yellowstone zircons (Rivera et al., 2014, 2016; Stelten et al., 2015; Wotzlaw 348 et al., 2015; compositional data are provided in supplementary materials). While analyses optimised 349 for trace elemental abundances did not return high U and Th values, some spots analysed for U-Pb did 350 find elevated values similar to those observed by Bindeman et al., (2008) in their two Mesa Falls 351 analyses.

LA-ICPMS  ${}^{206}$ Pb/ ${}^{238}$ U ages derived for zircons from the MFT range from 1.133  $\pm$  0.073 Ma to 352 353  $1.634 \pm 0.079$  Ma (individual uncertainties at 2 standard error). As illustrated in Figure 8, with the 354 exclusion of three outliers from the total population of 65 grains that were used for final age 355 determination, the average ages of the chemically abraded and untreated grain populations are 356 indistinguishable. However as is shown in the supplementary material, the proportion of discarded 357 analyses due to common Pb contamination is significantly reduced following chemical abrasion. The 358 distribution of ages is close to a normal distribution. It is therefore possible to report a single 359 crystallisation age of the zircons of  $1.322 \pm 0.024$  Ma. While this age is within uncertainty of the preferred eruption ages defined by both <sup>40</sup>Ar/<sup>39</sup>Ar and ID-TIMS (Figure 3), the higher precision ID-360 TIMS measurements of Wotzlaw et al., (2015) and Rivera et al., (2016) provide better estimates of 361 eruption age. The value of the LA-ICPMS ages provided here is that the oldest age returned is 1.634 362

- 363 Ma. Zircon ages older than the inferred eruption age have been reported from the SIMS study of
- Bindeman et al., (2008) which returned zircon core ages of  $1.45 \pm 0.03$  Ma and  $1.49 \pm 0.05$  Ma and a
- 365 single rim age of  $1.49 \pm 0.05$  Ma. The ID-TIMS studies of Wotzlaw et al., (2015) and Rivera et al.,
- 366 (2016) report zircon ages as old as  $1.327 \pm 0.017$  Ma and  $1.568 \pm 0.011$  Ma respectively.

367

368 ----Figure 8 Zircon U-Pb diagram----

369

# 370 **5. Discussion**

# 371 **5.1 The origin of 'too old' sanidine**

372 Clearly the MFT, in addition to all other Yellowstone rhyolites so far dated via <sup>40</sup>Ar/<sup>39</sup>Ar (e.g.

373 Gansecki et al., 1998; Ellis et al., 2012b; Rivera et al., 2014, 2016; Singer et al., 2014; Stelten et al.,

2015; Matthews et al., 2015; Troch et al., 2017), contains sanidines that are slightly, but distinctly,

375 'too old' to be considered juvenile to the magma which erupted. Below we assess the potential reasons

376 for the presence of this older sanidine.

377

### 378 **5.1.1 Entrainment of pre-existing material**

379 One potential source of older sanidines within the Mesa Falls Tuff would be the country rocks through 380 which the Mesa Falls magma transited en-route to the surface. For such material to be incorporated in 381 the MFT magma and not be fully re-equilibrated would require the material to be taken in only a few 382 months prior to eruption. Diffusional calculations for Yellowstone rhyolites by Gansecki et al., (1996) indicate that even xenocrysts of 1000 Ma would return juvenile ages after only several years 383 384 immersion in the magma. Gansecki et al., (1996) proposed that the older sanidine and plagioclase they 385 observed was sourced from the Eocene Absaroka volcanics and older basement rocks, on the basis of slightly more anorthitic plagioclase compositions (up to  $An_{40}$ ) and rare  ${}^{40}Ar/{}^{39}Ar$  ages as old as 370 Ma 386 but more commonly around 25-50 Ma. The use of the 'split-grain' approach allows us to address this 387

