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dark count rates and the deleterious effects of after-pulsing, which generally result in a need for both gated operation and much reduced overall count rates [11] compared with Si-based SPADs. In this Letter we present ranging results obtained with a nanopatterned superconducting single-photon detector (SSPD), which has demonstrated single-photon sensitivity at longer wavelengths with lower jitter of 68 ps FWHM and a Gaussian profile [12]. We estimate the minimum resolvable surface separation using the same approach employed previously [10]. The data from these measurements were processed using two different algorithms (described below) in order to determine the surface-to-surface separation. The minimum resolvable separation was 4 mm.

The SSPD is a 100 nm wide niobium nitride wire patterned in a meander line covering a $10\ \mu\text{m} \times 10\ \mu\text{m}$ area [13]. The SSPD is fiber-packaged and mounted in a cryogen-free, commercial cryocooler [14] operating at a temperature of 3 K. The detector is mounted on a stage with passive temperature stabilization, and the fiber coupling is not affected by cryocooler vibrations. The detector is coupled via a single-mode telecommunications fiber to the rest of the optical ranging system. This type of detector offers several important advantages: single-photon sensitivity at infrared wavelengths ($\sim 1\%$ detection efficiency at $\lambda \sim 1550\ \text{nm}$), coupled with low timing jitter ($< 70\ \text{ps}$), short recovery time ($< 10\ \text{ns}$), and low dark counts (< 100 counts per second). With the SSPDs used in this Letter, we were able to utilize a pulsed laser emitting at 1550 nm, resulting in a ranging system that is more eye-safe. The main contribution to the background count level in the measurement is from ambient light; however, this is reduced by spectral, spatial, and temporal filtering [7]. Operation at the 1550 nm wavelength, as described in this Letter, allows these filtering techniques to be applied with much greater efficacy, as the solar background is greatly reduced at this wavelength, making the system considerably more robust against changes in the ambient light level.

The transceiver (Fig. 1) comprises a standard 200 mm diameter aperture telescope with a 1550 nm wavelength pulsed laser diode (pulse FWHM less than 50 ps). The eyepiece of the telescope was modified for imaging onto a CCD camera for initial alignment and for the collection of the reflected laser signal, which is focused into a multimode optical fiber. An in-line bandpass interference filter (FWHM $\sim 15\ \text{nm}$) centered at a wavelength of 1550 nm was used to reduce the background light at the detector. All measurements were taken during ambient daylight conditions with an average laser power of $\sim 5\ \mu\text{W}$. A pulse-pattern generator (operated at 40 MHz) was used to clock the laser driver and synchronize the signal for data acquisition.

To determine the surface-to-surface separation, two identical air-filled retroreflecting corner cubes were attached to translation stages and mounted on an optomechanical rail that was positioned at a range of approximately 330 m from the telescope. The retroreflectors were used to provide two closely

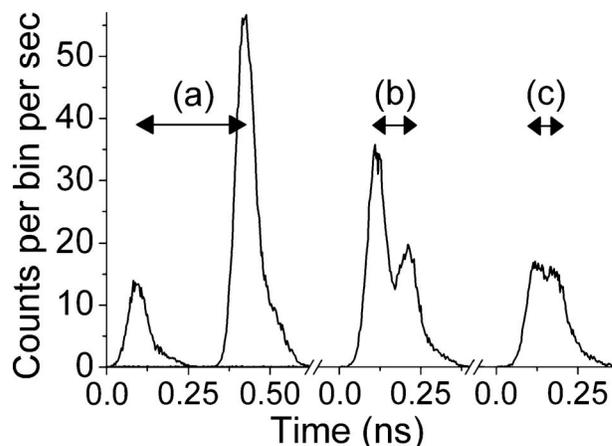


Fig. 2. Three separate examples of the time responses of the system where the target consisted of two corner-cube retroreflectors separated by (a) 50, (b) 15, and (c) 10 mm. The acquisition time was 30 s in each case. The heights of each trace are slightly different due to variations in optical alignment among the three measurements.

spaced reflecting surfaces with variable separation from 500 mm down to a minimum of 1 mm. The optics of the laser diode were configured so that both corner cubes were illuminated at the range of 330 m, and the photon returns were collected by the telescope.