by looking at compositions of melt inclusions within the dated sanidines. It is clear that all of the 388 389 sanidines in our study crystallised from A-type rhyolitic magmas as illustrated by the relatively flat 390 REE patterns with steep negative Eu anomalies (Fig. 5) and the relatively low Sr contents, rather than 391 melts generated from subduction-related magmatism. The lack of Eocene ages in our LA-ICPMS U/Pb 392 zircon dataset further suggest limited incorporation of significantly older materials, consistent with 393 other studies of Yellowstone zircon (e.g. Bindeman et al., 2008; Watts et al 2012; Stelten et al., 2015). 394 Gansecki et al., (1996) proposed that the slightly higher anorthite contents of the plagioclase may 395 reflect a subduction-related origin, but such compositions are readily observed within the earlier 396 Yellowstone-related volcanism in the Snake River Plain (Ellis et al., 2013). We speculate that the pre-397 Yellowstone ages reported by Gansecki may be a consequence of using a 'degassed' basalt to promote 398 coupling between clear feldspar and laser during these early analyses.

399 Alternatively, the 'too old' sanidine could be from entrainment of slightly older volcanism 400 from the Yellowstone episode. The pre-MFT geology is poorly known, particularly in the area of the 401 caldera which would have been devastated by the MFT eruption. It is entirely conceivable that this 402 region contained numerous rhyolitic lavas and tuffs, which are no longer preserved. The potential for 403 such preservation bias is clear; the pre-MFT record is poorly known at Yellowstone, with only 6 small 404 volume lavas (5 post Huckleberry Ridge Tuff and the earlier Snake River Butte lava) known 405 (Christiansen, 2001). When compared to the post Lava Creek Tuff record of volcanism, where at least 406 40 separate rhyolitic units are known from a roughly equivalent period of time, the difference is stark. 407 Based on the mapped distributions of rhyolites from Yellowstone, a likely potential substrate for 408 entrainment would be Huckleberry Ridge Tuff member B which underlies the whole source region of 409 the MFT (Christiansen, 2001; Fig. 6). The majority of the 'too old' sanidines have ages intermediate 410 between those of the MFT and the HRT which would allow partial re-equilibration of the Ar inventory 411 during the period between entrainment and eruption. Although in many cases (e.g. major elements and 412 trace elements) compositional information derived from the sanidines may be equivocal, the Pb 413 isotopic record demonstrates that there is no contribution from HRT member B (Fig. 6). This 414 conclusion is supported by the U-Pb measurements in zircon (even at the relatively low precision 415 afforded by LA-ICPMS), which preclude significant involvement of HRT with none of the 65 zircons

416 measured here (either in the chemically abraded or untreated populations) returning an age >2 Ma 417 which could have crystallised in the HRT (Fig. 8). However it is noteworthy that the Island Park lava 418 dome series which erupted from the same caldera system immediately following the MFT eruption 419 have Pb isotopic compositions in sanidines which are identical to MFT (Troch et al., 2017) suggesting 420 a localised geographic isotopic signature.

421

#### 422 **5.1.2 Excess Ar**

423 An alternative explanation for the subtly older sanidine ages in the MFT could be the existence of 424 excess Ar disseminated either within fluid inclusions or within defects in the crystal lattice (e.g. Esser 425 et al., 1997; Renne et al., 1997; Winick et al., 2001; Stelten et al., 2015). The lack of observable fluid 426 inclusions within sanidines during inspection under a binocular microscopes suggests the excess Ar 427 component may reside within the crystals themselves. The excess Ar explanation is attractive because 428 it allows for juvenile and 'too old' sanidines within the MFT to have the identical geochemistry and Pb isotopic compositions, with only the age being disturbed, as illustrated in Figure 9. Studies of the 429 430 gas and fluid geochemistry of the Yellowstone system have commonly detected a mantle component, principally in <sup>3</sup>He/<sup>4</sup>He (e.g. Craig et al., 1978) with recent work also documenting the role of thermal 431 432 metamorphism of cratonic lithologies in contributing to the degassing budget (Lowenstern et al., 2014). Taking a conservative lower mantle <sup>40</sup>Ar/<sup>36</sup>Ar value of 40,000 (Burnard et al., 1997; Graham, 433 434 2002) as a mixing component indicates that only small (typically <2%) mantle components are 435 required to explain the too-old sanidine (Fig. 9). While partial re-equilibration of sanidine sourced 436 from the Huckleberry Ridge Tuff could also produce a data array similar to that observed in Fig. 9, as 437 noted above the Pb isotopic compositions of the sanidines within the MFT preclude recycling from 438 HRT member B. If the subtly older sanidines in the MFT are indeed due to the presence of excess Ar, 439 this feature highlights the difficulty in using the term 'antecryst'. We note that previous studies dating 440 MFT sanidine using step-heating (Lanphere et al., 2002; Rivera et al., 2016) found older ages in the 441 lower temperature steps that in the case of the Lanphere et al., (2002) study were interpreted as excess 442 Ar. While these older ages could be interpreted as re-entrainment of older components of the same

| 443 | magmatic system the total overlap in all geochemical and isotopic parameters measured here would            |
|-----|---|
| 444 | require the recycled material to be compositionally identical to the juvenile component. The ability to     |
| 445 | directly couple the age of the sanidine to the compositional information contained in the crystal and       |
| 446 | discriminate between these two possibilities (i.e. antecrystic vs. excess Ar) highlights the utility of the |
| 447 | new split grain method.   |

- 448
- 449 ----INSERT FIGURE 9 Inv. Isochron----
- 450

# 451 **5.2 Generation of the Mesa Falls Tuff**

452 The combination of geochemical, isotopic, and geochronological records from sanidine and zircon 453 allows the petrogenesis of the MFT to be investigated in greater detail. The steep negative Eu anomaly 454 and low Sr contents in MFT melt inclusions indicate that both the Eu and Sr have been sequestered 455 into plagioclase during earlier stages of evolution and subsequently removed by fractionation. In terms 456 of radiogenic isotopes, bulk Sr, Nd, and Pb isotopic values (Doe et al., 1982; Hildreth et al., 1991) Nd 457 isotopes in glass separates (Nash et al., 2006), Hf isotopes in zircons (Wotzlaw et al., 2015), and the 458 Pb and Sr analyses from sanidines all indicate limited contributions from surrounding crustal 459 lithologies, particularly given the strong isotopic leverage provided by the Archean Wyoming craton 460 (Doe et al., 1982). Indeed, the requirement of crustal contamination to pass from typical Yellowstone 461 basalt to MFT-like compositions is slight. The large variability observed within sanidines in <sup>87</sup>Sr/<sup>86</sup>Sr 462 space is typical of the isotopic heterogeneity observed in many large silicic magmatic systems (e.g. 463 Charlier et al., 2007) and results from the rhyolitic liquids with low Sr contents being particularly susceptible to assimilation. Published oxygen isotopes reveal a normal- $\delta^{18}$ O magma (Hildreth et al., 464 1984; Bindeman and Valley 2001) in contrast to much of the younger volcanism at Yellowstone. The 465 normal-  $\delta^{18}$ O signature along with the trace elemental and radiogenic isotopic evidence above 466 467 indicates that the MFT, like the other large-volume explosive rhyolites at Yellowstone, may most

simply be explained as a magma generated by fractionation-dominated processes with limitedassimilation.

470

# 471 **5.3 Potential of the split-grain method**

472 The MFT data raise interesting questions for crystal-specific studies. Given that the similarity in 473 appearance, major and trace elemental composition of the subtly older crystals within the MFT makes 474 their distinction without the associated high-precision geochronology impossible. The requirement to 475 define the provenance of crystals is also clear for the analytically challenging projects involving 476 radiogenic isotopes or diffusional modelling whereby necessarily few crystals are typically studied 477 (e.g. Knesel et al., 1999; Morgan et al., 2006; Till et al., 2015). We note that in the studies referred to 478 the target mineral was sanidine and we propose that the split-grain method outlined here has the 479 potential to generate exciting avenues of research coupling geochronology and geochemistry in major 480 phases.

481

# 482 **6. Conclusions**

483 The main conclusions of this study are:

1. For the first time, we demonstrate it is possible to couple high-precision <sup>40</sup>Ar/<sup>39</sup>Ar geochronology with in-situ major, trace and isotopic determinations of sanidines using the split-grain approach. This technique has significant potential to help understand the complex age spectra often observed in volcanic rocks. Indeed, the geochemical aspect of the split-grain approach could readily be extended to include other avenues of research such as diffusion modelling in the sanidine, O isotopic studies, or determination of volatile contents of melt inclusions as appropriate to the sample.