In Fig. 2 we see three typical photon-return histograms as recorded by the photon-counting card. In this example the corner cubes were separated by a distance of 50 mm in Fig. 2(a), 15 mm in Fig. 2(b), and 10 mm in Fig. 2(c). An acquisition time of 30 s was used in each case. In these examples, the 4096 channels, or bins, of the photon-counting card were divided over 12.5 ns, giving a bin width of 3.073 ps. Photon returns were recorded at a rate of $\sim 1000/\text{s}$ (dependent on the optical alignment). In Fig. 2, the instrument response of each surface return is 70 ps FWHM; the slight shoulder on the right-hand side of each individual surface return is due to the asymmetric shape of the laser output pulse.

To analyze the results, two different approaches were implemented. The first method is straightforward: the position of each peak is determined by its centroid from which the surface separation can be estimated. This is a satisfactory approach for well-separated return peaks; however, as the two surface return signals merge, the process becomes increasingly problematic.

For peak separations that overlap, we used a previously developed algorithm based on reversible jump Markov chain Monte Carlo (RJMC) techniques [15]. This technique determines an estimate of the number of scattering surfaces, their positions, and the amplitudes of the returned signals from a distributed target. The RJMC approach is an extension of the basic Markov chain Monte Carlo (MCMC) algorithm designed to allow jumps between different solutions with different numbers of surface scatterers. Therefore, it is possible to consider the number of peaks as an unknown. In these experiments the only *a priori* information that we include is

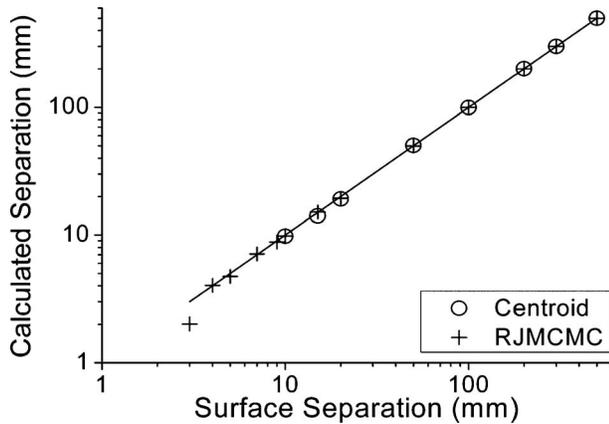


Fig. 3. Calculated separation versus surface separation. An acquisition time of 30 s was required for each data point; photon returns were recorded at a rate of $\sim 1000/s$. The black line represents the case of perfect agreement.

the instrumental response of the detector. We initialized RJMCMC with ten uniformly distributed peaks to represent our lack of knowledge about the actual distribution of the target return. We are able to extract the positions and amplitudes of the returned signals from appropriate prior distributions, as explained in [16].

Figure 3 shows the separation calculated using both centroid and RJMCMC methods. At large separations both processes perform well. For surface separations of less than 10 mm the target returns were sufficiently close to render the centroid method wholly unreliable. However, it is clear that using the RJMCMC algorithms enables a much lower separation to be resolved, 4 mm being the minimum reliably demonstrated with these measurements.

We have successfully demonstrated subcentimeter ranging resolution at a standoff distance of 330 m at a wavelength of 1550 nm using time-correlated single-photon counting with a superconducting niobium nitride nanowire single-photon detector. This low jitter detector, used in conjunction with advanced signal processing algorithms, has permitted depth resolution in the millimeter range for the first time to our knowledge using this ranging technique. Furthermore, operation in the 1550 nm spectral region enables eye-safe ranging to be carried out in daylight conditions.

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