490 2. By combining geochemical, isotopic, and chronological information from sanidines, for the
 491 first time we illustrate how by interpreting the data separately (i.e. just the <sup>40</sup>Ar/<sup>39</sup>Ar data or just the
 492 trace element or isotopic data) may affect the interpretation of the same population of crystals.

493 3. The <sup>40</sup>Ar/<sup>39</sup>Ar data indicate the 'too old' sanidine grains contain excess Ar, a feature 494 consistent with the elevated <sup>3</sup>He/<sup>4</sup>He found in the Yellowstone system. The occurrence of excess Ar 495 explains the compositional and isotopic similarity between juvenile and 'too old' sanidine. Isotopic 496 compositions of the 'too old' sanidine crystals reveals that they do not represent partially re-497 equilibrated material scavenged from the underlying Huckleberry Ridge Tuff member B. Such a 498 conclusion is only possible using the split grain method described here.

499 4. Radiogenic and stable isotopic compositions of both bulk samples and individual juvenile
500 sanidines do not require high degrees of any potential assimilant, a conclusion in good agreement with
501 previous studies of the MFT.

502

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# 687 Figures

Figure 1: Location map showing the outcrop of the Mesa Falls Tuff ignimbrite and the inferred source
for the deposit (after Christiansen, 2001). Inset shows the bulk compositions of the Mesa Falls pumice
compared to other Yellowstone rhyolites.

691 Figure 2: Example of a sanidine grain used in this study (grain 64) illustrating the results of splitting

the grains (further examples are in supplementary materials). Cartoon beneath illustrates what the

693 various portions of the grains are used for.

Figure 3: <sup>40</sup>Ar/<sup>39</sup>Ar ages for the Mesa Falls Tuff with grains in dark blue reflecting those used to
estimate the eruptive age of the deposit while those in light blue are excluded.

Figure 4: Compositions of the sanidine crystals used in this study with darker blue fill representinggrains that returned an eruption age and light blue fill representing those from the older population.

gruins that retained an eraption age and right onde rin representing those from the order population.

698 **Figure 5:** Rare earth element (REE) patterns from sanidine-hosted melt inclusions illustrating similar

699 compositions between the populations classified as eruption age and subtly older. Similarity in REE

700 patterns in inclusions indicates that the older sanidines grew from liquids generated in a similar

701 manner to the juvenile sanidines. Insets show typical normalised REE patterns from the Yellowstone-

702Snake River Plain province (Ellis et al., 2013) and from the Eocene Washburn volcanic centre (Feeley

703 et al., 2002).

Figure 6: Pb isotope compositions of dated sanidines from the MFT. Notably, the Pb isotopic

composition of MFT sanidines does not overlap with the composition of sanidines in Huckleberry

Ridge Tuff member B (based on data from Watts et al., 2012 and Stelten et al., 2013). This lack of

isotopic agreement comes despite HRT member B outcropping above the inferred source of the MFT,

the Henrys Fork caldera (H) illustrated in lower inset map (after Christiansen 2001). Upper inset
shows all sanidine Pb isotopic data from Yellowstone.

710 **Figure 7:** Sr isotopic composition of sanidine from the MFT. Grey bar highlights the bulk <sup>87</sup>Sr/<sup>86</sup>Sr

values published for MFT by Doe et al., (1982) and Hildreth et al., (1991). The range in <sup>87</sup>Sr/<sup>86</sup>Sr is

much larger than the range in Pb isotopes (Fig. 6).

713 **Figure 8:** U-Pb ages of zircons from the MFT. No zircon dated from the MFT returns an age

consistent with crystallisation in the underlying HRT, consistent with the Pb isotopic results.

715 **Figure 9:** Inverse isochron diagram showing the MFT data. The older sanidines are found in a wedge

reflecting the presence of excess Ar. An inverse isochron based on the age of the Huckleberry Ridge

717 Tuff is shown in purple. The green squares on the y axis represent mixing of an atmospheric

component with various proportions (up to 3%) of a deep mantle source with <sup>40</sup>Ar/<sup>36</sup>Ar of 40,000 (after

719 Graham, 2002).

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